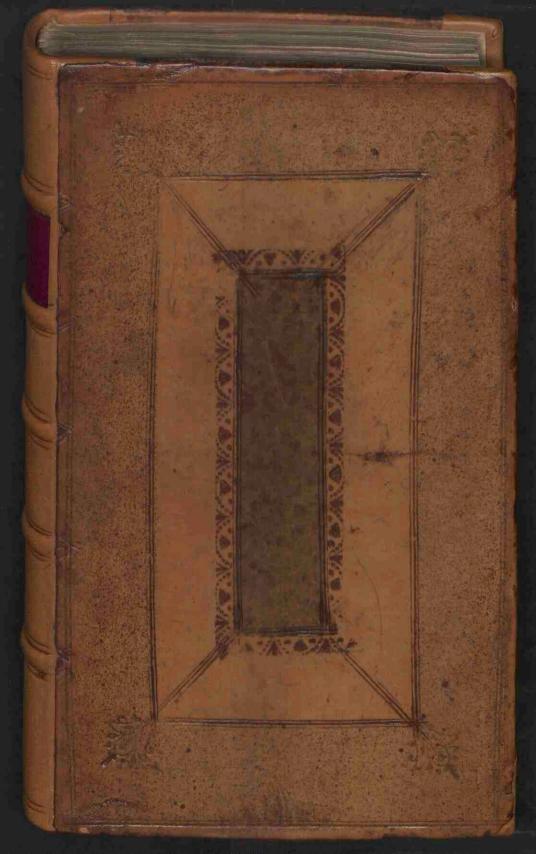
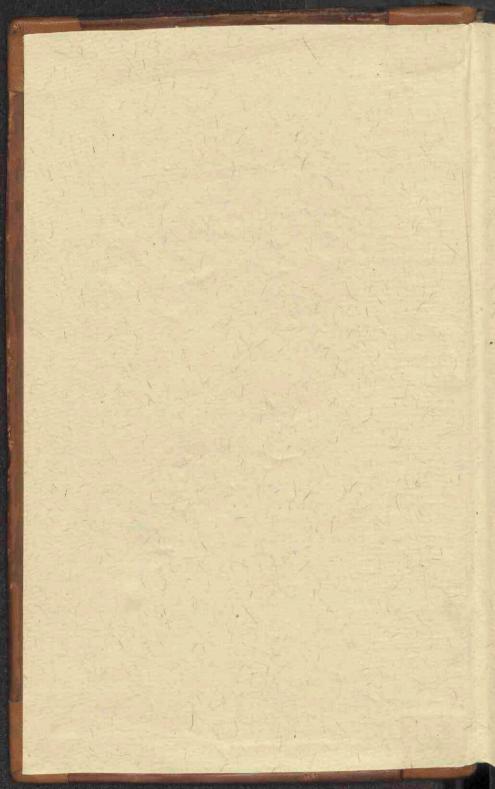
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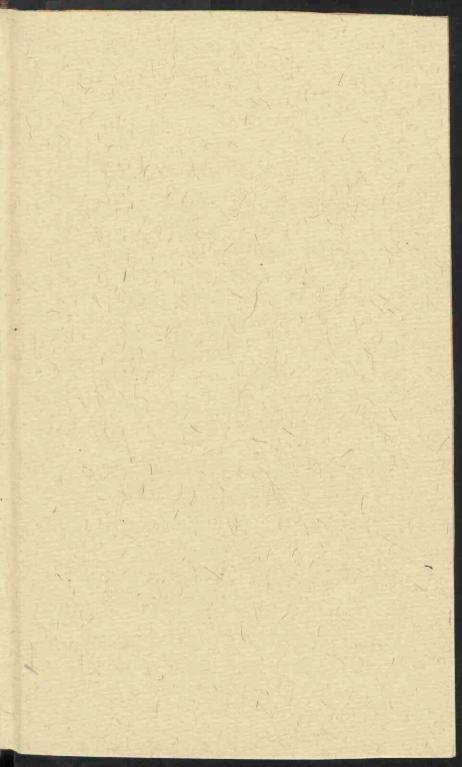
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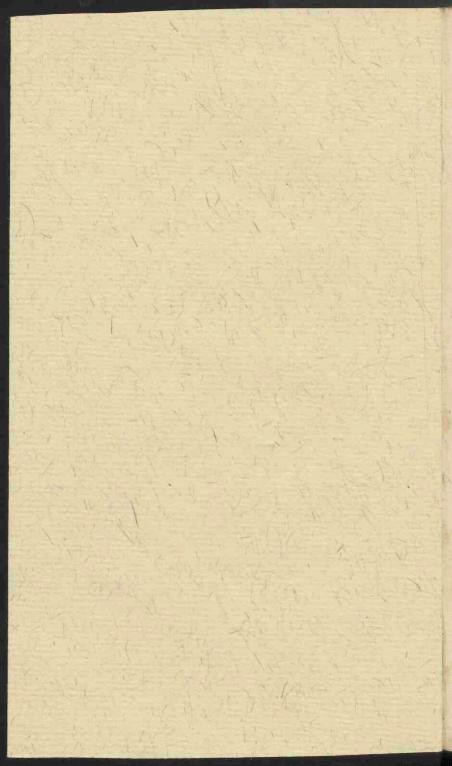
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Mathematical ELEMENTS Natural Philosophy, Confirm'dby EXPERIMENTS; OR, AN INTRODUCTION TO Sir ISAAC NEWTON's Philosophy. VOL. II. Bei Written in LATIN By WILLIAM-JAMES 'S GRAVESANDE. Doctor of Laws and Philosophy, Professor of Mathematicks and Aftronomy at Leyden, and Fellow of the Royal Society of London. Translated into ENGLISH By J.T. Defaguliers, LL.D. Fellow of the Royal Society, and Chaplain to his Grace the Duke of CHANDOS. The Fourth EDITION. LONDONS Printed for J. SENEX in Fleet-Street, W. INNYS in St. Paul's Church-yard ; and J. OSBORN and T. LONGMAN in Pater-Noster-Row. MDCC XXXI.

Utrechts Universiteits Museum

Ro. Co



Moft Noble and Right Honourable THOMAS Lord PARKER,

Baron of Macclesfield,

Lord HIGH CHANCELLOR of Great Britain, &c.

My Lord,



F whilft Your Lordfhip's Hours are taken up with an Employment of the great-

eft Fatigue as well as the higheft Honour, there can be any Time allowed for Recreation, your a 2 Lord-

iv DEDICATION.

Lordship makes that your Diversion, which few can attain to without painful Application and laborious Study. To become a skilful Mathematician and a found Philosopher, and at the fame Time shine in other Parts of Learning, requires a great and extenfive Genius : But to lead a Life of Bufinefs, and be eminent in the Law, where Reputation is only got by conftant Practice, as well as Brightnefs of Parts; and yet, in those few Minutes of Leifure that are allowed to breathe in, and are, as it were, ftolen from Sleep, to play with all the Intricacies of Lines and Numbers, to view and understand the System of the World, the Proportion, Symmetry, and Harmony of its feveral Parts; to be acquainted with all the Experiments of Confequence that have been made, and be able to contrive new ones as useful as instructive; was only re. ferved * ATTA 2.3

DEDICATION.

V

ferved to Men of uncommon Capacities, and to none in fo eminent a Degree as to your Lordship.

THE Honour I have had of being admitted into your Lordship's Conversation, has given me an Opportunity to know which of the Mathematical Sciences are chiefly lik'd by you: And as *Astronomy* and *Optics* feem to have the Preference, I thought this Translation of Dr. 's GRAVESANDE'S Second Volume would not be unacceptable.

HERE you have the Principles of the common Optics reduced into a fmall Compafs, and confirmed by new Experiments of the Author's own Invention : A fine Application of the Action of Electric Bodies to difcover the Nature of Fire : And Sir IsAAC NEW-TON'S Doctrine of Light and Colours a 3 proved

vj DEDICATION.

proved by the most confiderable of his Experiments, which Dr. 's Grave fande performs with an Apparatus very ingenioufly contrived, and nicely exprefied by curious Figures. The laft Part of this Volume not only leads a Beginner gradually on from the moft fimple to underftand the most difficult Phanomena of Astronomy; but gives fuch a phyfical Account of the Celeftial Motions as muft be fully fatisfactory to the beft Geometricians. There your Lordship will fee with Pleafure, that there are Professors abroad who understand the PRINCIPIA; and have fo just a Value for that excellent Book, as to take Pains to propagate the wonderful Truths which it demonstrates, fo as to make them plain to fuch Philofophers as are not yet able to go through all the difficult Propositions from whence they are deduced. But I will detain your Lordship

DEDICATION. vij

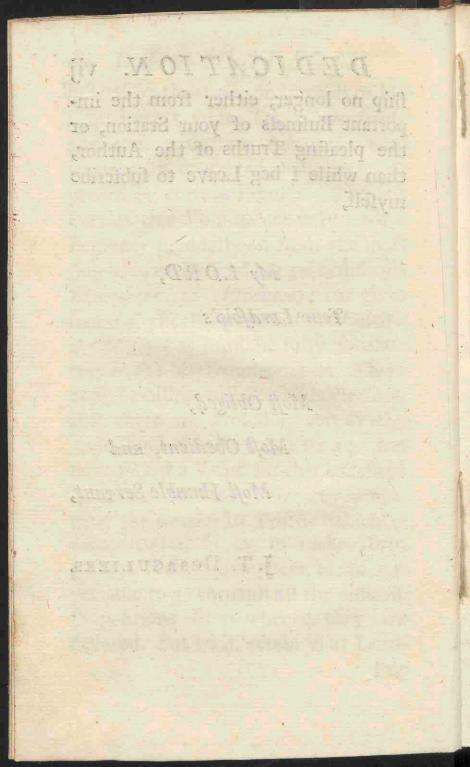
fhip no longer, either from the important Bufiness of your Station, or the pleasing Truths of the Author, than while I beg Leave to subscribe myself,

My LORD,

Your Lordship's

Most Obliged, Most Obedient, and Most Humble Servant,

J. T. DESAGULIERS.





The AUTHOR's PREFACE.



Spoke of the Method of Reasoning to be used in Natural Philosophy, in the first Chapter of the first Volume, and in the Preface endeavoured to vindicate the Goodness of the Method I have followed. There are several remarkable Specimens of this

His

Method in the present Volume, which evidently shew Sir ISAAC NEWTON'S great Superiority of Genius above all other Philosophers.

Before him, Naturalifts were in the Dark in numberless Things relating to Light, and especially to Colours. For Instance, whoever suspected before, that the Opacity of Bodies depended upon their Interstices, so that a Body becomes transparent, when these Interstices are filled with a Medium of the same Density as the Particles of the Body itself. His Account of the Planetary System, and particularly of the Motion of the Moon, is not less worthy of eternal Praises, being likely to carry Astronomy to a greater Pitch of Perfection than the nicest Observations alone could possibly do. For if a Manis acquainted with the Laws that govern the System of the World, he will be able to make a better Use of his Observations, and to compute the Motions of the beavenly Bodies more exactly than if he had nothing but Observations to direct him.

It was my Defign in these two Books, to give my Reader a general Notion of the chief Things discovered by Sir ISAAC NEWTON in Natural Philosophy, and thereby to encourage him to the Study of the more abstruse, and, at the same Time, more sublime Parts of Mathematics, after he has learned the first Principles of Geometry, to fit him for the reading of these plain Elements. He will, as it were, go to the Fountain Head, when he reads the Writings of our great Philosopher, which will reveal such Things to him, as were anknown to the profoundest Philosophers before him; and which, though published, are still a Secret to all but thorough Mathematicians.

I have only a few Words more to fay to the English Reader, concerning the two English Translations of this Work.

As tedious and diffasteful as an Author's Complaints generally prove, they cannot, however, be difallowed, when they are grounded upon such an Injury done to the Author, as it is his Reader's Interest to be informed of.

Soon after the Publication of the first Volume of these Elements, it was translated into English, and printed

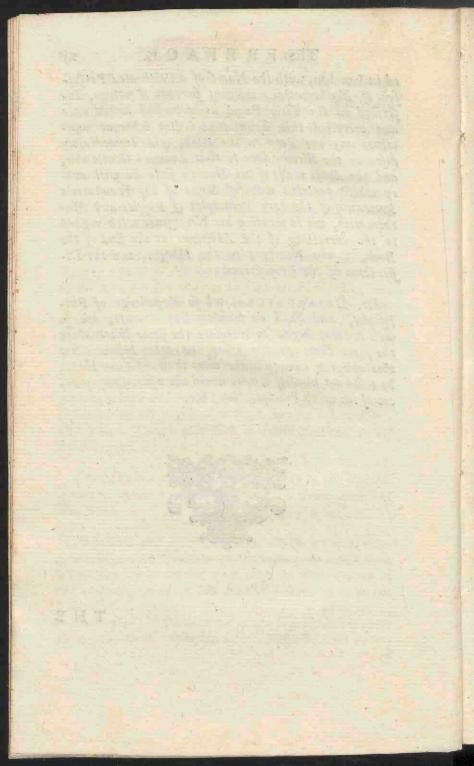
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ed in London, with the Name of a Celebrated Profefor of Mathematics, eminent for his Writings, infcribed on the Title Page, as if he had looked over and corrected this Translation. But whoever examines any one Page in the Book, will immediately discover the Wrong done to that Learned Gentleman, and the Abuse made of his Name; fince he will every where perceive manifest Signs of the Translator's Ignorance of the very Principles of Physics and Mathematics, not to mention his Negligence with regard to the correcting of the Additions at the End of the Book, to the Numbers in the Margin, and the Difinction of the Propositions.

Dr. DESAGULIERS, whose Knowledge of Philosophy, and Skill in making Experiments, are so well known, began to translate the same Work about the same Time as the other, or rather before. But this obliging him to make more than ordinary Haste, be could not himself wholly avoid the usual Consequences of too much Precipitation, &c.



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Mathe-



Mathematical ELEMENTS

OF

Natural Philosophy,

Confirm'd by

EXPERIMENTS.

Vol. II. BOOK III.

PART I. Concerning FIRE.

CHAP. I.

Of the Properties of Fire in general.



HO we know feveral Properties of Fire, yet we are ignorant of a great many things relating to it. I fhall not invent Hypotheles, but reason from Experiments.

but reafon from Experiments, and leave untouched what is

not fully known. Fire eafily penetrates thro' all Bodies, however denfe and hard they are. For we have never yet known Vol. II. B any

Mathematical Elements Book III.

any Body, that by the Application of Fire has not been heated in all its Points; that is, in every Part.

545 Fire moves very fwiftly, as appears from Aftronomical Obfervations.

2

- 546 Fire unites it felf to Bodies; for when they are brought to the Fire, they grow hot, as we faid
- 547 before; and in that Cafe they expand or fivell: Which Expansion is also observed in such Bodies,
- 548 whole Parts do not cohere, in which Cafe they allo acquire a great Degree of Elasticity, as is observed in Air and Vapours.

549 That Fire is attracted by Bodies at a certain Diflance from them, will be flown in the following * 611 Part of this Book. *

If any Bodies are violently moved againft one another, they will grow hot by fuch a Friction, and that to a great Degree ; which fhews that all Bodies contain Fire in them : for, by rubbing, Fire may be put in Motion, and feparated from Body, but can by no means be generated that way.

Having laid down these general Heads, we must examine several Things more particularly.

CHAP. II.

That Fire adherers to Bodies, and is contain'd in them; where we shall also speak of Electricity.

A S we have already faid, one may prove that Fire is contain'd in all Bodies, becaufe there are no Bodies but what may be heated by Attrition; * and that it coheres firmly with the Parts of Bodies, appears in Smoke and Vapours : for Smoke and Vapours are made up of Parts feparated from Bodies, and agitated (fome-2 times

times very violently) by the Fire that is join'd with them.

There are, befides, feveral remarkable Phænomena arifing from Fire, contained in Bodies, fome of which we fhall here mention; amongft which, there are fuch as relate very much to Electricity, for which Reafon, we muft alfotreat of the Phænomena of Electricity.

DEFINITION.

Electricity is that Property of Bodies, by which 551 (when they are heated by Attrition) they attract, and repel lighter Bodies at a sensible Distance.

Experiment 1.] Take two Pieces of Rock 552 Cryflal and rub them together, and immediately they will appear luminous all over, tho' they do not acquire any fentible Heat by that Attrition. Light (as well as Heat) is a Proof that there is Fire in a Body. The greateft Light is in those Points, where the Bodies touch one another.

Experiment 2.] Take a Glass Tube 15 or 18 553 Inches long, and one Inch in Diameter, and rub it with a Linnen or Woollen Cloth, and it will emit Light in the Dark.

Experiment 3. Plate I. Fig. 1.] This Tube, 554 heated by rubbing, has a very fentible Electricity; for if light Bodies, fuch as Pieces of leaf Gold, and Soot, be laid upon a Plane, and the Tube be brought near them, they will be put in Motion, attracted, repelled, and driven feveral Ways by the Tube. The Tube acts at different Diflances, according to the different State of the Air; fometimes at the Diffance of one Foot; when the Air is full of Vapours the Effect is diminifhed.

There

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There is one Thing remarkable, and very hard to explain in this Experiment, concerning the Direction of the Attrition; when you rub the Tube, one End of it hisheld in one Hand, whilft it is rubb'd with the other; which, if it bedone from the Hand that holds towards the other End of the Tube, the Effect will not be fenfible; but if you rub from the free End of the Tube towards the End held in the Hand, the contrary will happen. And this happens indifferently, whether you hold the open or the flut End of the Tube in your Hand.

In the following Experiments (Plate I. Fig. 2.) 556 Glass Globes are fwiftly whirl'd about ; to perform which, there must be a cylindric Neck at each End of every Globe; but only one of thefe Necks is to be open, and both are to have Brafs Ferrels, fuch as are reprefented at G; a Cock E must be forew'd on to the Ferrel at the open Neck, and the opposite Brass Ferrel must have a little Wheel r of about an Inch and a half Diameter, join'd to it, with a small Brafs Axis flanding out : There is just fuch another Piece of Brassfcrew'd on to the Cock, fo that it may be taken on and off at Pleafure. Thefe Axes go about a Quarter of an Inch into the Pillars S S, that Support the Globe, and are Centers for it to whirl upon when it turns about its Axis.

The Pillars SS ftand upon a ftrong horizontal Board of about an Inch and a half thick, framed into three other Boards, as may be feen in the Figure: In that which ftands foreright, there is an Hole f that you may come at the lower Part of the Pillar, to make it faft with a Nut or Screw. The other Pillar S is likewife faften'd by a Screw applied under the horizontal Board, and moved forwards and backwards in a Slit of 4 or 5 Inches long, before it be made faft, in order to take the I

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Glafs Globes in and out, and to rake in fuch as are larger or fmaller than others, according to the different Experiments to be made.

5

There is a great Wheel R which is turn'd round by means of the Handle M, and thereby gives a very fwift whirling Motion to the Globe G.

In the Side of the upper Board, there is a Slit, along which the Pully t may be moved, by means of the Screw c, in order to keep firetch'd the Rope that goes round the Wheels R and r.

Experiment 4 & r. Plate II. Fig. 1.] Apply a 557 Glafs Globe of about 8 or 9 Inches Diameter, to the Machine above-mention'd, and let it be briskly whirl'd in a dark Place, the Hand all the while being held against it, to give it Attrition.

If the Globe be exhausted of its Air, it will appear all luminous within, but mostly fo where the Hand touches the Glass.

But if the Globe has Air in it, and being 558 whirl'd in the fame Manner, the Hand be applied to it, no Light appears, either in the inner or outer Surface of the Glafs; but Bodies at a fmall Diftance from the Glass (as for Example, at a Quarter of an Inch. or nearer) become luminous; and fo only those Parts of the Hand held against the Glass, which terminate, or rather environ the Parts that immediately touch the Globe, are luminous.

Experiment 6. Plate I. Fig. 2.] Take the Globe 559 made use of in the foregoing Experiments, and put it in between the Pillars, to whirl it as before: Then take a Brafs Wire a b d, circularly bent in the upper Part, and fix it fo that its curve Part may be about 4 Inches off of the Globe, with small Threads hanging from it, which being extended towards the Center of the Globe come

come within a quarter of an Inch of the Surface of the Globe.

Whirl the Globe and apply the Hand, and immediately the Threads will be moved irregularly by the Agitation of the Air; but when the Glass is heated by the Attrition, all the Threads are directed towards the Center of the Globe, as may be feen in the Figure : And if the Hand be applied a little on one Side, or nearer the Pole of the Globe, the Threads will be directed towards that Point of the Axis which is under the Hand.

560 If the Air be drawn out of the Globe, this whole Effect ceafes.

561 Experiment 7. Plate I. Fig. 3.] Take another Globe like the former; only differing in this, that the open'd Neck must have a larger Opening than that of the Globe G, fo that you may put into it a round flat Piece of Wood o, that has a Brass Wire for its Axis. In order to fix this Piece of Wood in the Middle of the Globe, its Axis must be firmly forew'd to the Middle of the Cover that is join'd to the open Neck of the Globe at b: and the Cock E is also join'd to the Middle of the Cover on the outfide.

To the wooden Circle fmall Threads are faften'd, which, being extended, almost touch the inner Surface of the Globe.

Turn the great Wheel fo as to whirl the Globe, and rub it till it becomes warm, as was done in the former Experiments; if you ceafe to whirl the Globe, and the Hand be taken off, the Threads will immediately ftretch themfelves out like Radii from the Center towards the Surface of the Globe, yet they hardly remain one Moment at reft; for tho the Globe be exactly thut, thefe Threads will be put in Motion, as appears

6

appears by blowing against the Globe, tho you ftand two Foot off, or farther, from the Globe. If you bring your Finger towards the Globe, tho you do not touch it, the Threads next to it will be attracted by the Finger, and directed towards it; nay, and fometimes they fly from it. If you apply the whole Hand to the Globe, the Threads will be mov'd violently and irregularly.

And if all the Air be drawn out, as in the 762 foregoing Experiment, the whole Effect ceases; and the Threads (both before and after the Friction) only hang down by their Gravity.

If we attend to all the foregoing Experiments, the following Conclusions feem to be naturally deduced from them, which we do not give out as certain, but very probable ; for we must always diffinguish Certainty from Probability.

That Glass contains in it, and has, about its Sur- 563 face, a certain Atmosphere, which is excited by Fri-Etion, * and put into a vibratory Motion; for it at- * 554, tracts and repels light Bodies; * the finalleft Parts 559, OI. of the Glass are agitated by the Attrition, and by * 554. reason of their Elasticity, their Motion is vibratory, which is communicated to the Atmosphere above-mention'd; and therefore that Atmosphere exerts its Action the Yarther, the greater Agitation the Parts of the Glafs receive when a greater Attrition is given to the Glafs.

The Fire, contained in the Glass, is expelled by 564 the Action of this Atmosphere ; at leaft it is mov'd with it. For when light Bodies are put in Motion at a Diftance from the Glass the Bodies also * 559; 558. become lucid at a Diftance. *

It is plain, that this Atmosphere and Fire is more 565 eafily mov'd in a Place void of Air : For if the Globe be exhaufted of its Air, there can be perceiv'd no Light, nor any Effect of Electricity on the outfide of the Globe *, but the infide of the * 557 Globe 560. B 4

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Globe appears very luminous, and this Fire is perceiv'd to be in greater Quantity in this Experiment, than in that immediately after men-55⁸ tion'd *.

8

But the Action of the Electricity does alfo * 56z ceafe on the Infide when the Air is drawn out *, which feems to overthrow what we faid of the more cafy Motion of the Atmosphere of Glafs in Vacuo: But yet it is not probable that it fhou'd move no whither in this Cafe. On the contrary, it feems to go the fame Way as the Fire, and to move that way where there is the leaft Refiftance; and that the ceafing of the Action of Electricity is to be attributed to the Want of Air, 566 by means of which, the Threads are moved by the * 605 Atmosphere; as we fhall hereafter fhew *; and the * 607 Fire, which freely penetrates all Bodies, acts upon them violently by means of the Air or Vapour.

But laying afide Conjectures, tho they have a great many Experiments for their Foundation, let us return to other Things relating to Fire; as feveral Experiments are to be made *in Vacuo*, we must deferibe the following *Machine* contriv'd to give Attrition to Bodies *in Vacuo*.

967 Plate II. Fig. 2.] Ler M be the Air Pump de437 forib'd before *, LL the Brafs Plate of the Pump
on which the Glass Receivers are fet; on each Side of the Plate there is a wooden Pillar, AD, AD, which Pillars fland on the Board that carries the Air-Pump Plate, having their Bafe upon it, and a Part below the Bafe, and going thro' the Board with an Hole in it, to receive two crofs Pieces of Wood or Wedges, that make them fast during the Esperiment, but fo that the faid Pillars may be taken away afterwards.

The Glafs Receiver, in which the Experiments are to be made, is about 9 Inches high, and 6 in

in Diameter, with a Cover * that has annex'd to * 44° it a Box or Collar full of oil'd Leathers thro' which a Brafs Axis paffes; and left in the whirling round of the Axis the lower End fhou'd move, Part of the Box is made fquare, fo as exactly to fit a fquare Hole in the Horizontal Board E E that preffes on the Cover, and is made faft by bringing down the Nuts or Screws B B at each End, without bearing upon the Shoulders of the Pillars.

Towards the upper Part of the Pillars, there is a narrower horizontal Board FF prefs'd down alfo and faftened with Nuts, which has fix'd to it, on the under Side, a Piece of Brafs g, with an Hole in it for the upper End of the above-mentioned Axis to turn in, whilft the biggeft Part of this Axis, being in the middle of the Glafs, has an outfide Screw upon it, with the two Wing-Nuts d, d, moveable upon it, in order to faften feveral Bodies upon the Axis.

The lower End of the Axistums in the Hole c. The Braß Spring f f is joined to the Piece c, which forews it down to the Air-Pump Plate at the Hole where the Air is drawn out; but there must always be made a fmall Hole or Paffage in the Piece c, for the Air to be drawn out at.

Having exhausted the Receiver, the Axis above-mentioned (by reafon of the oiled Leathers in the Collar P) may be turned round without admitting any Air; but to give it a quick Motion, there is a little Brass Wheel r of about 2 Inches Diameter, with Points in its Groove, that the Rope, that turns it round, may not flip.

The great Wheel R, of about 3 Foot Diameter with its Frame, is brought up close to the Air-Pump, and made fast to its Foot by a Screw.

The Rope, that goes about the leffer Wheelr, is brought down over Pullies obliquely placed in 9

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in the upper Part of the Pillars C, C, and being over them, comes down to the great Wheel which it goes round; fo that by turning the vertical Wheel R; the Brafs Axis, above defcribed, is carried round very fwiftly, whereby a Motion is communicated to Bodies in Vacuo for feveral Experiments.

586 Experiment 8.] Take a Glass Globe of three Inches Diameter, or of two Inches and a half Diameter, with an Hole on each Side, where it may alfo have cylindric Necks. The Axis abovemention'd must go thro' those Holes, in order to give the Globe a whirling Motion; Pieces of Cork must be put on each Side of the Globe to cover its Holes, and made fast with the Nuts d, d, in the Manner that they appear in the Figure, to hold fast together the Plates or little Wheels, thro' which the Axis goes.

This Globe, thus fix'd, will be fwiftly mov'd in Vacuo with its Axis, by turning the Wheel R: To caufe Attrition, a Piece of Woollen Cloth must be tied on to each Side of the Brass Spring f, f, which by its Elafticity preffes the Globe hard on each Side.

Making the Experiment in the dark, the Globe will appear luminous; and if the Motion be continued till the Globe grows hot by the Attrition, the Light will indeed be increased, but will appear fix'd in the Places where the Attrition is made.

569 It follows from that Experiment, that the Fire, contained in Glass, does not want Air to make it vifible, for it grows hot, and fhines when both the internal and external Air are taken out.

570 Experiment 9.] Take a round Piece of Wood of 2, or 2 - Inches Diameter, and about + an Inch

Inch thick, and let it have feveral Hollows cut in its Edge, that it may be encompass'd round with Beads of Plaister, which may be made fast to it with Threads going thro' them. Let this Piece of Wood, thus prepar'd, be made fast upon the Axis above-mention'd, in the fame Manner as the little Globe had been fasten'd; then give Attrition *in Vacuo*, as in that Experiment, and Light will be thereby produced in the dark.

Experiment 10.] That Quickfilver contains 571 Fire is plain, from Experiments made upon it, in Vacuo. For if Mercury well clean'd be fhak'd about in an exhausted Glass, it will appear luminous.

If you put Mercury into a Glafs Globe, the 572 Globe may be whirl'd round, as in the former Experiments, which will be delightful to fee, if the Glafs be mov'd flowly. If Mercury has no Tin mix'd with it, it may be clean'd with boiling hot Vinegar.

The Globe, that has the Mercury in it, may 573 be exhaufted by fcrewing on to it a Pipe about 2 Foot long, the lower End of which must be fcrew'd on to the Hole in the Middle of the Plate of the Air-Pump.

Then, if you cover both that Hole and the Hole thro' which the Air is drawn out with one of the Receivers above-mention'd, the Air will be cafily drawn out of the Globe.

Plate I. Fig. 4.] Take a Plate O, and hav-574 ing fcrew'd on to it the Pipe E E, whole other End is fcrew'd to the Air-Pump, put the Glafs Receiver R upon it, and pump out the Air from it. The Pipe, thro' which the Air is drawn out, mult fland up beyond the Plate in the Receiver about 4 or 5 Inches, be bent, and have a fmall Hole in it; which mult be taken ken care of, for fear any Mercury shou'd get into the Air-Pump.

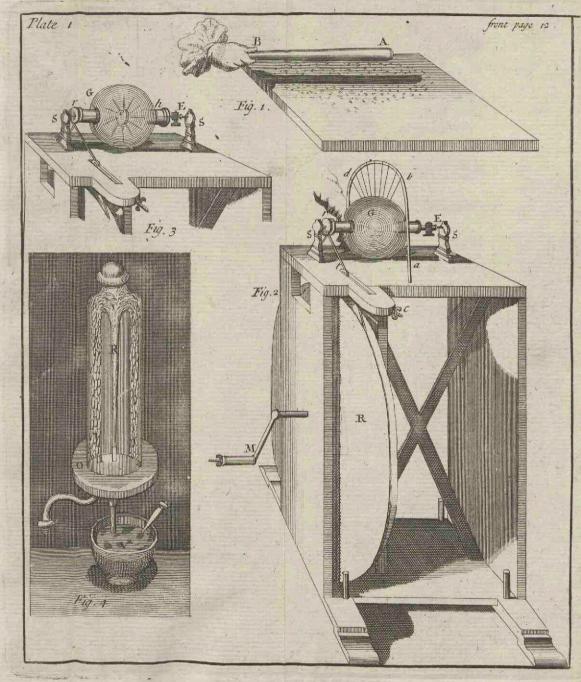
The middle of this Plate, has a Brafs PipeB; whofe lower End, that is without an Hole, comes down almost to the Bottom of the Veffel V, that contains very clean Mercury : There is an Hole in the Side of the Pipe, exactly shut by the Pin A. The upper End of this Pipe has a very small Hole in it, and comes up thro' the Plate into the Receiver.

The Height of the Receiver R is about 16 Inches, and its Diameter 4 Inches; pump the Air out of it, and then open the Hole in the Side of the Pipe B, and the Mercury by the Preffure of the external Air will go up to the Tube, and will fpout very firongly into the Receiver, firiking against the Top of it.

The Experiment must be made in a dark Place, and the Mercury will appear luminous.

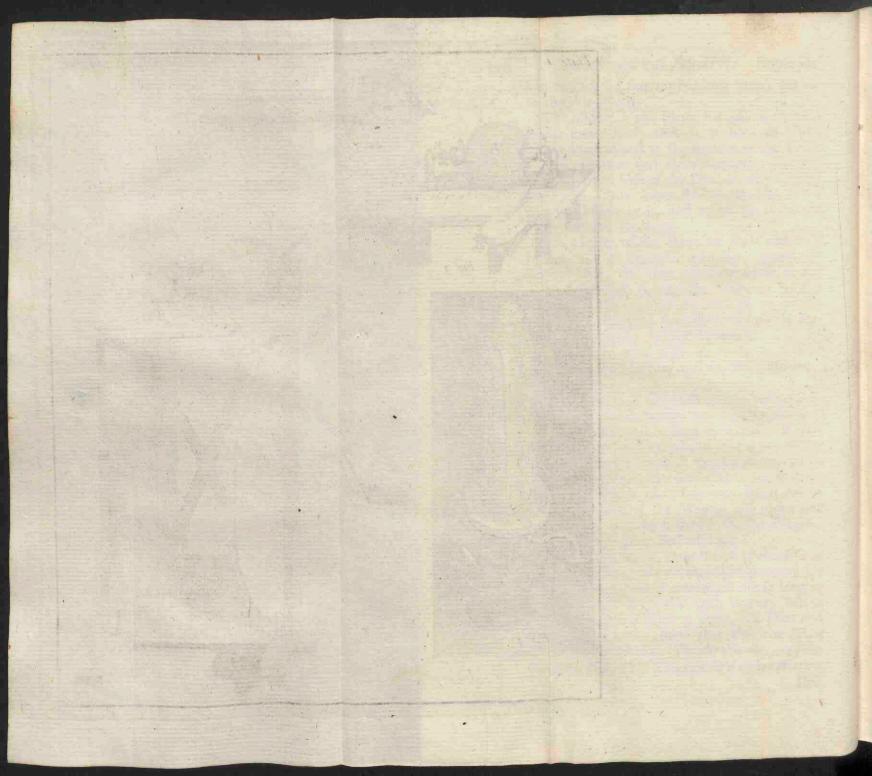
575 Experiment 11.] Chymifts make a certain Preparation of Urine, which they call *Phofphorus of* Urine; which must be kept in Water. If it be roll'd up into finall cylindrical Pieces like a Pencil, and you write with it upon a Paper, when you hold the Paper in a dark Place, you will fee Letters of Fire. The *Phofphorus* itfelf, when you take it out of the Water, will grow hot and finoke; all which shews that the *Phofphorus* contains a great deal of Fire.

576 In this Experiment, we may observe that Water acts upon the Fire contain'd in the Phosphorus; for it keeps it in, fo that it cannot get out fo long as the Phosphorus is cover'd with Water; but as foon as the Water is taken away, the Heat and Smoke immediately shew that the Fire issues out of the Phosphorus. The Air alfo does, in some measure, keep in the Fire contain'd in hot Water; that



1.

Part Part



that is, hinders it from going out of it fo fast as it will do in Vacuo.

Experiment 12.] Take two Veffels that are equal and alike, and put an equal Quantity of boiling Water into each of them. Let one of them be fet upon the Air-Pump Plate under a Receiver, and pump out the Air; when you are pumping, the Water boils violently by the Action of the Fire that goes out of it; and then foon becomes only lukewarm, whilft the Water in the other Veffel, which remain'd in the open Air, has fcarce loft any of its Heat.

One may observe fomething like this in fhi-578 ning Wood; for fome Wood that is rotted in the Ground, fhines when it is taken out; the Earth that encompass'd the Wood kept in the Fire, which goes out when that is remov'd, and the Wood continues to fhine for fome Days. In an exhausted Receiver, the Light is foon destroy'd, and is not reftor'd by the Re-admission of the Air.

But it cannot be eafily determin'd how Fire is confined in a Body by the ambient Bodies, nor is it eafy to find out what Action produces this Effect; it is hardly probable that Preffure is here much concerned, fince Fire does eafily penetrate all Bodies by its Subtilty.

CHAP. III.

Concerning the Motion of Fire, where we Shall Speak of Heat and Light.

W E have faid that Fire moves very fwiftly *, and this Motion, in different Cir- * 545 cumftances, produces different Effects. Heat and Light

Light are to be attributed to the different Mo-579 tions of Fire. Heat, in a hot Body, is the Agitation of the Parts of the Body and the Fire contained in it, by which Agitation, a Motion is produced in our Bodies, which excites the Idea of Heat

- 580 in our Mind. Heat, in refpect of us, is nothing but that Idea, and in the hot Body is nothing but Motion. Here we must call to mind what has been faid of the Senfations in general (N. 502.) which also may be referr'd to Light.
- 581 When Fire enters our Eyes in Right Lines, by the Motion that it communicates to the Fibres in the Bottom of the Eye, it excites the Idea of Light; of which Motion of the Fibres we shall speak more
- * 716 particularly hereafter. * A rectilinear Motion is
- 582 the Motion of Light, as it appears from its being eafily flopped by an Obffacle. On the contrary, fuch a Motion is not requir'd in Heat: and that an irregular Motion is more for it, may be proved, becaufe the Rays, that come directly from the Sun to the Top of a Mountain, produce no fenfible Heat; whilft in the Valley, where the Rays are agitated with an irregular Motion by feveral Reflexions, there is often produced a very intenfe Heat.

DEFINITION.

A Body is faid to be Lucid, when it emits Light, that is, when it gives Fire a Motion in right Lines.

583

3 That there is not always Light, where there is Fire, is beyond all doubt; for we daily fee hot Bodies that do not fhine.

584 But whether there be any Body without Heat, cannot be determin'd. Heat, in all Bodies, is a Motion that may be infinitely diminifh'd; and there may be fuch a Motion, tho it be not fenfible to us, becaufe often we cannot difcover any

any thing of Heat; for certainly there are a great many lucid Bodies, in which we can perceive no fenfible Heat.* Concerning which it * 571 may be observ'd, that no Heat is sensible to us, unless the Body, that acts upon our Organs of Sense, has a greater Degree of Heat than that of our Organs. Which shews us, that the Judgment of our Senses, concerning Heat, is wholly uncertain.

When the fmalleft Parts of which any Body is made up are agitated, either by Attrition, or the Action of Fire applied externally to it, or any other Way; the Fire is feparated from the fmall Particles and agitated in the Body; then alfo the Fire and the Particles of the Body act upon one another by their Attraction, as is prov'd by the Experiments which we fhall hereafter mention.* By this Action fome Parts are * 611 feparated from the Body, and carried off from it by the Motion of the Fire. And this is the Caufe why bard Bodies are often fet on Fire by a 585 violent Attrition.

Hence we deduce, that the burning of Bodies 586 is a Separation of their Parts by the mutual Action of the Fire, and those Parts on each other; some of those Parts, carried off by the Motion of the Fire, 587 make Flame and Smoke.

We fee befides, that a Body, that is burn'd by the Application of Fire, is not only diffolv'd by the Action of the external Fire, but also by the Action of the Fire contained in the Body itfelf; and that the Heat is encreased by the Application of new Fire, and the Augmentation of the Agitation of the Fire which the Body itfelf contains; and therefore that the Heat is not in 588 proportion to the Quantity of Fire.

As to the Motion of Light, it is plain, as we faid *, that it is perform'd in right Lines; but * 58, whether

whether it be fucceffive or inflantaneous, is difputed; that is, whether at the fame Moment that a Body begins to fhine, the Light is fenfible at any Diflance, or whether the Light goes on fucceffively to Places more and more diffant.

589 It feems clearly to follow from feveral Aftronomical Obfervations, that that Motion is fucceffive, and Philofophers did not long doubt of it; but by fome later Obfervations, the Conclufions, drawn from the former, are weaken'd, and we are obliged to confels that the Motion of Light has fomething unknown to us.

To fay that a Motion from one Place to another is not fucceffive, implies a Contradiction, and it can fcarce be doubted, that Light moves from one Place to another.

For we observe that Fire is carried off in Vapours and Smoke; in which Cafe, Fire carries with it the Bodies to which it adheres, and yet is often mov'd very fwift: If the Subtilty of Fire be confider'd, it will eafily be found that it is immenfly retarded by the Bodies to which it adheres; and, that, as foon as it is freed from them, it is mov'd with a very great Velocity.

There are feveral other Things very well worth obferving, in refpect to Light and Heat, but a great many of them are hard to explain. In natural Philosophy, when we are ignorant of the Causes, we must only mention the Effects.

591

We fee feveral heated Bodies that become lucid, if their Heat be increafed; fuch are Metals: They emit Fire by the Agitation of their Parts, but if the Motion of the Parts be increafed, Part of the Fire is mov'd in right Lines, and the Body fhines.

Thus

Thus if Smoke be made hotter by applying 592 Flame to it, it is chang'd into Flame; that is, it becomes a lucid Body.

Experiment 1. Plate II. Fig. 3.] Having blown out the Candle C, which fends out the Smoke F; let another lighted Candle A be brought to it, and the Smoke of the first will be chang'd into Flame, and that fucceffively quite to the Candle C, which is thus lighted by the Candle A, tho it be 7 or 8 Inches from it.

We have faid that Air acts upon Fire *; one * 577 of its Effects upon it, that in many Cales is not to be overlook'd, fhews itfelf in Refpect of Light. For the Presence of the Air is often necessary for the 593 Production of Light or Prefervation of Fire, which may be observ'd in the burning of all those Bodies that go out when the Air is taken away from them.

Experiment 2.] Put a lighted Candle under the Receiver of the Air-Pump; exhauft the Receiver, and the Candle will immediately go out.

Plate II. Fig. 2. Experiment 3.] Take a round Plate of Steel of about 3 Inches Diameter, with an Hole in the Middle to receive the Axis which is to be mov'd in Vacuo *, and let it be prefs'd * 567 and fix'd between two round Pieces of Wood by the Help of the Screws $d_1 d_2$; let Pieces of Flint be fasten'd to the Spring f, f, fo that by the Motion of the Axis the Steel Plate may rub nimbly against the Flints. As long as there is any Air in the Receiver the Attrition will produce Sparks of Fire; when the Air is drawn out, the you continue the Attrition, you can perceive no Light; but upon re-admitting the Air, it becomes again fenfible. We

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594 We observe also on the contrary, that very often the Absence of Air is necessary for Light, as might be observed in the Experiments already
* 557 mention'd *. And lastly, That taking away the 571 Air, the same Light, which may be seen in the open 595 Air, is sometimes increased.

18

* 575 Experiment 4.] With the *Pholphorus* * abovemention'd write or draw Figures upon Paper, and they will fhine in the dark, as we faid before; but the Paper will become much brighter in Vacuo.

CHAP. IV.

Of the Dilatation occasion'd by Heat.

 ⁵⁹⁶ A LL Bodies are dilated by the Action of ⁵⁴⁷ Fire*; but that Dilatation changes as the Heat changes; fo that it feems to depend rather upon the Motion than the Quantity of the Fire; for Bodies are expanded as well by rubbing as by applying Fire to them externally.

> Experiment 1. Plate III. Fig. 1.] Take a fquare from Bar about 3 Foot long and $\frac{1}{4}$ of an Inch thick, and having laid it upon a Board between two firm Obffacles or Standards O, O, between one of the Obffacles and one End of the Bar, thruft in an Iron Ruler about 4 or \int Inches long, (made Wedge-Fashion, being $\frac{1}{4}$ of an Inch broader at R than at r, and divided into equal Parts) as far as it will go; and observe the Division that is against the End of the Bar, which must be a little oblique, that it may the better apply it felf to the Ruler.

Then let the Bar be heated, either by Attrition or Fire; then let it again be put between

Book III. of Natural Philosophy. the Standards, and the Ruler thrust in between: As the Ruler will not go in fo far as the Divifion that was observed before, it is plain that the Bar is lengthen'd by Heat.

Experiment 2. Plate II. Fig. 4.] That Liquids, 597 as well as Solids, are dilated by Heat may be prov'd in the following Manner. Take a Glafs Ball G that has a fmall Tube E D for its Neck, and fill it with fome Liquor up to what Height you pleafe in the Tube, and the Liquor will afcend in the Tube, when you heat the Ball; a fmall Degree of Heat will produce this Effect, even tho you use Mercury, the denseft of all Fluids. The Experiment will be made in the fame Manner, but better, if instead of a Ball you use an hollow Cylinder G with the long Tube B A join'd to it; for the whole Liquor will be fooner heated in 'a Cylinder than in a Ball.

If the Tube ED or BA be divided into equal 598 Parts ; or the Tube with its Ball or Cylinder be fitted to a Frame on which equal Divisions are mark'd, the Heat may in fome Sort be meafured by this Machine : For the Liquor rifes or fublides in the Tube, as the Heat of the neighbouring Bodies is increas'd or diminish'd. Such Inftruments are call'd Thermometers, and are of common Ufe. They do indeed fhew a Change 599 in the Heat, but it is uncertain whether they shew the Degree of Heat ; that is, it is not known what Relation there is between the Change of Expansion and the Change of Hear, fo as to enable us to compare together the Degrees of Heat, by comparing the Degrees of Dilatation.

If the Ball G or Cylinder C be fuddenly heated, 600 the Liquor in the Tube will immediately defcend, but immediately after rife. By reafon of the fud-

den

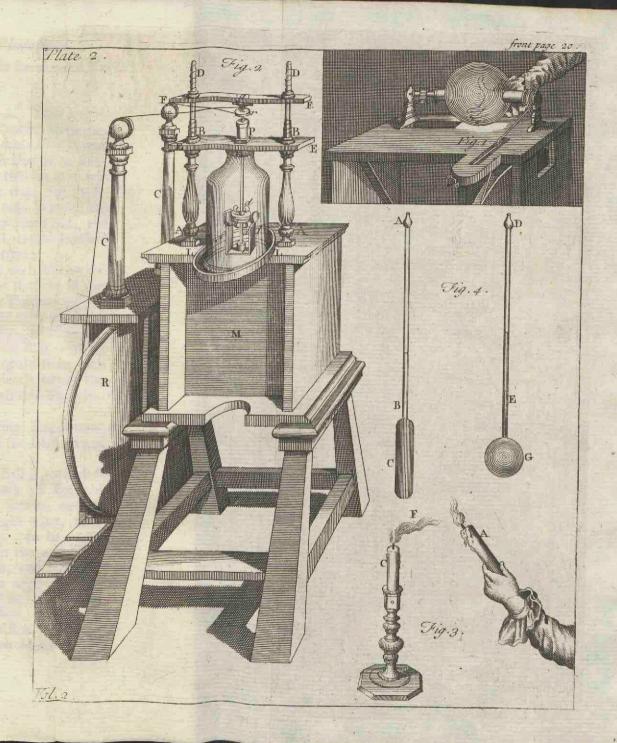
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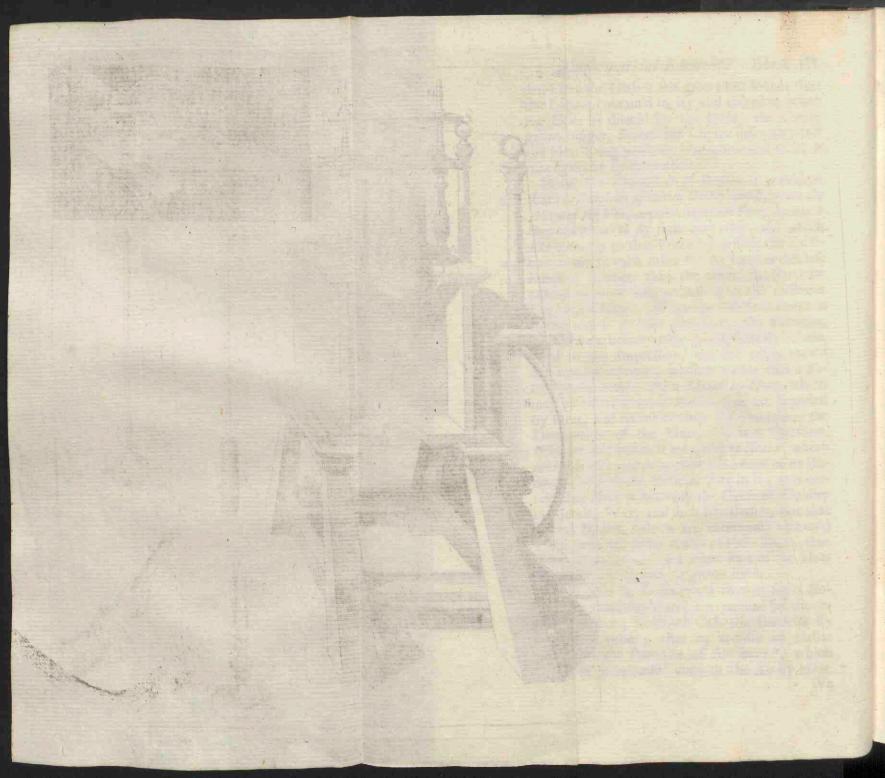
den Heat the Glass it felf grows hot sooner than the Liquor contain'd in it; and therefore when the Glass is dilated by the Heat, the Cavity grows bigger, so that the Liquor descends; but the Heat being presently communicated to it, it rifes again by its Expansion.

20

From the Expansion of Bodies it is evident 601 that the Particles of which Bodies confift, from the Action of the Fire, acquire a repellent Force, by which they endeavour to fly from each other, and which acts contrary to that Force by which the Parti-* 31 cles come to each other *. As long as this laft Force is ftronger than the other, the Particles cohere more or lefs, according to the different Degree of Heat. When the repellent Force is almost equal to the attractive, the Particles, which were before firmly join'd, fcarcely cohere, yield to any Imprefilon, and are eafily mov'd one among another; whence we fee that a So-602 lid Body is chang'd into a Liquid by Heat, which may be observ'd in all Bodies that are liquified

- by Heat, and return to their first State upon the Diminution of the Heat. It is a Question,
- 603 Whether all Fluidity is not owing to Heat? which cannot be determin'd, becaufe we know of no Body that is entirely without Fire in it; it is certain that Heat is not only the Caufe of Fluidity in Metals, Wax, and fuch like Bodies, but that feveral Bodies, which are commonly reckon'd Fluid, are only fo by reafon of their Heat; thus
 604 Water is melted Ice; for when Part of the Heat of the Water is gone, it grows fix'd.
- 605 Heat may be fo increas'd that in fome Bodies the attracting Force is overcome by the repelling Force; in which Cafe the Particles fly from each other; that is, acquire an elaftic
 * 431 Force, as the Particles of Air have *, which Elafficity is increafed even in the Air by Heat. *





Book III. of Natural Philosophy. We may observe this Effect in Smoke and Vapours.

Experiment 3. Plate III. Fig. 2.] 'Take an hollow Brafs Ball E, of about 4 Inches Diameter, with an Handle M; This Ball has a Pipe T with an Hole of hardly the twentieth Part of an Inch. Let this Ball be heated, and the Air in it will expand itfelf * and come out thro' the Pipe; * 597 then, immerging the Ball in cold Water, the Air will be again condens'd by the Cold, and the Water will go into the Ball, by the Prefiure of the Atmosphere upon its Surface.

If this Ball, thus partly fill'd with Water, be laid upon the Fire, the Moment that the Water is chang'd into Vapours, those Vapours will go out thro' T; but if the Heat be increas'd, fo as to make the Water boil violently, the Vapours, compress'd in the upper Part of the Ball, will, by their Elasticity, endeavour to recede from each other every Way, and rush violently out of the Pipe. Such an Inftrument is call'd an *Æolipile*. 606

Experiment 4.] The following Experiment 607 fhews a more fenfible Effect of the Elasticity of Vapours.

Plate III. Fig. 3.] Take the Ball E, which is alfo of 4 Inches Diameter, made of Brafs or Copper, but thicker than the former, and let it be placed upon a little light Cart, fuch as is reprefented in the Figure. In its upper Part, it has a fquare Pipe T. In the middle of this Pipe, there is a Separation, and the hind Part of the Pipe communicates with the Ball. There is an Hole of about $\frac{1}{2}$ of an Inch, in the middle of the Separation of the Pipe, which Pipe is open forwards. This Hole is flut up with a long Plate, that goes thro' two Holes in the C 3 Sides Sides of the Pipe, and applies itself to the abovemention'd Separation. As the Plate is made Wedge-wife, if you ftrike upon it with a Hammer, it will exactly thut the Hole.

Take this Ball off of the Carr, and open the Hole, and, having heated the Ball, immerge it in Water, to let it be in Part fill'd, as in the laft Experiment. Then, having that the Hole, fet the Ball again upon the Fire, till the Water boils violently; and then if you make it fast to the Cart, and open the Hole, the Vapour will fly out violently one Way, and the Cart be carried with the Ball the contrary Way,

The Vapour, being very ftrongly compress'd, 608 endeavours to recede equally every Way, and therefore opposite Preffures deftroy one another : but when the Hole is open'd, the Vapour, which goes out, does not prefs; therefore the Preffure one Way being taken off, that which acts in a contrary Direction prevails, and the Cart is mov'd along.

609 A Sky-Rocket rifes up into the Air, becaufe the Gun-powder, being fet on Fire, acquires an Elasticity, and its Parts endeavour to recede every Way: As the Pipe, or Cafe of it, is open at one End, it is lefs prefs'd that Way, and confequently at the other End, the Preffure prevails, and carrics up the Rocket.



BOOK

BOOK III.

[23]

PART II.

Concerning the Inflexion, Refraction, and Reflexion of Light.

CHAP. V.

Concerning the Inflexion of the Rays of Light.



A VING premis'd, in the foregoing Part, what regards Fire in general; let us again farther examine the Properties of Light and the Phænomena arifing from them.

The Things observed concerning Light are very wonderful, yet most of them are explained by a few Laws of Nature.

Light moves in right Lines *, can be inter- * 581 cepted by an Obffacle which wholly ftops fo much Light as comes upon it, but ftops no more.

DEFINITION.

Any Light confider'd according to the Direction 610 of its Motion, if it be all carried in the fame Direction, is call'd, a Ray of Light.

Such

Such is a Beam of the Sun's Light going thro' an Hole.

Fire, as has been faid, is attracted by Bo* 549 dies *; the notable Effects of which Attraction may be obferv'd in the burning of Bodies; they
611 are alfo fenfible in Light; for when Light paffes near Bodies, it is turn'd out of the firait Way; which may be diffinitly perceiv'd by the following Experiment.

Experiment 1. *Plate* III. *Fig.* 4.] Take a little Board Tabout 6 Inches long, and ashigh, with an Hollow *e e* in its Surface, in which two Plates of Steel flide, each of which has an Edge like a Knife; the two Edges may be brought together in the middle of the Board, and be join'd together exactly. At the Place where they meet, there is a fquare Hole in the Board, of almost an Inch, that a Beam of Light, let into a dark Room thro' an Hole of a Quarter of an Inch in Diameter, may pass thro' the faid fquare Hole, fo as to come to the Plates.

If the Diftance between the faid Edges be of about the 10th Part of an Inch, and the Light be made to pass between them, the Board being fet at the Diftance of 3 Foot from the Window, if the Light falls upon the Paper A, at 5 Foot beyond the Board, there will appear, on each Side of the transmitted Beam, a Light like that of the Tail of a Comet; which proves that the Light is inflected as it passes by the Edges of the Plates, as is plain from the Figure.

If you bring the Plates nearer (as for Example) within the hundredth Part of an Inch, inftead of the Light above-mentioned, you will, on each Side of the Light of the Beam upon the Paper, fee three colour'd Fringes, parallel to the Edges

Edges of the Plates; which will fill appear more diffinct, if the Hole in the Window be made lefs. I fhall hereafter explain whence thefe Colours arife. Now it is enough to deduce from this Experiment, that Light *is attracted* by Bodies that inflect its Rays; for if there was not a Motion towards the Body, the whole Beam would have continu'd in its direct Motion.

But the Action of Bodies, by which they act up- 612 on the Light to attract it, exerts itfelf at a fensible Distance; as is plain from the Experiment.

Experiment 2. Plate III. Fig. 4.] Every thing elfe being as in the former Experiment, if the Plates be brought nearer together within $\frac{1}{2}$. Part of an Inch, no Light will appear upon the Paper between the Fringes above-mention'd; fo that in this Cafe, all the Light, which paffes between the Edges, is inflected on either Side, fo as to produce the Fringes above-mention'd. Which plainly flews, that Steel acts upon Light at leaft at the Diffance of $\frac{1}{2}$.

It is also proved, that That Action is increas'd 613 as the Distance is diminish'd.

Experiment 3. Plate III. Fig. 4.] Things being fill in the fame Difpofition, leffen the Diffance between the Plates; and the Fringes will vanifh fucceffively, till the Plates being join'd together, no Light paffes between them. The Fringes, which vanifh firft, are those that are produced by the Rays which are least inflected; and those vanifh laft, which are produced by the Rays that are most inflected; that is, when the Edges come towards one another, the Shadow between the Fringes, made by each Edge, is continually increas'd, till, at last, the whole Light on each Side vanishes. Whence it plainly follows, that the

the Rays are more inflected, the nearer they pass by the Edges; that is, the Attraction is increased, as the Diffance is diminith'd.

614 But this Attraction has fomething particular, for the Attraction of one Edge is increased, as the other is brought near it. Which is plainly feen in this Experiment; for as the Edges are brought towards one another, the Inflection of the Rays continually becomes greater.

26

CHAP. VI.

Concerning the Refraction of Light and its Laws.

615 A LL that gives Paffage to Light, is call'd a Medium.

Giafs, Water, and a Vacuum itfelf are Mediums. When a Ray of Light goes out of one Medium into another, it is often turn'd out of the right Line.

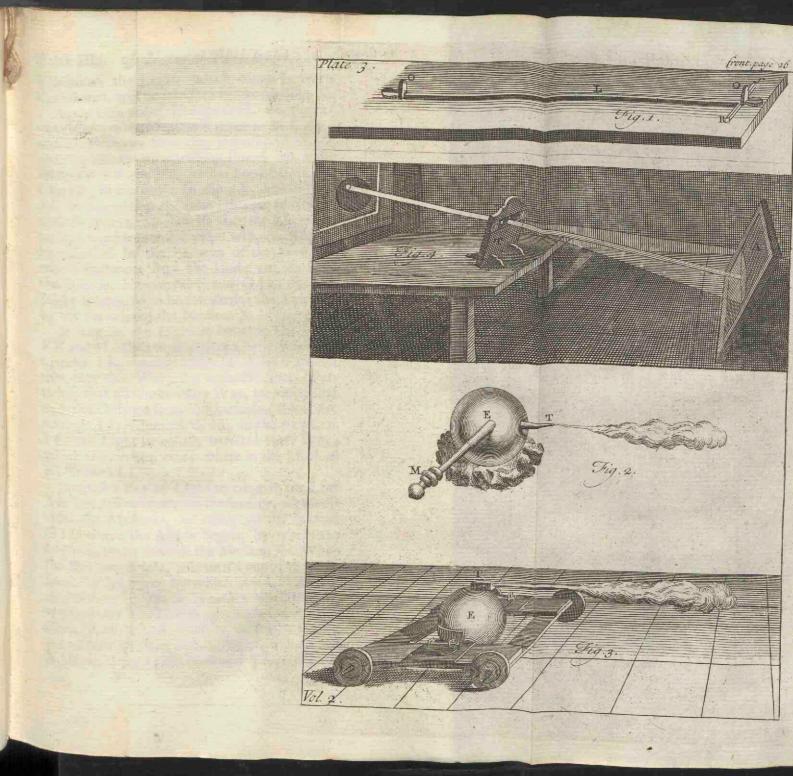
DEFINITION II.

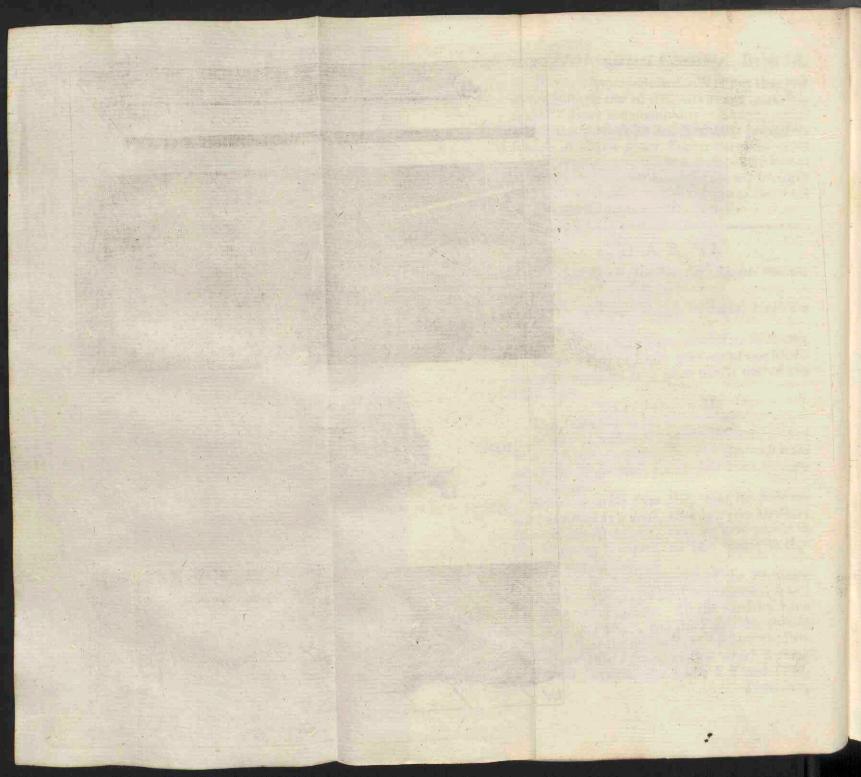
This Inflection is call'd Refraction.

616 In order to produce Refraction, the Mediums must be of different Densities, and the Ray must make an oblique Angle with the Surface that separates the Mediums.

618 Refraction arifes from this, that the Rays are more attracted by a denfe, than by a rare Medium; from which Attraction, that we have prov'd in the foregoing Chapter, all that relates to Refraction may be deduced.

619 Let E F be the Separation of the Mediums [Plate IV. Fig. 1.] and let X be in the denfer, and Z in the rarer Medium. All the Particles have
* 31 an attractive Force *, and this Force alfo obtains
* 611 in Light. Let the Diffance, at * which the Particles exert their Action, be that which is comprehended between the Lines E F and G H. Therefore,





Therefore, the Light, which comes between those Lines, will be attracted by the Medium X.

At the Diftance of the Line GH, only the extreme Parts of the Medium X act upon the Light; at a lefs Diftance, both they and other Parts act fo as to increafe the attracting Force, when the Diftance is diminifh'd, as has been before obferved*. In the denfer Medium X, let the Line * 613 I L be fuppofed at the fame Diftance from E F as G H is in the Medium Z. Let the Light enter into the Medium X, and it will, on all Sides, be attracted by the Particles of the Medium, whofe Diftances from the Light are lefs than the Diftance between E F and G H; for the Light is fuppos'd to be attracted at that Diftance by the Particles of the Medium X.

As long as the Light is between the Lines EF and IL, the attracting Force is the ftrongelt towards IL, because there are more Particles that draw that Way; but as the Number of Particles, that act the contrary Way, increases, that is, as the Distance from E F increases, the Force towards IL is diminish'd, till, in the very Line IL, the Light be equally attracted every Way; which also obtains every where in the Medium X, beyond IL.

Suppose a Ray of Light to come in the Line A a, and fall obliquely on the Surface, which divides the Mediums, or rather on the Surface G H, where the Action begins, by which the Light is driven towards the Medium X: When the Ray comes to a, it is turn'd out of the right Line, by the Force by which it is attracted by the Medium X; that is, by which it is driven towards it in a Direction perpendicular to its Surface. And, indeed, the Ray of Light is bent out of the right Line in every Point, as long as it is between the Lines G H and I L, between which

which the faid Attraction acts; and therefore between those Lines, it describes the Curve ab, * 208 in the fame Manner that Projectiles move *. Beyond the Line I L, the Action which bends the Ray ceafes; therefore it goes on then in a right Line, according to the Direction of the Curve in the Point b.

The Diftance, between the Line G H and I L, is very fmall; therefore, when we confider Refraction, we take no notice of the bended Part of the Ray, which we look upon as if it was made up of two ftrait Lines A C, C B, meeting at C, namely, on the Surface which divides the Mediums.

Thro' C draw N C M, perpendicular to the Surface of E.F.

DEFINITION III.

620 The Part A C of the faid Ray is call'd the incident Ray.

DEFINITION IV.

The Angle A C N is call'd the Angle of Inci-621 dence.

DEFINITION V.

622 The Part C B is call'd the Refracted Ray.

DEFINITION VI.

The Angle BCM is call'd the Angle of Re-623 fraction.

In this Cafe, When Light goes out of a rare into 624 a dense Medium, the Angle of Refraction is less than the Angle of Incidence, by reafon of the Inflexion of the Ray; for these Angles wou'd be equal, if the Ray AC continu'd to move ftrait on in the Line CD. The Ray CB does also come nearer to the Per-

Perpendicular CM; and therefore the Refraction is faid to be made towards the Perpendicular.

On the contrary, if the Ray goes out of a denfer 625 into a rarer Medium, it will recede from the Perpendicular, because the denser Medium attracts the Ray in the same Manner, whether it goes out of a rarer into a denser, or out of a denser into a rarer Medium. Therefore if BC be the Ray of Incidence, CA will be the refracted Ray; that is, Let the Ray come from either Side, and it will move 626 in the same Lines. Therefore, if there be two Rays, 627 and one comes out of a denser Medium into a rarer, and the other out of a rarer into a denser, and the Angle of Refraction of the one be equal to the Angle of Incidence and Refraction will be equal to each other.

Whence it follows, that The Direction of the 628 Ray is not chang'd, if it moves thro' a Medium termined by two Surfaces parallel to each other; for as much as it is turn'd towards any Side at its Entrance, fo much exactly is it turn'd the other Way as it goes out of the faid Medium.

If a Ray falls perpendicularly on the Surface that 629 feparates two Mediums, it will not be turn'd out of the right Line by the Attraction of the denfer Medium; because in that Case it acts in the Direction of the Ray.

Plate IV. Fig. 2.] To confirm what we have 630 faid by Experiments, take a Trough or open Box P, about a Foot long, 4 Inches wide, and of the fame Depth. Its two Sides *abcd*, *abcd*, are Planes of Glass parallel to each other, and this Box muft be fill'd with Water two Inches and half high.

The Experiments mull be made in a dark Place; the Light must be let in thro' a Slit in the moveable round Plate O, which is fix'd to the Wind-

Window-fhut, and is eafily turn'd round, that the Slit may be inclin'd at Pleafure : The Slit is 4 Inches long and ⁴ Inch wide.

The Light, that is let in, ftrikes upon the Looking-Glais S, and the Slit is fo inclin'd, and Speculum difpoled, that the Beam, that comes thro' the Slit, may be horizontally reflected by the Looking-Glais, the Beam being in a vertical Polition, that it may go thro' a vertical Slit in the Board T, of the fame Dimensions as that in the Plate O. This Board is us'd to diminish the Breadth of the Beam, which is continually increasing, by reason of the Light that comes from the Sides of the Sun.

Experiment 1.] Things being difpos'd, as we have faid in the Defcription of the Machine; let the Beam fall perpendicularly on the Surface $a \ b \ c \ d$, and it will pass perpendicularly thro' the Water and the upper Part of the Trough, and not be turn'd out of the Way at all, either going in or coming out; by which is confirm'd what has been faid (N°. 629.)

Experiment 2.] Now let the Beam fall obliquely on the Surface a b c d, at f g; the upper Part of the Beam will continue its Motion ftrait along b_3 but the lower Part of it, in the Water, will be bent towards *i*, coming towards the Perpendicular; which confirms (N°. 624.)

Experiment 3.] Every Thing being, as in the former Experiments, the Beam, which, at *i*, goes out of Water into Air, is turn'd out of its Way fromwards the Perpendicular, and in fuch Manner, as to move in the fame Direction as the incident Ray that goes into the Water at g; for it goes

goes on parallel to the Ray f b continued : which confirms (N^{\circ}. 625, 626, and 628.

In what has been faid hitherto, we have only confider'd the Attraction of the denfer Medium, becaufe it overpowers; but we mult not overlook the Action of the rarer Medium, becaufe it diminishes the Action of the denfer, which Action becomes fo much the lefs on the Rays of Light, as the Mediums differ lefs in Density. Therefore there is no Refraction where the Densities of Me-631 diums are equal; and the Refraction is greatest where the Densities of Mediums differ most.

The Laws of Refraction are deduc'd from the Acceleration which the Attraction generates; therefore this Acceleration muft be examin'd.

DEFINITION VII.

The Space, terminated by the Planes G H and 632. I L, is call d the Space of Attraction.

Plate 4. Fig. 1.] The Attraction acts between the Planes represented by these Lines, but no farther *.

The Direction of this Action is perpendicular to 633 the Surface that feparates the Mediums, and confequently to the Surface I L; and its Form is unequal at different Diffances from the Surface*. But * 619. it is equal at equal Diffances, because both Mediums are supposed homogeneous, and alike in all their Parts.

The Motion of the Ray AC may be refolved into two other Motions, according to the Directions AO and OC*, of which the first is * 192. parallel to the Surface E F, and the fecond perpendicular to that Surface : The Celerities of which Motions will be refpectively proportional to those Lines AO and OC, whilst AC denotes the Celerity of the Ray itself *. * 192.

The

634 The Motion, according to the Direction AO, is not chang'd by the Attraction perpendicular to the Surface IL, only the Motion according to OC is accelerated.

Keeping the Line AC, that is the Celerity of 635 the Ray, its Inclination may be chang'd, whereby alfo the Celerity in the Direction OC is alfo chang'd; which Celerity wholly vanishes, if the Angle A a G be very fmall. In which Cafe, if after Light has enter'd into a denfer Medium, its Motion be refolv'd into two, fo that the Direction of the one be perpendicular to the Surface I L, its Celerity must be wholly attributed to the Attraction fo often mention'd. For, asit enters into the Space of Attraction, a Motion is generated in that Direction; and, as it goes thro'. that Space, in which a new Action does every where act upon the Light in the fame Direction, it is continually accelerated. Which Acceleration obtains in every Passage of Light thro' the Space of Attraction; but it is different, according to the different Celerity with which Light comes perpendicular to the Surface which parts the Mediums.

If the Attraction was equable thro' the whole Space of Attraction, the Laws concerning the faid Attraction might be determin'd (as was faid of the Acceleration of heavy Bodies after N° 1 30.) by Help of the rectangular Triangle P Q R, (*Plate* IV. *Fig.* 3.) in which the Lines parallel to the Bafe represent the Celerities, whilf Portions of the Area of the Triangle represent the Spaces gone thro'.

But here we have always the fame Space run thro', namely the Breadth of the Space of Attraction, becaufe we only confider the Motion which is perpendicular to the Surface, which feparates the Mediums; therefore that Space is always reprefented by equal Parts of the Area of the Triangle PQR. Let P d c reprefent fuch a I Part

Part when the Celerity is equal to o. The Light enters the Space of Attraction in the above-mention'd perpendicular Direction, that is, when an incident Ray makes a very fmall Angle with the Surface that feparates the Mediums; d c in that Cafe will denote the Celerity acquir'd by the Attraction, and with which the Light goes out of the Space of Attraction.

But if the Light goes perpendicularly into the Space of Attraction with the Celerity that is exprefs'd by f g, it will go out of that Space with the Celerity b i, supposing the Areas P d c and fgib equal to one another, as appears from what has been faid. The Triangle P dc, Pfg, P bi are Similar, and therefore their Areas are to one another, as the Squares of the homologous Sides d c, fg, b i; but the Sum of the Areas Pdc, Pfg, is equal to the Area Pbi; (by reafon of the equal Areas P dc and f g i h;) therefore alfo the Sum of the Squares of the Lines d c and f g is equal to the Square of the Line bi; whence it follows, that with the three Lines above-mentioned one may form a rectangular Triangle, whofe Hypothenufe will be h i.

Therefore, In a rettangular Triangle, one Side of 636 which is the Celerity with which Light enters the Space of Attraction perpendicularly, and the other Side the Celerity acquir'd in going thro' that Space, when the Celerity with which the Light enters into it is equal to 0; the Hypothenuse of the Triangle will be the Celerity with which the Light goes perpendicularly out of the Space of Attraction on the other Side: Which univerfally obtains, which way foever the Attraction is chang'd, in the Space of Attraction, according to the different Diffance of the Planes that terminate that Space. Which to prove,

Let us fuppose the Space of Attraction to be divided into two Parts, whether equal or un-Vol. II. D equal

34

equal, by a Plane parallel to the Surfaces with which it is terminated. Let us fuppole belides, that the Attraction is not the fame in those two Parts, but yet that it does not vary in one of These Parts are to be confider'd as two them. different Spaces of Attraction. Let A (Plate IV. Fig. 4.) be the Celerity which the Light acquires in going thro' the first Part of the Space, when it enters the Space with the Celerity O. Let B be the Celerity acquir'd in going thro' the fecond Part of the Space, when the fame Light enters that Part with the Celerity O. It is to be observ'd that this Demonstration relates to the Motion, perpendicular to the Surface, which separates the Mediums.

Let the Light enter the first Part of the Space above-mention'd with the Celerity O, it will come to the fecond with the Celerity A; if therefore a right angled Triangle be form'd with the Sides A and B, the Hypothenuse E D will express the Celerity with which the Light will go out of * 636 the Space of Attraction *.

If the Light enters the Space of Attraction with the Celerity F G, let the rectangular Triangle HFG be drawn with the Sides FG and A; the Hypothenufe HG will be the Celerity, with which the Light goes out of the first Part of * 636 the Space of Attraction*, and enters into the fecond; and if you draw the rectangular Triangle HGI, whole Perpendicular is equal to the Line B, you will have the Hypothenufe I G to denote the Celerity with which the Light goes out, and continues its Motion after it has run * 636 thro' the whole Space of Attraction *.

Now we must demonstrate that the Celerity I G is also the Hypothenuse of the rectangular Triangle NML, whose Side ML is equal to the Celerity FG, with which the Light enters the Space

Space of Attraction; and whole other Side L N is equal to the Line E D, which is the Celerity that the Light acquires in going thro' the whole Breadth of the Space of Attraction, when it has enter'd it with the Celerity O; which being alfo demonstrated in that Cale, in which two different Forces of Attraction act, it is plain that the Proposition of N° 636 is thereby prov'd.

But it is plain from the Confideration of the rectangular Triangles, that the Lines I G, and N M, are equal. The Square of the Line N M is equal to the Squares of the Lines N L and L M or F G : N L is equal to the Line E D, whole Square is equal to the Squares of the Lines E C and CD, or of the Lines A and B, which are equal to FH and HI: therefore the Square of the Hypothenufe N M is equal to the three Squares of the Lines FG, FH, and H I. To which three Squares, the Square of the Lines G I is equal; as having been prov'd equal to the Squares of the Lines H I and H G, which laft is equal to the Squares of the Lines H F and FG.

If the Space of Attraction be divided into ever fo many Spaces by plain parallel Surfaces, which terminate that Space, and the different Parts have different Forces of Attraction, the fame Demonstration will ferve; and the Number of Divisions may be made any how in infinitum; which Cafe obtains in the Refraction which Light fuffers going out of any Medium into another of different Denfity*: to which Refraction * 636 therefore the Rule of N°. 636 may be applied.

Plate IV. Fig. 5.] Let Z be the rarer, and X the denfer Medium, and let them be feparated by the Plane E F; let a Ray of Light AC fall obliquely on the Surface E F; let A C denote the Celerity of the Light in the Medium Z, and D 2 let

let that Line AC be fettled; that is, let it be always the fame, whatever the Inclination of the Ray is. With the Center C and the Semidameter CA, draw a Circle; and let NCM be drawn perpendicular to EF; from A, draw AO perpendicular to NC, and AQ to EF.

Let us conceive the Motion along A C refolv'd into two others, the one along A O, and * 192 the other along A Q or OC *; the Line OC will denote the Ray's Celerity perpendicular to the Surface E F, which Celerity alone is in-* 634 creafed by the Attraction of the Medium *.

Let CP be the Celerity which Light acquires, in paffing perpendicularly thro' the Space of Attraction of the Medium X; fuppofing the Celerity of Light at its Entrance to be O, the Hypothenufe O P of the rectangular Triangle PCO will be the Celerity of the Ray A C in the Medium X, according to a Direction perpendi-* 637 cular to the Surface E F*; the Celerity of the Light in the Direction AO or O C, parallel to

Light in the Direction AO or QC, parallel to *634 the Surface EF, is not chang'd*. Therefore let CV be equal to AO or QC, and VB perpendicular to EF, equal to the Hypothenusc PO, and draw CB; the Motion along CB is compounded of the Two, and this Line by its Situation determines the Direction, and by its Length * 190 the Velocity of the Light in the Medium X*; which

638 Celerity is not chang'd by the different Inclination of the Ray AC. For the Square of the Line CB is equal to the Square of the Line B V or PO, the Square of the Line C V or AO; but the Square of the Line PO is equal to the Squares of the Line PC and CO: Therefore the Square of the Line C B is equal to the three Squares of the Lines PC, CO and AO; which two last, if join'd, will give us the Square of the Semidiameter AC or CN; that is, CB is equal to

to PN, whole Square is allo equal to the Squares of the Lines PC and CN, and which undergoes no Change from the different Inclination of the Ray AC.

The Line CB does in T cut the Circle, which is defcrib'd with the Semidiameter CA; from the Points B and T, draw BS and TR perpendicular to CM; by reafon of the Similar Triangles CBS, CTR, BC will be to TC or CA, as BS to TR; which Lines therefore, by reafon that BC and CA are fettled, will have the fame Ratio to one another, whatever be the Angle of Incidence. TR is the Sine of the Angle of Refraction TCR, and BS which is equal to CV, that is equal to A O, is the Sine of the Angle of Incidence ACO.

Therefore, in every Inclination of the Incident 639 Ray, there is a settled and constant Ratio between the Sines of the Angles of Incidence and Refraction.

Now fince B C and C A which are as the Sines above-mention'd, do alfo denote the Celerities of the Light in the Mediums X and Z, it follows that those Sines are inversity, as the Celerities in the Mediums.

If the Medium Z be Air, and X Water, the 640; aforefaid Sines are as 4 to 3; and the Celerity of the Light in Air to its Celerity in Water, as 3 to 4. But if, Z being ftill taken for Air, X be Glafs, the Sines will be as 17 to 11. One Experiment determines this for all Mediums.

The Ratio between the Sines of any Angles is 641 the inverfe Ratio of the Secants of the Complements, as it appears in this Figure, fuppofing a Circle drawn with the Semidiameter CQ or CV; for then AC (which is equal to CT) and CB will be the Secants of the Angle ACQ, and BCV Complements to the Angles of Incidence and Refraction, and inverfly, as BS (which is D 3 cqual

equal to A O) and T R, which are the Sines of Incidence and Refraction in the greater Circle.

And this Proportion of the Secants gives us a Method of eafily reducing to Experiment the *Prop.* of N°. 639.

Experiment IV. Plate IV. Fig.6.] In this Experiment the Light is to be let into the dark Chamber thro' a Slit, in the fame Manner as in the former Experiments, and by means of a Looking-Glafs reflected thro' the vertical Slit in the Board T.

Take a Trough P, nearly of the fame Size as that which was us'd in the former Experiments, but which has only one End of Glaís, namely, the little Side a b c d. And let it be half fill'd with Water.

The vertical Beam of Light being made to fall obliquely on the Glass End of the Trough, the Part of it, which is above the Water, goes directly forwards, and at b falls upon one of the long Sides of the Trough ; but that Part of the Beam, which is refracted in the Water, goes along gi, and firikes against the fame Side of the Trough at i. Whatever the Angle be which the Beam makes with the End or Side a b c d, the Lines f b and g i will always be to each other, as 3 to 4; as is very eafily fhewn in feveral Inclinations, if you have two Scales, each with fmall Divisions on them; which Divisions in the one must be to the Divisions in the other, as 3 to 4 : For the greater Line g i will always contain as many of the great Divifions, as the leffer Line f b will contain of the fmall ones. The Angle, which f b makes with the Plane a b c d, is the Complement of the Angle of Incidence to a right

a right Angle, and the Angle made by the Line gi with the fame Plane is the Complement of the Angle of Refraction to a right Angle; therefore gb and gi are the Secants of the Complements of the Angles of Incidence and Refraction, which have a conftant Ratio to one another, as was to be confirm'd by this Experiment.

We have hitherto confider'd a Ray of Light 642 going out of a rarer into a denfer Medium, but the fame conftant Proportion of the Sines mention'd, N°. 639, holds good alfo in the contrary Motion of the Rays; the Angles A C N, M C B (*Plate* IV. Fig. ς .) are not chang'd, whether the Incident Ray be A C or BC*. In that Cafe, if * 626 BC be the Celerity of the incident Ray, C A will be the Celerity of the refracted Ray; for the Motion of the Ray, going out of X into Z, is retarded in the fame Manner by the Attraction towards the Medium X, as it is accelerated in the contrary Motion.

CHAP. VII.

Of the Refraction of Light, when Mediums are Jeparated by a plane Surface.

DEFINITION I.

R Ays proceeding from the fame Point, as a 643 Center, and continually receding from each other, are faid to be divergent.

DEFINITION II.

Those diverge more, which make a greater 644 Angle with each other.

D 4

DEFI-

DEFINITION III.

645 The Point, from which the divergent Rays proceed, is call'd the Radiant Point.

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646 The more the Rays diverge, supposing them at equal Diffances from each other, the less is the ra-

647 diant Point distant from them; and so on the contrary, the Rays are often mov'd by Refraction, as if they came from a radiant Point, tho they do not really proceed from such a Point; that is, if the Rays should be continued or produced back the Way from whence they come, they would meet in one Point. In that Case also, the Rays are faid to be divergent.

DEFINITION IV, and V.

648 Rays which concur in one Point, or would concur if they were continued, are faid to be convergent;
649 and thefe are more convergent, which make greater Angles between themfelves.

DEFINITION VI.

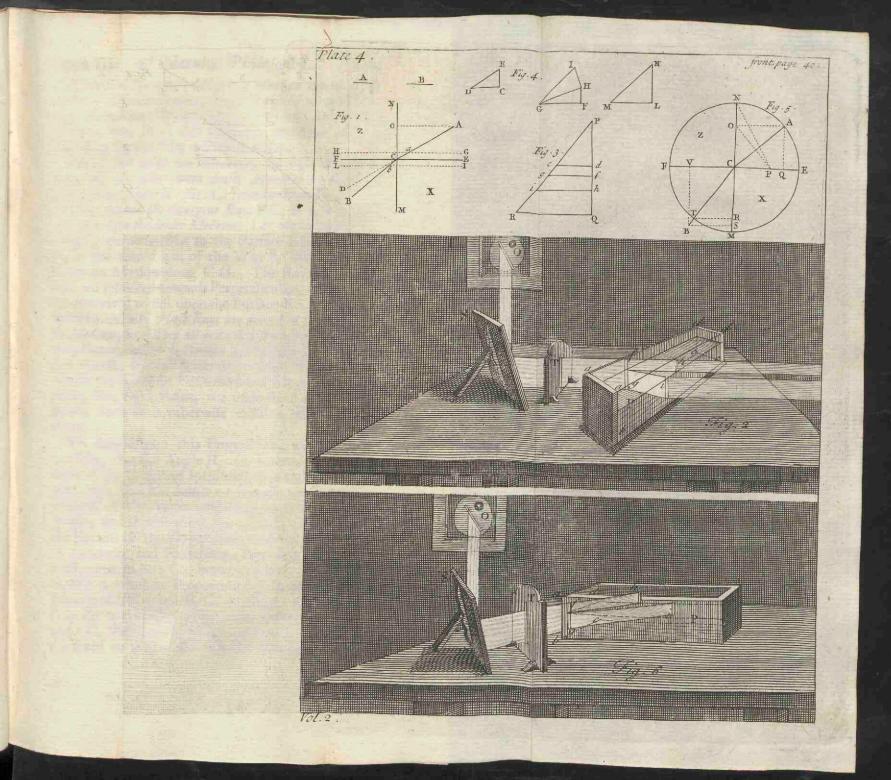
650 The Point of the Concourse of converging Rays, is call'd the Focus.

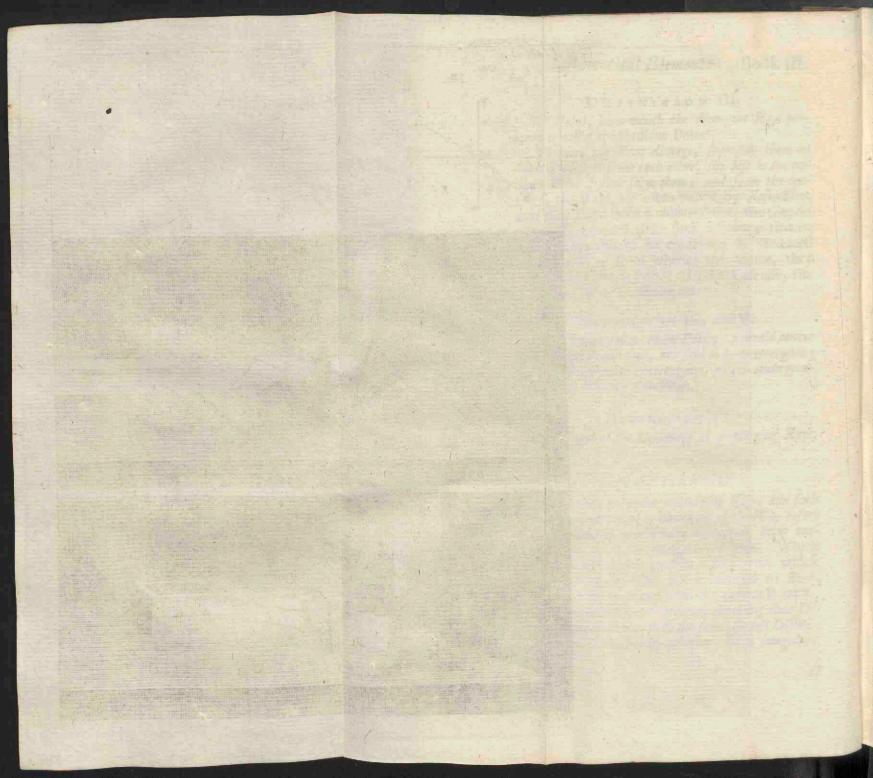
DEFINITION VII.

651 The Point, in which converging Rays, and fuch as (being intercepted or turned out of the Way before their Concourfe) would have concurred, being continued, is called the imaginary Focus. Which Name is also given to that Point from which those divergent Rays are conceived to flow,
* 647 which do not proceed from the radiant Point*.
652 The more the Rays converge, supposing their Di-

652 The more the Rays converge, supposing their Distance from each other to be the same, the less Distance from them is the Focus, whether real or imaginary.

If





If parallel Rays pafs out of any Medium into a-653 nother of different Denfity, they will also be parallel after Refraction: Because they are all equally inflected; for, in all this Chapter, we speak of Mediums separated by a Plain Surface.

Let X and Z be two Mediums, the last more 654 rare, and the other more dense, separated by the Plane E.S., (Plate V. Fig. 1.) from the Point RC, let there proceed the divergent Rays PC, Ro, Rn, and enter into the denfer Medium: Let one of them be RC, perpendicular to the Surface ES; this * 629 laft is not turn'd out of the Way *, but continues its Motion along CG. The Rays Ro, Rn, are refracted towards Perpendiculars, which are conceiv'd to fall upon the Surface ES in the * 624 Points o and n*. These Rays are moved in the denfor Medium, as if they all proceeded from the imaginary Focusr, which is farther distant from the Surface than R., if the Rays are not too much scattered; which yet is not to be underftood Mathematically; for, by a Point, we understand a small Space, fuch as is otherwife call'd a Phyfical Point.

To demonstrate this Proposition, we must confider, that the Angle R o C, is the Complement of the Angle of Incidence to a right Angle; and, that the Angle r o C is also the Complement of the Angle of Refraction to a right Angle; and therefore that the Lines R o, r o, are the Secants of the Complements of the Angles of Incidence and Refraction, supposing the Semidiameter to be o C; between which Secants, there is a constant Proportion *. In the small * 641 Divergence R o and R C, as also r o and r C, 639. they do not fensibly differ, and, between R C and r C, the Ratio is always constant; that is, r is fixed as well as R, tho the Inclination of the

the Ray be chang'd: And therefore Rn is refracted along n A, as if it had proceeded from r.

- 655 If the Rays are too much difpers'd, this Demonfiration will not ferve, and the Place of Concourfe r cannot be taken for a Point : In this Cafe a little Gircle must be imagin'd there, into which all the Rays concur, which will be the greater, the greater the Angle is that the divergent Rays make.
 - 656 If fome Rays, proceeding from R, are not too much difperfed, but fall very obliquely on the Surface ES, they will be refracted, as if they proceeded from a Point not very remote from the Point r: As is plain from what has been faid.
 - 657 The Rays, fuch as A n, B o, GC, which come converging from a denfer Medium X into a rarer Z, concur fooner than they wou'd do, if they fhou'd
- * 625 continue their Motion in a denfer Medium *, that is, become more convergent, and the real Focus is lefs diftant than the imaginary one. In this Figure, the imaginary Focus is r, and the
 * 625 real Focus R *. This Proposition therefore is
- properly the Inverse of the Proposition of N°.654.
 * 626 and therefore * both are prov'd by the fame Experiment.

Experiment I. Plate V. Fig. 2.] Thro' the Ball G, which is moveable in the Window-fhut, and has an Hole going thro' the Middle of it, let a cylindric Beam of the Sun come into the dark Room, and be reflected horizontally by the Looking-Glais S; let it then go through the Convex Lens of Glais that is fix'd in the Board or Stand T, and the Rays will meet together at R, and beyond R will move as if they proceeded from that Point; which therefore is the radiant Point.

Z

Convex

Convex Lenfes of Glafs are very common; we fhall hereafter mention their Properties*, it be-*697 ing needlefs to do it in this Place : Now we only want a radiant Point, and it is enough to fhew how to make it.

Take a Trough Box, P, whole Side *a bcd* is of Glass, and let it be fill'd with Water: The Rays, which diverge from the Point R, become less diverging when they go into the Water.

When convergent Rays, as HD, Ip, Lq (Plate 658 V. Fig. 1.) having their imaginary Focus at f, go from a rarer Medium Z into a denser X, they become less convergent *, and concur in the Focus F, * 624 which is more diffant from the Surface ES *, * 652 as appears by applying here the Demonstration given in N°. 654.

Rays proceeding from the Point F, and going out 659 of a denser Medium into a rarer, become more diverging, and move as if they came from f, which Proposition is the Inverse of the foregoing, and is confirm'd by the fame Experiment *. *626

Experiment 2. Plate V. Fig. 3.] Take the fame Box B, as was ufed in the former Experiment; but here let in two Beams of the Sun into the dark Room thro' two Holes in the movcable Plate in the Window-fhut; let them be both reflected horizontally, and transmitted thro' fimilar convex Lenfes; thereby the Rays of the Beams will become convergent, having their Focus's at the fame Diftance; but if the converging Rays are made to run into the Water in the Box thro' the Side abcd, they will be collected at a greater Diftance; which will plainly appear by comparing together the Situation of the Points F in the Air, and f in the Water.

CHAP.

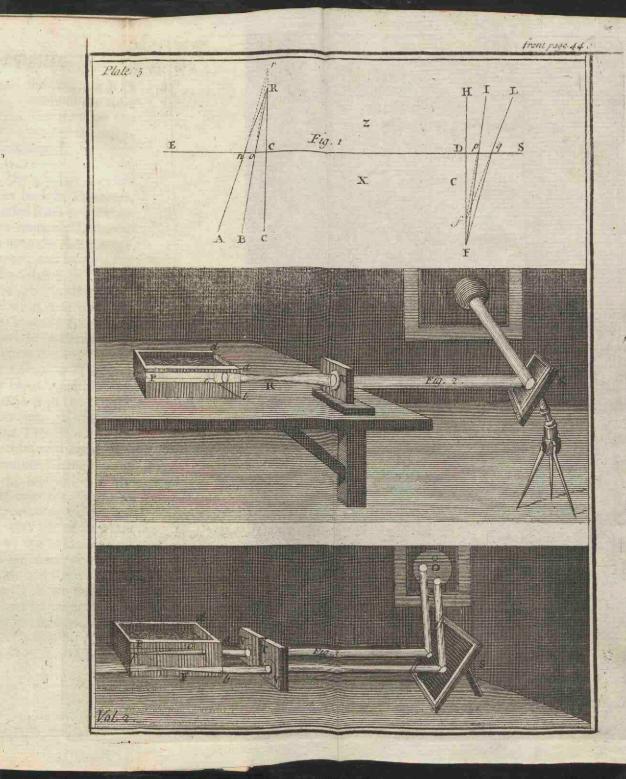
CHAP. VIII.

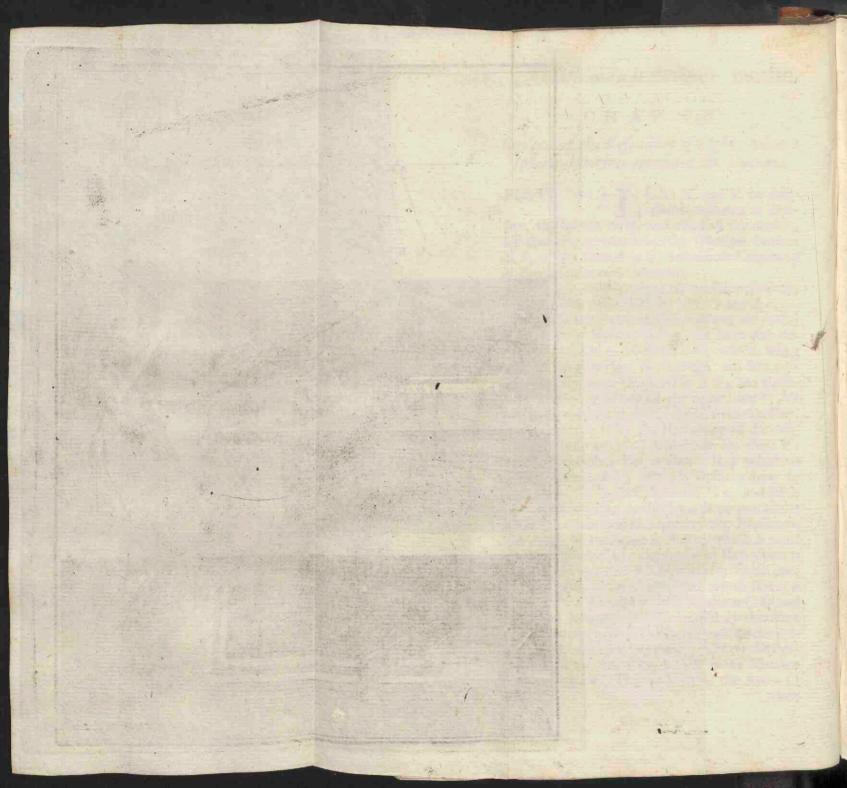
Concerning the Refraction of Light, when a spherical Surface sparates the Mediums.

Plate VI. Fig. 1.] **L** E T X and Z be Mediums differing in Denfity, the laft the rarer, and the firft the denfer; let them be feparated by the fpherical Surface E S, whole Center is C, and whole Convexity is towards the rarer Medium.

560 To begin, by examining the most fimple Cafe; Let us fuppofe parallel Rays, as B O and A n, going out of a rarer into a denfer Medium, and falling upon a convex Surface, fuch as we have just defcrib'd; let one of them be B O, which, being continu'd, goes thro' the Center, and falls perpendicularly upon the Surface E S; and there-

* 629 fore is not turn'd out of the right Line *. All the Rays, which are not too diftant from that Ray, come nearer to it by the Refraction of the denfer Medium, and are collected into one Point F: As for Example; Let A n be a Ray which is refracted along n F; thro' the Point n draw to the Center C the Semidiameter Cn, and let it. be continu'd to p; as this Line is perpendicular to the Surface which separates the Mediums, the Angle of Incidence is A n p, which is equal to the Angle n CO; the Angle of Refraction is C n F. If the Arc n O be a very fmall one, these Angles are as their Sines, whose Ratio is * 639 conitant *. Therefore these Angles n C O and C nF are increas'd and diminish'd in the fame Ratio, as long as their Difference, which the Angle n FO, which confequently follows the Proprotion of the Arc nO, which is the Measure of the Angle n CO; as long as the Arc n Odoes





does not exceed 15 Degrees, the Angle $n \neq O$ is increased and diminished fensibly in the same Ratio, as the said Arc; and therefore all the Rays, between A n and BO, do by their Refraction meet sensibly in the Point F.

Experiment 1. Plate VI. Fig. 2.] Let a cylindric Beam of the Sun, of an Inch in Diameter, made up of innumerable parallel Rays (which are fo on Account of the Sun's immenfe Diftance) be let into the dark Room, as in the first Experiment of the foregoing Chapter, and be reflected horizontally by the Looking-Glass S.

Fill with Water the Trough P, which is about 3 Inches high, and as wide, and one Foot long. Let it have a Glafs V made faft in one of its Sides: This Glafs muft be a Portion of a Sphere, thin, and every where of the fame Thicknefs, fuch as the Cryftal of a Pocket-Watch.

The Convex Part of the Glafs V muft be outwards, that the Water next to it in the Trough may put on a fpherical Surface. If the Rays above mention'd go into the Trough thro' this Glafs, becaufe the Glafs is thin, and has its Surfaces parallel, there is no fenfible Change in the Motion of the Light by the Refraction of the Glafs, and the Light enters into the Water in the fame Manner, as if there was no Glafs. Let the Trough be fo difpos'd, that one of the Rays may pafs thro' the Center of the fpherical Surface, and the others will come nearer and nearer to it, and at laft concur with it at F.

Plate [VI. Fig. 3] Again let X be the denfer Medium, and Z the rarer, and let them be feparated by the fpherical Surface E S, whole Center is at C, and whole Convexity is towards the rarer Medium : From the radiant Point R let Rays proceed,

ceed, and enter into the denfer Medium thro' the Surface above-mention'd; fo that of those Rays, that which is express'd by R O, being continued, may pass thro' the Center; this Ray is not refracted as it goes into the Water, and all the reft of the Rays come towards it by the Refra-Stion, and when they are not too divergent are colletted into one Point as F, in the fame Manner as was faid of parallel Rays; with this Difference, that the Focus F in that Cafe is more diftant. The fame Demonstration will also ferve here, as relates to parallel Rays, which is built upon this Foundation, that the Angle of the Incidence increases in the same Ratio as the Arc n O. which does also obtain here, when the faid Arc does not exceed 15 Degrees. Let Rn be a Ray of Light, and from the Center C thro' n draw Cnp: The Angle Rnp will be the Angle of Incidence : let it be divided into two Parts by the Line nq, parallel to the Line ROC; the Part p n q is equal to the Angle n CO, which is meafur'd by the Arc n O, and which therefore follows the fame Proportion as that Arc; and which also the Angle n RO (if it be very fmall) does follow, and is equal to the fecond Part of the Angle of Incidence, which also does wholly increase and diminish in the same Ratio, as the Arc n O; for the Ratio which holds in respect of every Part, taken fingly, will also hold in refpect of the Whole.

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The fame Demonstration may be applied to any diverging or converging Rays, which in any Cafe are refracted in passing thro' a spherical Surface; and which (as appears by this Demonstration) when they diverge but little, have their Focus real, or imaginary, or run parallel. It is enough to have observed this in general.

The

The Focus F of the Rays, that come from R, goes farther off when R is brought nearer; and fo on the contrary. For the radiant Point being brought nearer, if the Point n remains the fame, the Angle of Incidence is increas'd; and as it increafes, fo does alfo the Angle of Refraction F n C, and n F interfects R C at a greater Diftance.

Experiment 2. Plate VI. Fig. 4.] This Experiment differs from the foregoing only in this, that a Cylindric Beam of the Sun, reflected horizontally, muft be transmitted thro' the Convex Lens in the Board T, as was done in the Experiments of the foregoing Chapters, to form the radiant Point R, from which the Rays, going forward, diverging, are collected in the Water at a greater Diffance, than if they had been parallel.

As you move the Board T, the Point R alfo changes its Place : If this Point is farther off from the Surface that feparates the Mediums, F falls nearer to it: On the contrary, if R be nearer, F is farther off.

The radiant Point may be brought fo near to the 664 Surface above-mention'd, that the Focus will recede to an infinite Diftance; that is, that the refracted Rays will run parallel.

Experiment 3. Plate VI. Fig. ς .] Things being difpos'd as in the former Experiment; by removing the Board T, let R be brought nearer to the Trough; and it may eafily be fo difpos'd, as to make the refracted Rays become parallel.

Experiment 4. Plate VII. Fig. 1.] Now if the 665 Experiment be repeated, bringing nearer to the Trough the radiant Point R, the refracted Rays will become divergent; but they will diverge lefs than the

666 the incident Rays. If the Rays, which out of a rarer enter into a denser Medium thro' a convex Surface, be converging and directed towards the Center of that Spherical Surface, they will suffer no Re-* 629 fraction *. But if they be directed towards another 667 Point, fince they are refracted towards the Per-* 624 pendicular*, they will be fo inflected, that the Focus of these converging Rays will always be between the Center of the Surface which leparates the Mediums (to which all the Perpendiculars are directed) and the Point to which the incident Rays tend. That is, if the imaginary Focus of the incident Rays be at a less Distance than the Center, the refracted Rays will be lefs converging : But if this imaginary Focus be beyond the Center, the refracted Rays will be more converging.

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Experiment 5. Plate VII. Fig. 2.] Every Thing being in the fame Manner, as in the former Experiments, it is eafy to confirm these Propositions by Experiments; for the Board T may be fo dispos'd, that the convergent Rays shall enter the Water, fo as to have their imaginary Focus at any Distance in it.

668 From what has been faid hitherto it is eafy to determine what happens in a contrary Motion

* 626 of the Rays^{*}; that is, the Motion of the Rays from a denfer into a rarer Medium, the convex Surface remaining towards the rarer Medium.

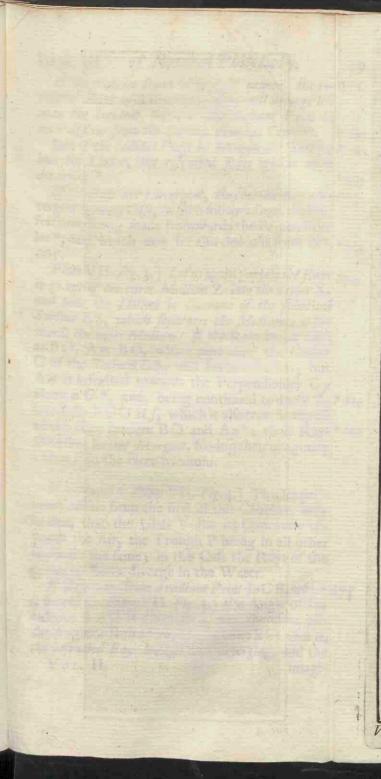
Parallel Rays, after Refraction, concur in a *644 Focus *:

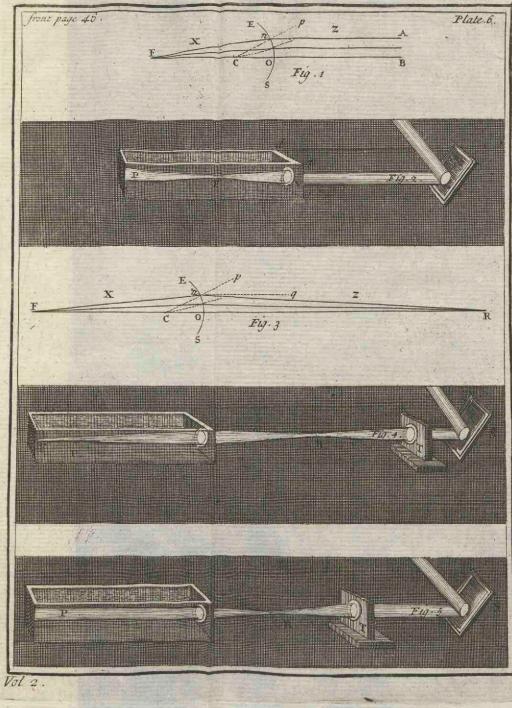
669 The Rays that come from a radiant Point meet
* 661 alfo in a Point or Focus *, and as that Point is brought near, the Focus goes farther off; and
* 661 alfo in a Point or Focus goes farther off; and

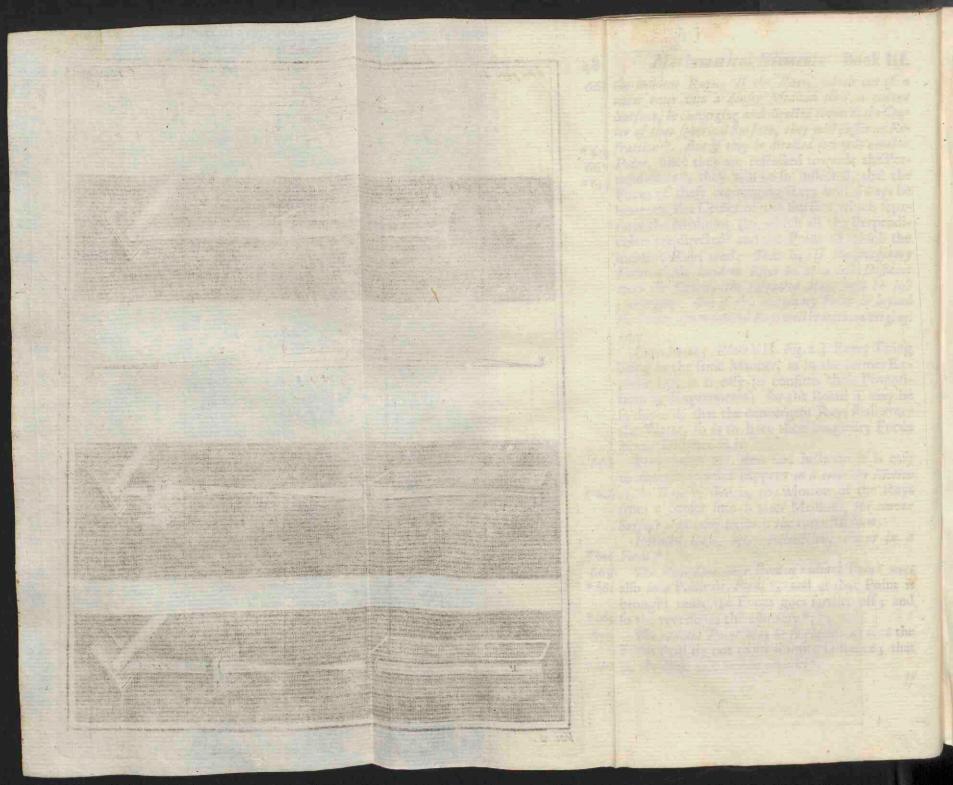
* 663 fo the reverse on the contrary *. 670 The radiant Point may be fo placed, as that the

Focus shall fly out to an infinite Distance; that * 660 is, the Rays will become parallel*.

If







If the radiant Point be brought nearer, the re-671 fracted Rays will diverge; they will diverge less than the incident Rays, if the radiant Point be more diffant from the Surface than the Center*. *667

49

But if the radiant Point be between the Surface 672 and the Center, the refracted Rays will be more divergent *. *667

If the Rays are convergent, they become more con- 673 vergent in every Cafe, which follows from the Refraction being made fromwards the Perpendicular*, and which may be alfo deduc'd from N°. * 625 665.

Plate VII. Fig. 3.] Let us again fuppose the Rays 674 to go out of the rarer Medium Z into the denser X, and that the Hollow or Concave of the spherical Surface ES, which separates the Mediums, is towards the rarer Medium. If the Rays be parallel, as BO, An, BO, which goes thro' the Center C of the Surface ES, will not be refracted; but A n is refracted towards the Perpendicular C p along n G*, and, being continued towards Z, *624 interfects BCO at f, which is also true in respect of the Rays between BO and A n*; those Rays *662 therefore become divergent, having their imaginary Focus f in the rarer Medium.

Experiment 6. Plate VII. Fig. 4.] This Experiment differs from the first of this Chapter, only in this, that the Glass V has its Concavity towards the Air, the Trough P being in all other Respects the fame; in this Case the Rays of the cylindric Beam diverge in the Water.

If Rays come from a radiant Point in CB, which 675 is beyond C (Plate VII. Fig. 3.) the Angle of Incidence A n C is diminifh'd, and therefore also the Angle of Refraction G n p grows less; that is, the refracted Rays become more diverging, and the Vol. II. E imagi-

* 646 imaginary Focus f comes nearer to C *; till by the radiant Point coming nearer, at length it coincides with the imaginary Focus at C; for in
* 629 that Cafe the Rays undergo no Refraction*.

50

676 If the radiant Point comes nearer between C and O, the imaginary Focus is farther from O than the radiant Point, for it is always between that Point and C, by reafon of the Angles of Re* 624 fraction being lefs than those of Incidence*.

Experiment 7.] Things being, as in the former Experiment; if you use the Board, with the Convex Lens in it, to form a radiant Point, the Experiments proving these Propositions may be easily made.

677 If the Rays are converging, and the Point of Concourfe be in the denfer Mediums, near the Surface which feparates the Medium, the refracted Rays will also converge, but lefs than the incident Rays.

678 If the imaginary Focus of the incident Rays recedes more and more from O, that is, if they converge lefs, therefratted Rays will also converge lefs; until, by the receding of the imaginary Focus, the Rays become parallel.

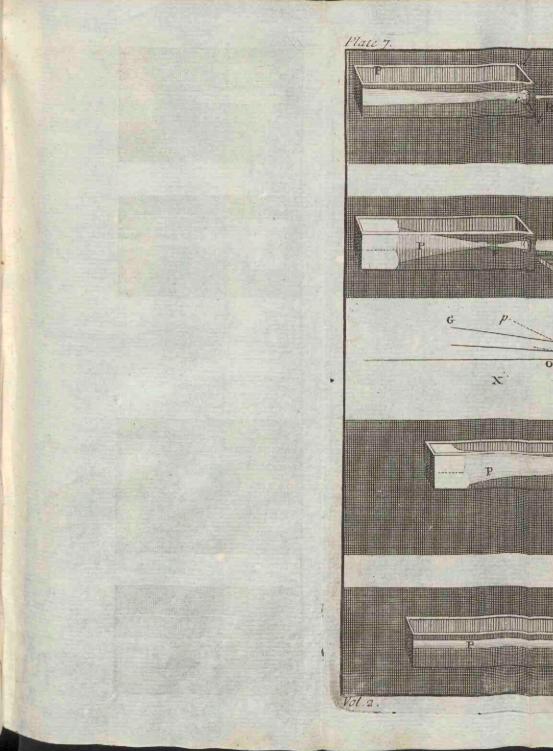
679 If the imaginary Focus recedes yet farther, the refracted Rays become divergent.

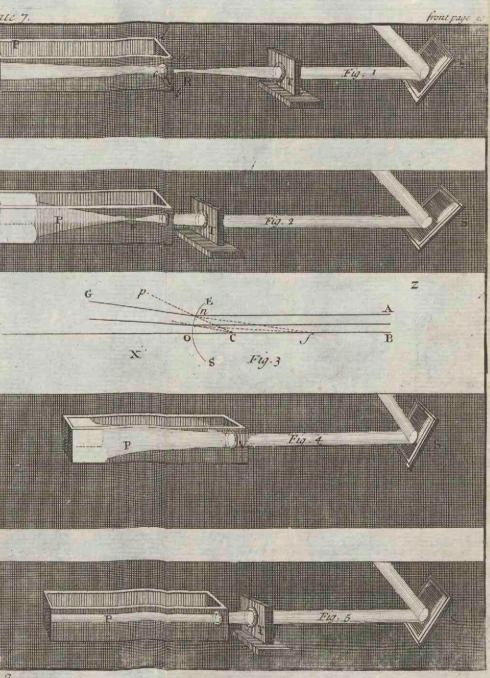
Experiment 8. Plate VII. Fig. 5.] Here the Board muft be fo plac'd in respect of the Trough, that the Rays may enter the Water converging; and the Phænomena above-mention'd may be feen, according as you remove the faid Board.

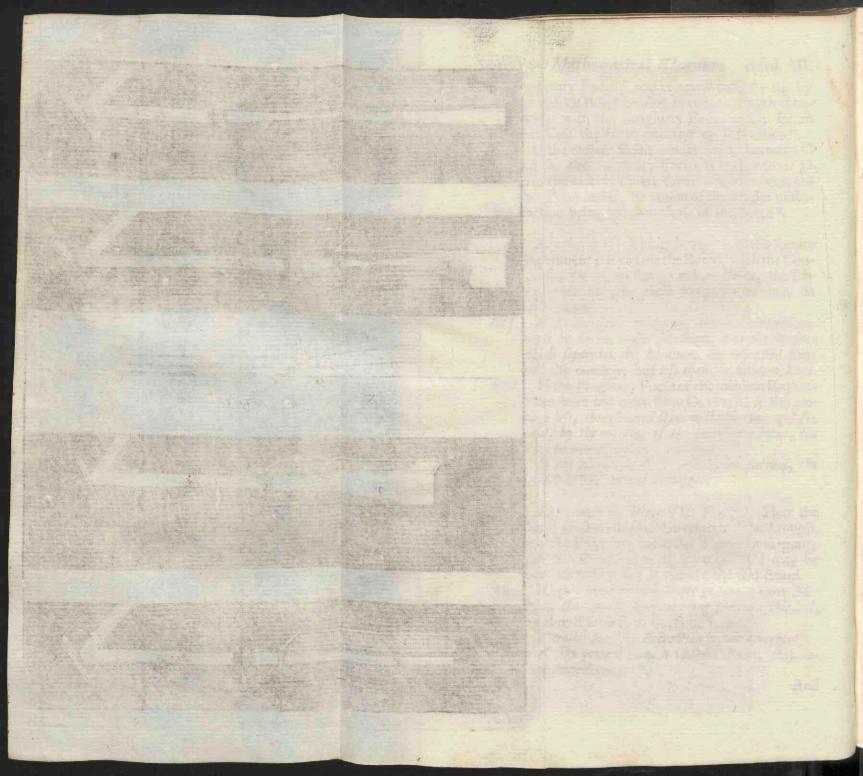
680 Rays which from a denser go into a rarer Medium, the concave Surface being towards this last, are almost subject to the same Laws.

678 Parallel Rays by Refraction become divergent*.
 626 If they proceed from a radiant Point, they be 681 come more divergent*.

And







And they diverge still more and more, as the radi- 682. ant Point is brought nearer *. * 678

SI

Converging Rays, which tend to the Center of 683 the fpherical Surface, undergo no Change*. * 629

If they converge more or less, the imaginary Fo-684 cus of the incident Rays is always between the Center of the Surface, which separates the Mediums, 685 and the Focus of the refracted Rays, * which may * 675, recede in infinitum, so as to make the refracted 676 Rays become parallel *. * 674

We have hitherto confider'd fuch Rays as are but little inclin'd to the Surface which feparates the Mediums; for we have mention'd fuch incident Rays as diverge but little, and one of which is perpendicular to the Surface that feparates the Mediums. The fame Propositions hold 686 good in oblique Rays, yet in that Cafe all the Rays are inflected, which does not happen fo in direct ones; for the Ray which is perpendicular to the Surface is not inflected. Oblique Rays alfo undergo a greater Refraction, that is, they are more inflected either to or from each other, 687 than direct ones, fuppofing the Circumftances the fame.

Plate VIII. Fig. 1.] Let Z be a rarer, and X a denfer Medium, E S the Surface feparating the Mediums, and having its Center at C; and the parallel Rays A n, Bm will come together at F. The Raysthat proceed from the radiant Point R will concur at F.

Plate VIII. Fig. 2.] If the Surface be turn'd fo as to have its Concavity towards the rarer Medium, the parallel Rays A n and B m will have their imaginary Focus at f; but its Diffance from the Surface E S, as alfo that of the above-mention'd Foci F and f in the Figure 1. is lefs, than if the Rays were direct.

E 2

688 All the Points of a lucid Body are radiant Points, and have each their particular Focus, which ferves to explain the following Experiment, made to confirm what has been faid of the oblique Rays.

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Experiment 9. Plate VIII. Fig. 3.] Take the Trough P mention'd in the first Experiment, and fill it with Water; let the Glass V be all cover'd, but a circular Part in the Middle of about half an Inch; have in the Trough a moveable white Plane T. If the Candle A befet at the Diftance of 3 or 4 Foot from the Trough, let the Plane T be mov'd backward and forward in the Water, and when it is come to the Diftance of the Foci of the Flame A, that Flame will appear exactly represented on that Plane, all the Foci forming a Picture. And this holds good, whether the Rays from the Candle fall obliquely or directly upon the Glass V, only that when the Rays are oblique (the Diftance of the Candle A remaining the fame) the Diftance of the Plane T from V must be less. In this Case also the Candle and the Glass V will not be in the fame right Line as the Reprefentation of it, as it happens when the Rays are direct, by which the Proposition of Nº. 686. is confirm'd.

- 689 The Candle is reprefented inverted, because the Rays, which proceed from different Points, interfect one another as they go thro' V, as may be plainly feen by looking at the first Figure. For which Reason, if there are two Candles, as A and B, the Representation of the last will be at b, and that of the first at a.
- 690 All the Changes that happen in Light, which we have taken Notice of in this Chapter, are fo much the more fensible, as the Surface separating the Mediums is more curve; that is, a Part of a leffer Sphere.

CHAP.

CHAP. IX.

Concerning the Motion of Light thro' a more denfe Medium. Where we shall take notice of the Properties of Lenfes.

THE Use of Glasses is common, they are 691 more dense than Air, and the Rays out of Air pass thro' the Glass into Air again. According to the feveral Surfaces that terminate the Glass, Light undergoes different Changes as it moves in it; which to determine, the Glasses, or any Mediums encompass'd with a rarer Medium, and terminated with different Surfaces, must be examin'd. If we confider only plane and spherical Surfaces, there are fix Sorts.

The Medium may be plane or flat on both Sides. 2. Plane on the one Side and convex on the other. 3. Convex on both Sides. 4. Plane on one Side and concave on the other. \int . Concave on both Sides. 6. Laftly, It may be terminated with a concave Surface on one Side, and a convex one on the other.

DEFINITION I.

If the Glass be made use of, and is not very 692 thick, Glasses, whose Figure is mention'd in the last five Cases, are call'd *Lenses of Glass*.

In the fecond and third Cafe, a Lens is faid to be convex; and if we diftinguish those two Cafes, in the fecond it is call'd plano-convex. And so in the fourth Cafe, it is faid to be planoconcave; tho both this Cafe, and the fifth, is generally referr'd to concave Lenses. But a concavo-convex Lens is referr'd to the concave or convex Lenses, according as the one or the other Surface is predominant; and that is faid

E 3

to be predominant which is most curve; that is, which is a Portion of a lefs Sphere.

DEFINITION II.

- 693 In every Lens, or Medium, terminated in any Manner above described, a right Line, which is perpendicular to the two Surfaces, is call'd the Axis. When both Surfaces are spherical, the Axes goes thro' both their Centers; but if one of them be plane, it falls perpendicularly upon that, and goes thro the Center of the other.
- 694 In the Passage of Light thro' a Medium, terminated by two plane Surfaces, the Direction of
- * 628 the Rays is not changed *; which is the Cafe in plane Glasses.
 - 695 It is the Property of all Sorts of convex Lenfes, that the Rays in their Paffage thro' them are inflected towards one another; so much the more, as
 - 696 the Convexity is greater: And fo concave ones, that the Rays are deflected from one another, according as the Concavity is greater. For the Direction of
- * 694 the Rays thro' a plane Glass is not changed ; * but, by inflecting one or both Surfaces, another Direction is given to the Rays: There are more inflected towards the Axis of the Lens, by reason of the Convexity of the Surface of the Glafs, and, by making the Surface concave, they are deflected from the Axis; as is plain in every Cafe, by comparing the Inflection in the plane Surface that is perpendicular to the Axis, with the Inflection in the fpherical Surface at any Diftance from the Axis. And the Difference of their Inflexions, that is, the Change of the Direction of the Rays, increases, as their Distance from the Axes does ; and it is to be observ'd in every Direction of the Rays, as well in oblique Rays, as in direct : but the Changes are greater in oblique Rays, because the Angles of Incidence are greater. From which 697 we deduce the following Properties of Lenfes.

That

That parallel Rays, by paffing thro' a convex 698 Lens, concur in a Focus.

That diverging Rays either diverge less, or run 699 parallel, or lastly, converge; in which Cafe, the radiant Point receding, the Focus comes nearer, and fo on the contrary : But this is the Cafe, when the radiant Point is farther diftant from the Lens than the Focus of parallel Rays.

Laftly, That converging Rays converge more, 700 when the Light emerges out of the Lens.

The same Things are observable in oblique Rays; 701 concerning which it is to be noted, that the Diftances of the Foci of the emerging Rays, are less than in the direct, and the other Changes more fensible *.

All these fame Things may be deduced from examining the double Refraction in the Entrance and Emersion of Light. And this double Refraction is visible in every Cafe, by the following Experiments; by which the aforefaid Properties of convex Lenfes are confirm'd.

Plate VIII. Fig. 4.] Take feveral Boxes like P, 702 with Water in them, and thro' which Light is transmitted thro' the Glasses V and V, which are placed in the oppofite Sides of the Box, and are diftant from one another about one Inch. These Glasses are thin; in the Box which reprefents a convex Lens, on each Side there is placed one, like that of the first Experiment of the former Chapter, which are fo difpofed as to have their Convexities without the Box. When a plano-convex Lens is to be reprefented, on one Side there is a plane Glafs. A concave-convex Lens is reprefented by two fpherical Glaffes, that are Portions of different Spheres, and the Convexity of the Portion of the greater Sphere, muft be turn'd towards the Infide of the Box.

Whilft the Light paffes thro' thefe Boxes, the Changes of the Light are visible to the Eye in its

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* 697

its Entrance into, and Emerfion out of a denfer Medium, and by the Affiftance of thefe, all Things, relating to convex Lenfes, are clearly demonstrable.

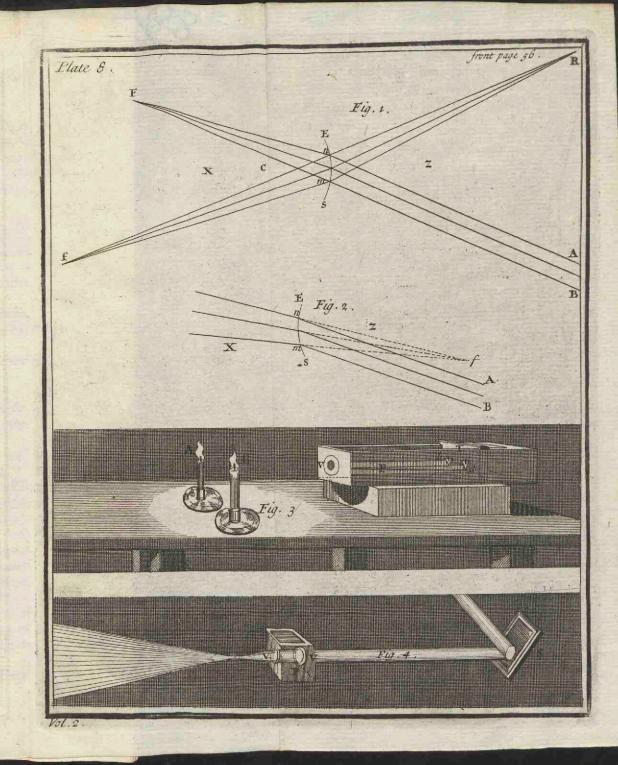
Experiment 1.] Ex. gr. Let P be one of the fore-mentioned Boxes, with the fpherical Glaffes V, V, the Convexities being placed outwards; let it be fill'd with Water; in a dark Room let a cylindric Beam of the Sun be horizontally reflected from the Looking-Glafs S; let this Beam enter the Box; the parallel Rays, of which it is form'd, will be inflected towards one another, and will converge ; at their Emerfion on the other Side, they will converge more, and concur in F. Experiments may be made of the incident Rays, which diverge or converge, by using the Board with the convex Lens, as in the Experiments of the former Chapter.

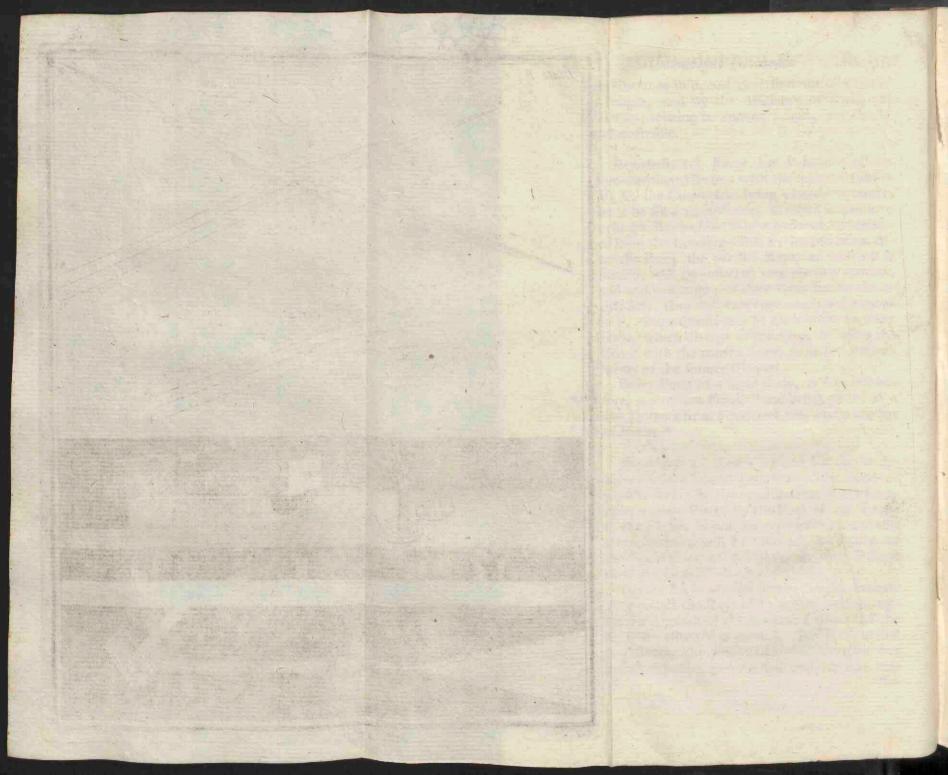
Every Point of a lucid Body, as was faid be-*688 fore, is a radiant Point, * and being placed at a due Diftance from a convex Lens, every one has * 698 its Focus. *

Experiment 2.7 Let a lighted Candle be removed from a convex Lens beyond the Focus of parallel Rays; at the oppofite Part of the Lens, upon a white Plane, by the Foci of the Points of the Flame, it will be represented; and this Representation will be inverted, by reason of the Interfection of the Rays in their Paffage thro' the Glass.

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Convex Lenfes are alfo burning Glaffes, becaufe * 698 they collect the Rays of * the Sun, which, upon the Account of the immense Distance of the Sun, are efteem'd as parallel. But Rays united in a Focus, (by reafon that the Fire, that was before difperfed, is now collected, and by reafon





fon of the Motion of the Fire according to various Directions,) do burn vehemently.

Experiment 3.] Take a convex Lens of any Magnitude, let it be fo exposed to the Rays of the Sun, that the Axis of the Lens may be in the Direction of the Rays; if any combustible Body be placed in the Focus of the Sun's Rays, it will burn.

When, by reafon of the Magnitude of the Lens, the Rays are not exactly enough collected, before they come to the Focus, they must be transmitted thro' another convex Lens that is lefs, by which they will be reduced into a fimaller Compass, fo as to burn more violently.

As for concave Lenfes, and their Properties, they may be cafily deduced from * what has * 696 been faid.

Parallel Rays become diverging, by passing thro' 704 a concave Lens.

Diverging Rays diverge the more.

Converging Rays activer ge the more. 705 Converging Rays either converge the lefs, or be-706 come parallel, or (as it happens in fuch as converge lefs) go out from the Lens diverging.

All which Things happen to oblique as well 707 as direct Rays, but more fentibly in the first. * *697

Plate IX. Fig. 1.] Boxes, to reprefent the Ef- 708 fects of concave Lenfes, are made in the fame Manner as those that represent the Effects of convex ones.* The Difference is only in the Posi- * 702 tion of the Glasses: In the first Box the Concavities of the two spherical Glasses V, V, are outwards. In the second, instead of one of those Glasses, you have a plane Glass. In the third. you have two spherical Glasses, but Sections of different Spheres; the Section of the greater Sphere has its Convexity outwards, and the on ther its Convexity inwards.

Expe-

Experiment 4.] Parallel Rays of the Sun, reflected horizontally in a dark Chamber by the Looking-Glafs S, muft be transmitted thro' the Box P, which is full of Water, and represents a Lens concave on both Sides : As they go into the Box they will begin to diverge, and as they go out they will diverge more.

The remaining Experiments, relating to concave Lenfes, are made in the fame Manner, as has been faid, in respect of convex Lenfes.

CHAP. X.

Of Vision; where we shall speak of the Make of the Eye.

THE Properties and Laws of the Refraction of Light, that we have explain'd, are of wonderful Ufe in reprefenting Objects to our Mind.

By these Laws, the Objects are beautifully painted in their proper Colours in the Bottom of the Eye; and this Picture, as I shall fay here-*716 after, * is the Occasion of the Ideas which are excited in our Minds concerning the Things that we see.

How this Picture is form'd in the Eye, cannot be explain'd, without examining a new Property of Light; namely, its Divifibility, which is paft our Comprehension.

DEFINITION.

709 A Body that is not lucid, and intercepts the Light, is faid to be opaque.

710 Several among all the opaque Bodies, when exactly polifh'd (except perfectly black Bodies, if 2 there

there are any fuch) have the Property of dividing Light; for they reflect the Light, fo that the Rays from every Point, being flruck back, are divided, and recede every Way, fo that all the fingle Points of the Body become as it were radiant Points, from which Light goes every Way.

Whence we deduce a Method of painting Ob-711 jects upon a white Plane; for all the Points of the enlighten'd Body, from which the Rays come upon a convex Lens, have their Focus of the other Side of the Lens.* The Foci of diftant * 699 Objects, tho not exactly, are fenfibly at the fame Diftance from the Lens; those Objects may by these Foci be represented in the fame Place; which Representation is inverted, (by reason of the Intersection of the Rays as they go thro' the Glass,) and fensible in a dark Place, in which Light comes in no Way but thro' the Lens, and only that Light by which the Objects are represented.

This will do wherever the Lens is placed, and in Refpect of all the Points of Objects, enlighten'd by Rays of Light, from which right Lines without Interruption may be drawn to the Lens; in this Manner the above-mention'd Divifibility in Light may be proved, and the Aptness that Bodies, that reflect Light, have to divide it.

Experiment 1. Plate IX. Fig. 2.] Make an Hole in a dark Place, over-againft feveral Objects that are at leaft ςo Foot off or farther. Let the Hole be V, and let it have a convex Lens in it that collects parallel Rays at the Diftance of about 4 or ς Feet; if a white Plane be placed behind the Lens a little farther from it than that Diftance, all the Objects above-mention'd will be painted upon it in very beautiful Colours. It is to be obferv'd, that the Lens muft be placed in a Po-

a Polition parallel to the Plane; and that by moving the Lensor the Plane, the Diftance muft be found at which the Objects are reprefented most exactly.

This Representation of Objects has great Affinity with that by which the Objects, that we fee, are reprefented in the Bottom of the Eye, as will appear from the Make of the Eye.

The Figure of the Eye, when taken out of the 712 Head, is nearly spherical; only the Fore-part is fomething more convex than the reft.

The Section of the Eye is reprefented Plate IX. Fig. 3.

The Part A A, which is most convex, is tranfparent, and call'd the Tunica Cornea.

The whole Covering of the Eye, except the Cornea, is call'd the Sclerotica, BAAB.

That Part of the Sclerotica, which is next to the Cornea, is call'd the Adnata, ot White of the Eye.

Behind the Cornea, on the Infide, is a Coat call'd the Uvea, which has in its Middle an Hole pp, call'd the Pupil.

The Uvea is made up of concentric circular Fibres, interfected at right Angles by frait Fibres. If the first are fwell'd, the last are relax'd, and the Pupil is leffened or contracted; and a contrary Motion of the Fibres increases or widens it.

In the Middle of the Eye, but nearer the Forepart, there is a transparent fost Body C C, like a convex Lens, whoft Hind-part is more convex than the Fore-part. It is call'd the Crystalline Humour. Its Axis coincides with the Axis of the Eye, that goes thro' the Centers of the Pupil and the whole Eye.

This crystalline Humour is fustain'd by finall Fibres or Threads, which are fix'd to all the Points of its Circumference, and likewife to the Infide

Infide of the Eye: They are inflected in the Form of an Arc, and every one of them is a Mufcle; they are call'd the *Ligamenta Ciliaria*, and two of them are reprefented by l C, l C. They all cohere to one another, and, together with the Cryftalline, make a Separation in the Eye, and divide it into two Cavities, one forwards p p, and the other backwards S S.

The Cavity, that is forwards, is fill'd with a Liquor like Water, call'd the Aqueous Humour.

The hind Cavity is fill'd with a transparent Humour, nearly of the same Density as the Aqueous Humour, but not so fluid, call the Vitreous Humour.

The hind Surface of the Eye within is lined with a Coat call'd the *Choroides*, which is again cover'd with a thin Membrane call'd the *Retina*.

At the back Part of the Bulb of the Eye, a little on one fide, is the Optic Nerve NN fo join'd to the Eye, that the Eye itfelf is, as it were, an Expansion of the Optic Nerve; for the expanded Coats of the Nerve form the Choroides and Sclerotica, and the Fibres, which make up the Retina, concurring, make the Marrow of the Nerve.

The Eye is moved in the Head by feveral Mufcles inferted in the Sclerotica; but we shall not treat of them here: As we only confider the Eye with respect to the Motion of Light, we purposely forbear to take Notice of any Thing elfe.

Rays that proceed from any Point, and enter the 731 Eye thro' the Pupil, go out of a rarer into a denfer Medium thro' a fpherical Surface; and therefore if that Point be at a due Diftance from the Eye, the Rays after Refraction will converge; * in the fame * 661 Manner as in the Experiment of No. 663. (Plate VI. Fig. 4.) in which the Glafs V reprefents the transparent Cornea of the Eye, whill the Water in

in the Trough is inflead of the aqueous Humour; and therefore, *fuppofing only the Cornea and the* 714 aqueous Humour, there will be in the Eye an in-710, verted Picture of the Objects. *

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Experiment 2. Plate IX. Fig. 4.] Let P be the Trough above-mention'd full of Water, and fet in a dark Place, which has an Hole about half an Inch wide, fo difpofed towards enlighten'd Objects at a certain Diftance, that one may fee them thro' the Hole; let the Glafs V of the Trough be applied to this Hole, and upon the white Plane T you will have an inverted Picture of the Objects: By moving the Plane backwards and forwards, you will find the Place where the Picture is most diftinct.

If the abovemention'd Picture, which we have imitated in this Experiment, was to be made in the Eye, it would be at too great a Diffance from the Cornea, and beyond the Bottom of the Eye; and therefore this Diffance is diminifh'd 715 by Help of the Crystalline Humour, which is denfer, but encompass'd with rarer Mediums; for the Rays, converging in the aqueous Humour, pass thro' the crystalline into the vitreous Humour; that is, out of a rare Medium thro' a denser, (which is terminated by two spherical convex Surfaces) into a rare Medium again; by *700 which Motion the Rays converge still more; *

and therefore they concur fooner, and the Picture above-mentioned falls within the Eye.

716 The Objects which, as we have explain'd, are reprefented in the Bottom of the Eye, are painted upon the Retina; and by the Motion of Light the finall Fibres, of which the Retina is made up, are agitated; by which Agitation, the Ideas of the Objects, painted in the Eye, are excited in the Mind. The Connexion between the Ideas and the Motions by which they

they are excited, is unknown to us, * as we faid * 502 before : In determining the Caufes of Senfations, we can go no farther than the Agitation of the Nerves.

The more exact the Picture above described is, the 717 more diffinit will the Objects appear. When the 718 Rays, coming from the same Point, are not exactly united upon the Retina, its Picture is not a Point, but a Spot, which is confounded with the Pictures of the neighbouring Points; in which Case the Vision is confused.

But when, according to the different Diffance of 719 the radiant Point, its Focus is brought nearer, or removed farther off, * it is necessfary that there should * 663 be a Change in the Eye, left the Place, in which the Picture is exact, should fall short of, or beyond the Retina, and fo the Vision should be confused.

But it is very difficult to determine what this Change is, and Philosophers are divided in their Opinions about it: I shall only observe in general, that it is not very probable that the Figure of the whole Eye is changed, in order to put back or 720 bring forward the Retina; and therefore we must expect to find this Change within the Eye.

For if the Figure of the Eye was changed, as this Change must be equally neceflary in all Animals, the Eyes of all Animals would undergo the fame Changes; for the fame natural Effects cannot have different Caufes. Now in the Whale the Sclerotica is too hard to be fubject to any Alteration of Figure. Befides, if there was fuch a Change in the whole Eye, it would arife from the external Prefiure of the Muscles, which would be different in different Positions of the Eye, and only regular in one Situation of it.

If now we examine the Eye within, it will appear neceffary that there fhould be a Change in the Crystalline; which by changing its Place

or

or Figure in the Eye, will produce the defired Effect; for the Rays that fall upon the Retina, before they are united, will be made to unite just upon the Retina, if the Crystalline becomes 695 more convex, * or if (its Figure remaining the

fame) it be brought forwards towards the Cornea.

721 That the Position of the Crystalline Humour is cafily changed, and that it is brought nearer to, or farther from the Retina, its Axis remaining the fame, is plain, because the ciliary Ligaments are muscular: When these Muscles are twell'd, and become shorter, the Hollow which their Inflection makes at C1, C1, becomes lefs; by which means the vitreous Humour is compress'd, and therefore it prefies upon the Crystalline, and pushes it forwards farther from the Retina; which is necessary when we look at near Ob-*663 jects.*

699 From an Experiment, that we shall hereafter 753 mention, it has been demonstrated, that there is

- 753 another Change in the Eye that acts contrary to this; and we fhall fhew what is the Occafion of it. The fecond Change is also to be re-
- 622 ferr'd to the Cryftalline; which (when it is drawn by the ciliary Ligaments, to make it recede from the Bottom of the Eye) becomes alfo flatter, and therefore it must recede farther than if its Figure was unchangeable; that is, the Change becomes more fensible; which we shall shew to * 730 be of Use.*

These Changes in the Eye have their Limits, for which Reason also Objects appear only distinct 724 within certain Limits, which are different Distances 725 according to the Difference of People's Eyes; and very often in the same Man, both Eyes have not the same Limits; which is almost of the same Use, as if the Limits of both Eyes were more distinct from one

one another: For one may fee an Object diffinctly enough with only one Eye. In fome Perfons also the nearest Limit of one Eye is farther off than the farthest of the other: In which Cafe near Objects and distant Objects are feen diflinctly, but the intermediate ones appear confused.

The Picture in the Bottom of the Eye, as has 726 been faid, * is inverted; whence a Queftion a- *714 rifes, why we fee Objects erect? To which we anfwer by asking another Queftion; Whether it is more eafy to conceive the Connexion between an Idea in the Mind, and an erect Figure, than an inverted one? We confels, that we have no Notion of that Connexion in either Cafe: But Experience teaches us, that there is a Connexion between an inverted Picture in the Eye, and the Idea of an erect Object; and further than this we do not know.

If we look at the fame Object with both Eyes, 727 it will appear fingle; but this happens only when the Object is painted in correspondent Points of each Retina; which probably happens from the meeting of the Optic Nerves. For it is obferv'd in all Animals, which fee the fame Object with both Eyes, that the Optic Nerves meet and feparate again before they go to the Brain; but in Animals which fee different Objects with each Eye, the Optic Nerves go feparately from the Eyes to the Brain.

Only one Point at a Time can be feen diffinitly, 728 namely, that which is represented in the Axis of the Eye; if we look at one Point with both Eyes, we must fo direct the Eyes, that their Axes continued shall meet in that Point; which happens when we have our Eyes intent upon any Point.

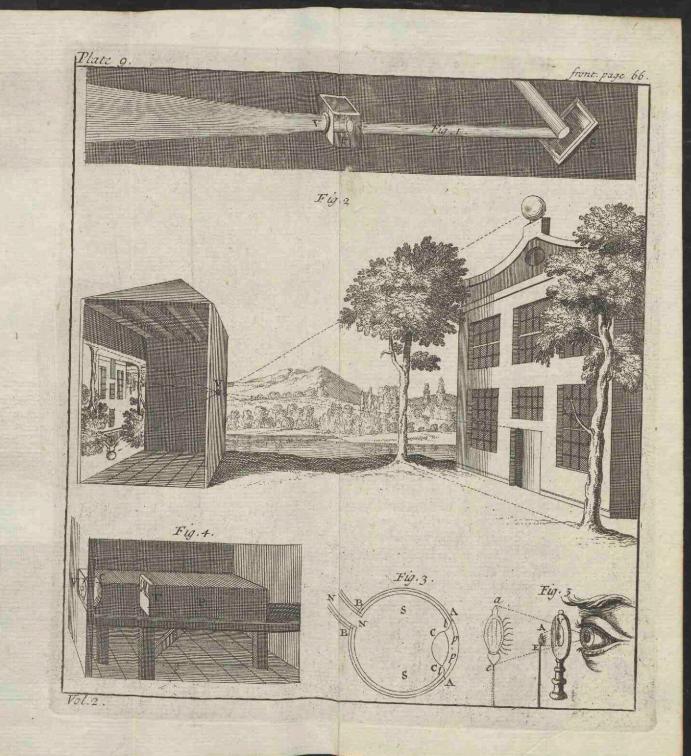
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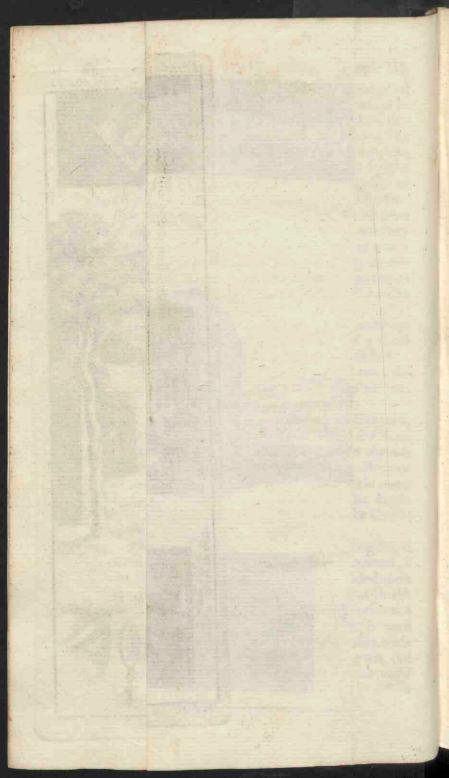
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By this Direction of the Axes we judge of the Diftance of Objects; for the Situation of the Eyes alter according as the Axes make a different Angle, which Angle depends upon the Diftance of the Object: Whence it happens, that without perceiving when we do it, by Ufe we get a Habit of judging of the Diftance of 729 Objects by the Direction of the Axes; which is fenfible to us, becaufe it depends upon the Motion of the Eye, that we feel. Therefore we may fee the Ufe of having two Eyes placed at a certain Diftance from one another; as long as this Diftance of the Eyes bears a fenfible Proportion to the Diftance of the Objects, we can judge of it pretty certainly.

730 We can also judge of leffer Distances with one *721 Eye alone; because in the Variation of a small

- 723 Diffance, the Change in the Eye is fenfible.*
- 731 In great Diftances, if we look at known Objects, we judge from the apparent Magnitude and the Colour.
- 732 It is impossible to judge of very great Distances, except the same Objects be seen from different Places. The apparent Magnitude of an Object depends
- 733 upon the Bigness of the Picture in the Bottom of the Eye, which Picture depends upon the Angle under which an Object is feen, that is, the Angle which is form'd by Lines drawn from the Extremities of the Object to the Eye.
- 734 This apparent Magnitude is to be diffinguish'd from the Magnitude which our Mind attributes to the Object that we see, which last is founded upon the Judgment, whose Foundation is not in the Appearance alone. Every Body knows that the Object appears the less according as it is more distant; whence it happens, that, according to the greater Distance of the Object, if it be known, the apparent Magnitude of the Object is increas'd in





in the Judgment that our Mind makes of it, which we do without any Attention to it. Therefore the fame Object at the fame Diftance appears of a different Magnitude, if we judge differently of the Diftance.

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We have a remarkable Example of this in re-fpect of the Sun and Moon, which appear greater 735 when near the Horizon, than at a greater Height; tho, as is known to Altronomers, the Picture of the Sun in the Bottom of the Eye is the fame in both Cafes, and that of the Moon is lefs when it appears nearer the Horizon; we cannot judge of the Diftance in either Cafe, * but it appears greater near the Horizon, by reafon of the Interposition of the Fields and Part of the Heavens. If we fee the Bodies above-mention'd thro' a Tube, this apparent Distance vanishes, as also the Magnitude which is deduc'd from it. From our Childhood upwards, and fo continually, we join the Idea of Diftance with the Increase of apparent Magnitude, (which is neceffary for making a true Judgment concerning the Magnitude) whereby the Ideas are fo clofely join'd, that they cannot be feparated, not even in those Cafes, in which we know that they lead us into Error. Logicians teach us, how many Errors are to be attributed to Ideas fo join'd.

CHAP. XI.

Of Vision thre' Glasses, and how to correct fome Defects of the Eyes.

A N Object is visible, because all its Points 736 a Point appears in that Place from whence diverging 714 Rays are emitted.

If

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737 If Rays inflected any how enter the Eye diverging, the visible Point will be in the imaginary Focus of the Rays; for the Rays enter the Eye exactly in the fame Manner that Rays would do that came directly from that Focus; and to have them unite upon the Retina, the fame Situation of the Cryftalline is requisite; fo that, in respect to a Speétator, it is no Matter whether those Rays, after Refraction, or these directly enter the Eye; and there will be the fame Motion in the Eye, when ^{*719}, it fits itself for diffinct Visions.*

739 When Objects are feen thro' a plane Glafs terminated with parallel Surfaces, they appear to be nearer than when feen with the naked Eye. Let A (Plate X. Fig. 1.) be a vifible Point; the Rays, going from it and entring the Eye, are between A b and A b; thefe, being refracted in the Glafs V V, move along b c, b c, and go out thro' c d, c d,

- * 628 which are parallel to the Lines A b, Ab: * now because bc, bc, are refracted towards the Perpen-
- * 624 dicular, * cd, cd, fall between b A and b A; that is, they interfect at a, which is nearer than A; therefore the imaginary Focus of the Rays which enter the Eye is at a, in which the Point A ap-
- * 737 pears to be.*

740 That Point also appears to be more enlightened, when feen thro' the Glafs above-mentioned. For all the Rays between Ab and Ab enter the Pupil between d and d; but as the Lines Ab, Ab, are parallel to the Lines c d, c d, and thefe are between those, Ab and Ab being continued would fall beyond d and d; therefore if the Glafs was taken away, the Rays, which now enter the Pupil, would take up a greater Space, and therefore would not all enter the Eye.

Plate

^{73°} A Point appears the more enlightened, the greater 738 Number of Rays coming from it enter the Eye.

Plate X. Fig. 2.] The apparent Magnitude of an 741 Object is increased by the Interposition of a plane Glass; the Object AE is feen by the naked Eye under the Angle AdE; but if you use the Glass V V, by reason of the Refraction thro' A b c d and E b c d, it will be seen under the Angle c d c, which is greater than the last. But yet the Object 742 is not greater in Proportion to the increased apparent Magnitude; * for it appears to be at a less Di-*734 ftance.*

The Increase of apparent Magnitude is fo much 743 the greater, as is the Difference of the Angles A dE and c d c; whose Difference increases as the Intersections of the Lines A d with b c, and E d with b c, come nearer to the Points b and b; which obtains as the Object is brought nearer to the Glass; and therefore it is the greatest of all, when the Object touches the Glass; which shews that Objects, even in the Glass itself, must appear magnified.

And in general, the Eye being placed in a rarer 744 Medium, the Object that is feen in a denfer Medium appears bigger, and is alfo brought nearer by the Refraction.* This is every Day confirm'd by *659 Experience, when we look at Objects in the Water.

Let there be a Point A feen thro' a convex Lens 745 VV, (Plate X. Fig. 3.) and the Rays A b, A b, at c d, c d, will go out more diverging, as if they came from a; * therefore that Point appears to be * 699 at a greater Diffance.* It appears also more en-* 737 lightened; for the Rays come nearer to each other 746 as they go thro' the Glafs, * and are also reduced * 695 into a lefs Space, wherefore a greater Number of them must enter the Pupil.

The apparent Magnitude of an Object is also 747 increased; that is, the Object feen thro' a convex Glass is seen under a great Angle, which ap-F 3 pears

pears from a Sight of the Figure; the Object A E, feen with the naked Eye, is feen under the Angle A d E, (Plate X. Fig 4. and r.) but now it is feen under the Angle c d c, which is greater; (in Fig. 4.) the Rays A b, E b, which are convergent, converge more as they go out of the 700 Lens; * or diverging (as in Fig. r.) they come * 699 to the Eye converging.* Therefore the Object appears magnified, both from the Appearance of * 745 it being farther, * and its Magnitude being in-* 734 creas'd; * and therefore the Magnitude, that we 748 attribute to an Object, does not follow the fame Proportion as the apparent Magnitude; for which Reafon we fhall not delay Time in Demonftrations about it : But we fhall obferve in general,

749 That the Angle, under which an Object is feen thro' a convex Glass, diminishes as the Eye recedes from the Glass; whils the Object is not more distant from the Glass than the Point in which parallel Rays are collected: But if the Object is farther off, the apparent Magnitude is increased as the Eye recedes.

750 After the fame Manner, if the Eye be between the Glass and the Focus of parallel Rays, the Angle abovemention'd is diminiscreased as the Object is farther removed; the Eye being placed at a greater Distance, the fame Angle is increased as the Object is farther removed; in which last Case the Object may be to far removed, as not to be visible beyond the * 754 Glass, as will be faid anon.*

Allo in those Cases in which Objects are vifible, they do not always appear diffinctly.

751 For, that a Point may appear diffinet, it is required that the Rays that proceed from a Point should *736 enter the Eye diverging, * and that the imaginary *727 Focus of those Rays, in respect of the Spectator, 724 should be within the Limits of distinct Vision.*

75² If the Object be removed beyond the Focus of parallel Rays, the Rays flowing from a Point of

of the Object, enter the Eye converging; * which * 699 Cafe is impossible to the naked Eye : In this the Vision is always confused, and the Eye disposes itfelf to as to have the Vision the least confused that may be; from this Disposition we judge of the Diffance, as we do in every Cafe in which we judge of it with only one Eye.* 730

But this Diffance does not always appear to 753 be the fame; whence may be deduced what is faid of the Change of the Figure of the Cryftalline.* For if, fuppofing the Crystalline moveable, * 722 its Figure be unchangeable, in every Situation of the Object and the Eye, in which the Rays, coming converging from a Point, enter the Eye, there will be the leaft Confusion, if the Crystalline comes back towards the Retina as far as may be; fo that in every Cafe there would be the fame Disposition of the Eye, and the fame Judgment concerning the Diftance; which, as has been faid, is contrary to Experience : But if it becomes flatter as it recedes from the Retina, there will be Changes in the Eye, which agree with the feveral Judgments made of the Distance in different Situations of the Object and the Eye.

If in the last Case, in which Rays, coming from a 754 Point converge, the Eye be fo removed, that the Rays unite before they come to the Eye, in all the Points in which the Rays concur, there will be radiant Points; which are the Foci of all the Points of the Object, by which the Object is reprefented inverted upon a white Plane; # and which are * 711 visible Points in respect of the Eye, to which the Rays can come after their Interfection.* In 736 that Cafe the Object appears inverted, because we don't fee the Object itfelf, but its Representation behind the Glass, which we have faid to be inverted.* # 711

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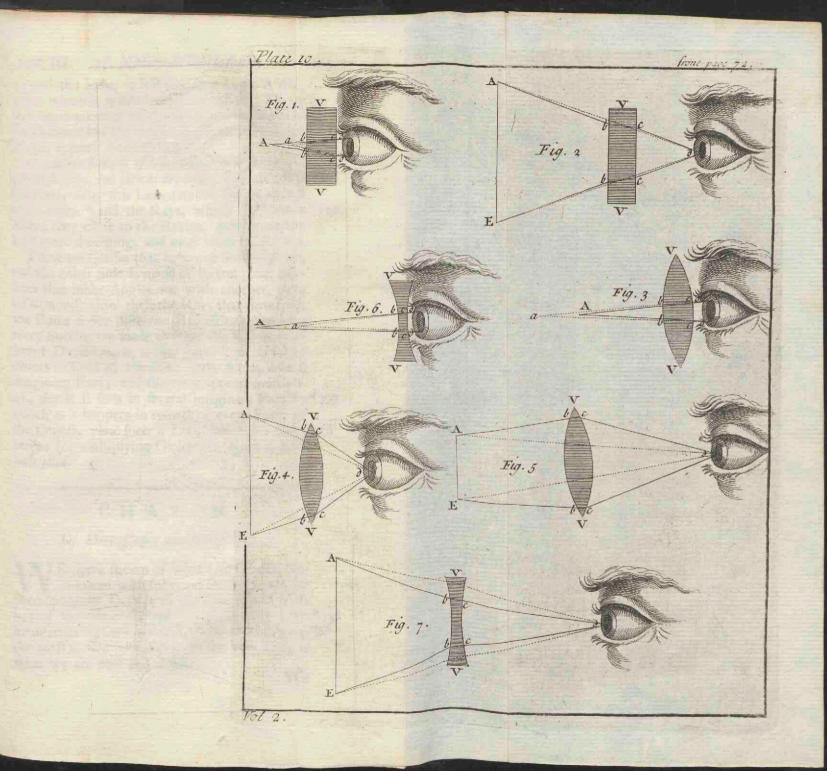
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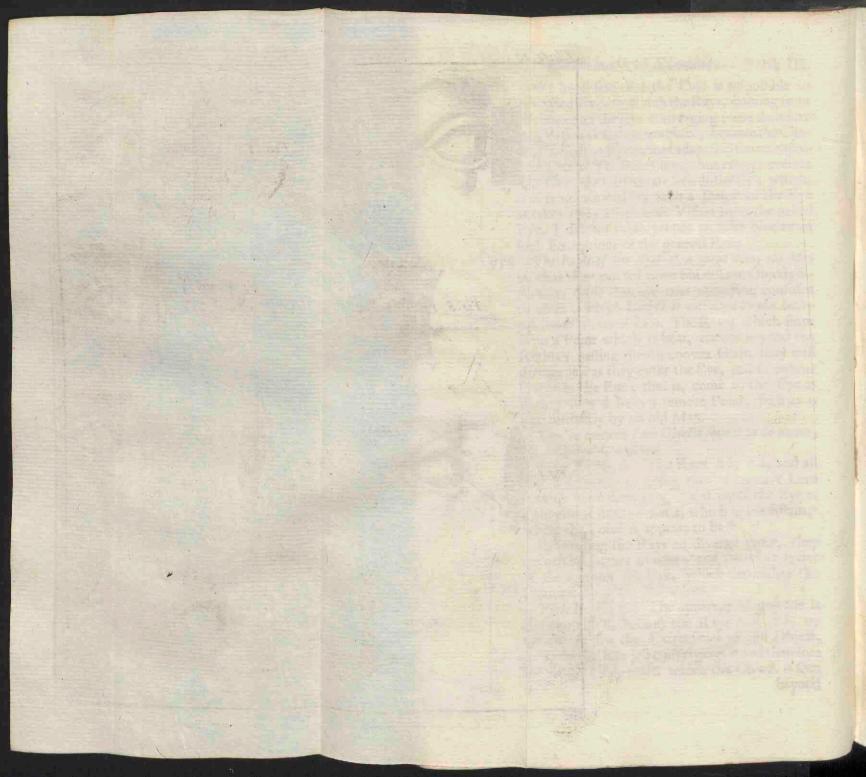
- 755 We have faid that the Cafe is impoffible to the naked Eye, in which the Rays, coming from a Point, enter the Eye converging; and therefore fuch Vision is always confused, because the Conflruction of the Eye cannot adaptitles to an impoffible Cafe : Yet fometimes, but rarely, even in that Cafe, the Objects are seen distinctly; which, as it is occasioned by such a Detect of the Eye as takes away all distinct Vision from the naked Eye, I did not think proper to take Notice of fuch Exceptions of the general Rule,
- 756 The Fault of the Eyes of a great many old Men is, that they can fee none but diftant Objects diflinctly, those that are near appearing confused to them; which Defect is corrected by the Interposition of a convex Lens. The Rays, which flow from a Point which is near, concur beyond the Retina; passing thro' a convex Glass, they will diverge lets as they enter the Eye, and so concur fooner in the Eye; that is, come to the Eye as if they flow'd from a remote Point, such as is feen diffinctly by an old Man.
- 757 Thro' a concave Lens Objects appear to be nearer, lefs enlighten'd, and lefs.

Plate X. Fig. 6.] The Rays Ab, Ab, and all that are between, going thro' a concave Lens
705 become more diverging, * and enter the Eye as
⁶⁵² if they came from a Point a, which is lefs diffant,*
⁷³⁷ where the Point A appears to be.*

By making the Rays to diverge more, they are carried farther afunder; and therefore fewer of them enter the Eye, which diminishes the ⁷³⁸ Illumination of the Point feen.*

Plate X. Fig. 7.] The apparent Magnitude is also diminish'd, because the Rays A b, Eb, by which we see the Extremities of the Object,
706 come to the Eye less converging, * and therefore the Angle c d c, under which the Object is seen beyond





beyond the Lens, is less than the Angle A d E, ander which it is feen by the naked Eye: Therefore upon account of the Diminution, both of the Diftance and the Angle above-mention'd, the Object appears diminifh'd.*

A concave Lens is of Use to those who see no Ob-758 jetts distinctly, but such as are near; such are call'd Myopes; thro' this Lens remote Points appear to be near, * and the Rays, which did concur * 757 before they came to the Retina, now enter the Eye more diverging, and meet upon the Retina.

There are Glaffes that have one Surface plane, and the other Side is made of feveral plane Surfaces that make Angles one with another, (like a Diamond) thro' thefe the Rays that flow from one Point fuffer different Refractions; and by every Surface are made to enter the Eye in a different Direction, as if they came from different Points: That is, the fame Point forms feveral imaginary Foci; and therefore appears multiplied; for it is feen in feveral imaginary Foci *: * 737 which as it happens in refpect of every Point of the Object, thro' fuch a Lens, which is a Poly-759 hedron (or multiplying Glafs) the Object appears multiplied.

CHAP. XII.

Of Microscopes and Telescopes.

WE have shewn of what Use Glasses, terminated with spherical Surfaces, are for correcting the Defects of the Eyes of old Men and of the Short-fighted *. How they ferve * 756 for discovering the smallest Objects, and bringing 758 the most distant (as it were) to the very Eye, is what we are now to consider.

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We have faid that the convex Glaffes magnify * 747 the Objects*; which magnifying depends upon the Refraction of the Rays as they go thro' a convex Lens ; whence it follows, that it is increafed, if, in the fame Circumstances, the Refraction be increased; which Effect may be produc'd, by augmenting the Convexity of the Lens; which will be the more convex, as the Surfaces that terminate it are Segments of a lefs Sphere; which can only be had in very fmall Glaffes.

DEFINITION I.

760 Such small Lenses are call'd Microscopes.

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761 By a Microscope small Objects are vastly magnified ; by which means, Things, which wou'd be invisible to the naked Eye, are very di-

ftinctly feen.

DEFINITION II.

762 The Space feen thro' the Microfcope, that is, the Circle in which Objects are visible thro' the Microscrope, is call'd the Field of the Microfcope.

Experiment 1. Plate IX. Fig. 5.] If we look at the fmall Object A E thro' the Microfcope V, * 745 it will appear magnified at a e *.

There are also compounded Microscopes, made up of two or three Lenfes ; what Foundation they depend upon, will be fufficiently fhewn by explaining one of those which is made up of two Lenses.

Plate XI. Fig. 2.] Take a small Lens that is very convex, as V V, and let the Object A E be plac'd at fuch a Distance from it, that all its * 710 Points shall have their Force beyond the Lens; 699 let the Object be brought fo near, that the Foci * 699 may be remov'd to a e *, and you will there have the Representation of the Object, very much enlarg'd,

enlarg'd, which will be fenfible if you receive it there upon a white Plane *.

Experiment 2. Plate XI. Fig. 1.] The Lens abovemention'd must be made last in the End of a Tube at V; and the other End of the Tube, which is wider, must be cover'd with a very thin Paper CC; the Object A E must be so plac'd, that the Foci of the Points of that Object shall be just at the Distance of the Paper: If then the Object be well enlighten'd, you will have its Representation inverted, visible thro' the Paper at ae. By moving the Object, you will have the true Position that brings the Representation upon the Paper to be distinct.

Plate XI. Fig. 2.] All the Points of the Reprefentation *a e* are radiant Points, and therefore visible*; which, being seen thro'a large Micro-*736 fcrope VV, shews the large Representation *a e* 754 at ac*; that is, the Rays coming from the Ob-*761 ject A E, after the Refractions thro' both the Lenses VV, VV, will enter the Eye, as if they came from an Object at *a e*.

Therefore thro' fuch a compounded Microscope 763 the Object appears inverted, and much more magnified than thro' a single Microscope.

DEFINITION III, and IV.

In this Microscope the smallest Lens, which is 764 next to one Object, is called the Object-Glass, and the other the Eye Glass.

This laft ought not to be too fmall; for the Points of the Reprefentation *ab*, tho they be radiant Points, do not emit Light every Way; only the Rays, which passthro' the Object-Glass, interfect one another in the feveral Points of the Reprefentation *ab*; which Reprefentation therefore will not be visible, unless the Rays that go thro' the Object-Glass do also go thro' the Eye-Glass.

765 Glass. The Field therefore (or the Space that the Microscope can take in) depends upon the Magnitude of this Lens.

The Eye also must be so plac'd, that all the Rays that come to the Eye-Glass, going thro' it, shall come to the Eye; which is done by placing the Eye at d, the Point in which all the Rays, which come from the Center of the Object-Glass, and pass thro' the Eye-Glass, are collected.

Objects appear bright enough thro' Microfcopes, becaufe they are very near the Glafs; and fo the fame Number of Rays pafs thro' a finall Lens as wou'd not pafs at a greater Diftance, unlefs thro' a great Hole: Yet often, where Objects are the most magnified, they must be enlighten'd by 767 Rays collected thro' a convex Lens and thrown upon them.

The Aftronomical Telescope much refembles the compound Microscope.

DEFINITION III.

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⁸ Inftruments, fitted to fee diffinet Objects, are call'd Telescopes.

769 That which we now treat of is call'd the Aftronomical Telescope, because it is not so fit for seeing Objects upon Earth; for it represents them inverted: But Astronomers do not much regard the Position of the Appearance of the Object.

This Telescope confiss of two convex Lenses, the one an Object-Glass, which is plac'd next to the Object, and the other an Eye-Glass, plac'd next to the Eye. By Help of the first, distant Objects are represented at a certain Distance be-711 hind the Lens *, as near Objects are in the compound Microscope. If this Representation be observed thro' an Eye-Glass, it will appear enlarg'd and inverted, as we have faid concerning the Microscope.

It

It is plain alfo, that in this Cafe, as well as in 770 the Microfcope, the Field depends upon the Breadth of the Eye-Glass; as alfo, that the Place of the Eye 771 must be determined in the fame Manner for the Telefcope as for the Microfcope *: For the Aftronomic * 766 Telefcope differs from the compounded Microfcope only in this, that in the Microfcope the Lenfes are more convex, and therefore lefs proper for looking at diffant Objects, cfpecially in refpect of Object-Glaffes. In the Microfcope, the Object-Glafs is more convex than the Eye-Glafs; but the contrary obtains in the Telefcope.

Telescopes cannot be too long for observing the Stars: But, if they are above 20 Foot long, they are of no Use for seeing Objects upon Earth; because of the constant Trembling of the Air, which is too sensible in Glasses that magnify very much.

A fhort Aftronomic Telescope will serve to see Ob-772 jetts upon Earth, by adding to it two convex Lenses, which are also call'd Eye-Glasses; the three Eye-Glasses are alike, and less convex than in the Aftronomic Telescope, the Object-Glassemaining the fame.

Plate XI. Fig. 3.] Take an Object-Glafs VV, 773 which reprefents a diftant Object at e a; then take befides three Eye-Glaffes DD, DD, DD; the first must be for plac'd, that the Rays, coming from the Point of the Reprefentation e a, shall become parallel when they have pass'd the Lens*; * 699 in that Cafe the Rays, which come from the middle Point of the Object-Glafs, will be collected at G; the Second Lens must be for plac'd, that these Rays which are collected at G (where they interfect one another, and move as if they came from that Point) may go out parallel after they have pass'd thro' it *; which being perform'd, * 669 the Rays coming from the Object-Glafs to e, and

and there interfecting and forming that Point of the Reprefentation of the Object, being refracted thro' the first Lens, pass by G parallel to one another; thro' the Second Lens they are refracted * 698 in the Direction D e, and collected at e *, fo as to make it the Point of a new Reprefentation. In the same Manner the Point of a new Reprefentation corresponds to the Point a of the second Representation; which being also true concerning the intermediate Points, there will be found an erect Representation of the Object at a e.

Experiment 3. Plate XI. Fig. 4.] Let three little Boards D, D, D, with Eye-Glaffes in them, that collect parallel Rays at the Diffance of about 5 Inches be moveable upon a Plane between two Rulers, in fuch Manner that the three Glaffes may be in the fame Line as the Hole V, thro' which alone the Light enters into the Room, and in which there is an Object-Glafs, which is fix'd in a fhort Tube, in order to exclude all the fide Light.

This Object Lens is fuch as is able, at the Diffance of 3 Feet from V, to reprefent diffant Objects inverted at F; which Reprefentation will be visible, if you let the Rays fall upon a * 711 white Plane in that Place *. Five Inches farther from F, you must place the first Eye-Glass, and 10 Inches from that, you must place the Second; at f, which is five Inches from it, you will have an erect Representation of the same Objects, which will also be visible, if receiv'd upon a white Plane.

Plate XI. Fig. 3.] If the Reprefentation ac be 774 feen thro' a third Eye-Glafs, fuppofing the Eye at o, in which the parallel Rays a D and a E are collected, the Object appears magnified, brought near, and erect; for it is feen under the Angle Do D,

DoD, when with the naked Eye it wou'd appear under a very fmall Angle; it will also appear near, because, tho it be seen beyond a e, yet the Diflance, at which it appears, has no fensible Relation to the Distance of a very distant Object.

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Experiment 4. Plate XI. Fig. 4.] Supposing every Thing as in the foregoing Experiment; let there be a third Eye-Glass plac'd 10 Inches from the Second, and 5 Inches from that a little Board, or Eye-Stop, with an Hole O; if the Eye be plac'd at O, the Object, as has been faid, will appear erect, magnified, and near. If the Board O be difplac'd, that is, be brought nearer or remov'd farther off, the Field of the Telescope is diminish'd; because there is but one Situation of the Eye, in which all the Rays which pass thro' the Eye-Glasses can come to the Eye.

We are to take notice, that the Eye-Glaffes, made Ufe of here, are not convex enough in refpect of the Object Glafs V; but they are best for making the 3d Experiment.

All the Points of the Object do allo appear more 775 enlighten'd: For the Rays which, coming from any Point of the feveral Points of the Object Glafs, do interfect one another in a Point of the Reprefentation, by reafon of the fmall Diffance of the Eye-Glafs from that Reprefentation, are but little difpers'd before they come to the Eye; fo that they all go into it. And therefore the Illumination, given by the Telefcope, is to that which is given by the naked Eye, as the Surface of the Object Glafs, thro' which the Rays pafs, to the Surface of the Pupil.* *738

One may alfo with two Lenses make Telescopes, 776 which shall shew Objects erect, enlighten'd, and magnified. These must be shorter; for if they be above one Foot long, they become almost useles, because

because their Field will be fo finall; that is, they will take in fo little of an Object.

Plate XI. Fig. 5.] Let V'V be an Object-Glafs, the inverted Reprefentation of a diffiant
^{*711} Object will be at e a*: The Rays are fo intercepted by the concave Lens D D, that fuch of them, as come from the Center of the Lens V V, are inflected as if they proceeded from the Point
^{*705} f*; by the fame Refraction in the Rays which concurr'd at a, become diverging*; having their imaginary Focus at a; which alfo happens in refpect to all the Points of the Reprefentation e a, and, inflead of it, you have an imaginary Reprefentation which is crect at a e; that is, the Rays enter the Eye as if they came from the Object plac'd at ae.

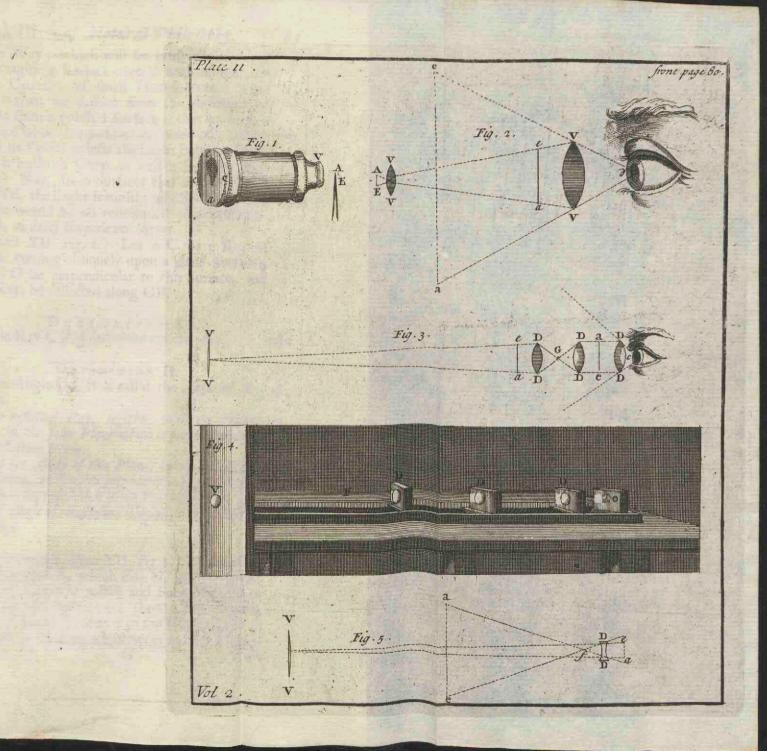
777 The Rays, in all refpects, go out diverging from the Eye-Glass; and therefore the Eye must be brought as near as possible to the Eye-Glass.
778 In this Telescope, the Eicher

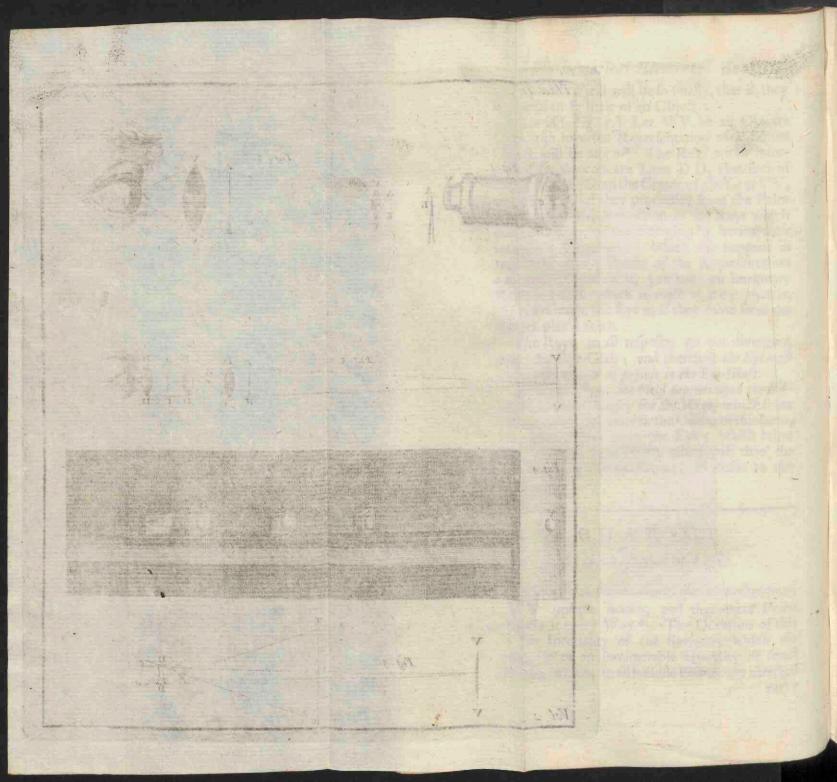
In this Telescope, the Field depends upon the Bigness of the Object Glass; for the Rays, which from a Point come obliquely to the Center of this Lens, very often do not enter the Eye; whilst other Rays from the same Point, which pass thro' the Lens near its Circumference, do come to the Eye.

CHAP. XIII.

Of the Reflexion of Light.

W E have shewn that Light is reflected from opaque Bodies, and that every Point *710 reflects it every Way *. The Occasion of this is the Inequality of the Surfaces, which are made up of an innumerable Quantity of small Planes, which, in all sensible Points, are directed every





every Way; which will be eafily conceiv'd, if we imagine a Surface cover'd with an innumerable Quantity of fmall Hemifpheres. That this is true, we deduce from the Reflexion of Light from a polifh'd Surface; that is, from a Surface whole Inequalities are taken off; which 779 in all its Points reflects the Light only one Way, which holds in Curve as well as in plane Surfaces: Nay, from Surfaces that are not at all Polifh'd, the Light is moftly reflected that Way, and it would be all reflected if they were polifh'd, as daily Experience fhews.

Plate XII. Fig. 2.] Let A C be a Ray of Light coming obliquely upon a plane Surface; let C O be perpendicular to this Surface, and the Rays be reflected along CB.

DEFINITION I.

The Ray CB is call'd the reflected Ray.

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DEFINITION II.

The Angle O CB is call'd the Angle of Re- 781 flexion.

The reflected Ray, together with the incident 782. Ray, in the fame Plane which is perpendicular to the reflecting Plane.

For the Action of this Plane, by which the Light 783 is reflected, is directed perpendicularly to a Plane, which is fuppofed like in all Points.

The Angle of Reflexion is equal to the Angle of 784 Incidence.

Experiment 1. Plate XII. Fig. 1.] Take a plane Looking-glass S, which may be set in any Pofition by Means of a Ball and Socket join'd to the Foot that suffains it; thro' a Hole, in the Plate of Metal L, that is in the Window, let in a Sun Beam of about a Quarter of an Inch Diame-Vol. II. G ter

82

ter into the Room; the Glass must be so dispofed that the Beam may come thro' an oblique cylindric Cavity (of the same Bigness as the Beam) made in a little upright Board T: If you turn this Board Side for Side, the reflected Ray will go thro' the same Cavity. This holds good, whatever the Inclination of the Cavity be, as may be demonstrated by using different Boards.

By which Experiment N°. 779. is prov'd, as well as Nº. 784.

785 Plate XII. Fig. 2.] If the reflected Ray becomes the incident Ray; that is, if the Light comes along the Line B C, it will return in the Line C A, that is, the first that was the incident Ray will become the reflected Ray; as appears from the Equality of the Angles BCO, OCA.

From this Equality of the Angles of Incidence and Reflexion, we farther deduce, that 786 the Light, after it has fallen upon a Body, recedes from it with the fame Force that it came upon it. Let the Motion along A C be refolv'd into two r192 Motions along A O and OC *, fuppofing A O parallel to the reflecting Plane, and OC perpen-

dicular to it. Let AO be continued; the Motion in that Direction is not alter'd from the Action of the Plane: Therefore let A O and OB be equal; if the Light recedes from the Plane with the fame Force, with which it came upon it, the Motion occasion'd by the Repulfion is represented by CO, and in that Case "190 the reflected Ray goes thro' B*; that is, the Angle OCB is equal to the Angle OCA, which agrees with the Experiment.

787 As to the Reflexion of Light, it is to be obferv'd, that Light does not run against the solid Part of Bodies, when it is reflected by them; but that it is reflected in those Places, where it cou'd very freely pa/s.

pafs. I shall prove this by feveral Experiments, by which many other wonderful Properties of Reflexion are discover'd.

It is a common Experiment observed by every 788 Body, that when Light is mov'd thro' any Medium, as for Example, Glafs, Water or Air, it does not undergo a fensible and regular Reflexion; but that it is reflected there, where two Mediums of different Density are separated; so it is reflected in the Surface of Water or Glafs.

Cou'd Light in fuch Quantity firike against the Particles, just where the Mediums are separated, whereas it moves thro' both the Mediums for a great Space without firiking against any fuch Particles; Are there more of those Particles near the Surface than elsewhere? Light also is more 789 abundantly reflected in a denser Medium, when it comes against the Surface of a rarer; than when on the contrary, moving in a rarer Medium, it firikes against the Surface of a denser.

Experiment 2. Plate XII. Fig. 3.] In a dark Place in which the Light enters thro' an Hole in the Plane L, let there be placed a triangular Prifm of Glafs A B; let the Light enter the Prifm thro' one Side; if it comes to the next Side making an Angle of Incidence greater than 40 Degrees, it is wholly reflected, and does not at all penetrate into the Air; but Light moving in Air is never wholly reflected by the Glafs.

But if the Reflexion be made by the firiking of Light against the folid Parts of Bodies, there must be more such Parts in Air than in Glass; for if Light was reflected from the Glass itself into the Air, the Light would never come to the Separation of the Mediums; that the Light can allo go out of Glass in the very Places; where it is reflected, is prov'd by following Ex-G 2 periments.

periments. Therefore in the Neighbourhood of the Glais there must be fo many Parts in the Air, that there may be no Way for the Passage of the Light, to cause it to be wholly reflected into the Glais; but it is plain that there are no such Parts, because Light comes thro' the Air in all Directions quite to the Glais. Even in the fame Place of the Surface which separates the Glais and Air, the Light which comes from one Side is reflected, whils that which comes from the other Side is transmitted. Which clearly proves that the Light is reflected in the very Place where it can go thro'.

790 Experiment 3. Plate XII. Fig. 3.] Every Thing
 being as in the former Experiment, if the Obliquity of the Light be chang'd, Part of it will pais thro' into the Air.

Who wou'd conceive that Light, which paffes from Glass into Air, and does not run against the folid Parts, shou'd all of it (by a little increasing the Obliquity) run against those Parts; when in each Medium, as has been already faid, there are Pores and Passages in all Directions?

791

Experiment 4. Plate XII. Fig. 3.] Take a Glass triangular Prifm AB, moveable about its Axis; which is made fo by fixing brass Plates to its Ends, with brass Wires perpendicularly fix'd to them: The Prifm must be fo laid upon the Trough P, as that the faid Wires may bear upon the Brims of it, which are made a little hollow to receive them, yet fo as to let them turn, that the Prifm may move freely about its Axis: Let it be fo plac'd as to reflect the Light in the fame Manner, as in the fecond Experiment. Let the Trough be fill'd with Water up to the Prifm; then the Light, which, ftriking againft the

the Air, was wholly reflected, now running againft the Water, does partly enter into it, and is only reflected in part.

Which Experiment does not at all agree with a Reflexion made by a Stroke upon the folid Parts.

In the third Part of this Book we fhall alfo fhew, that thin Plates, which reflect Light, will transfmit it, if they become thicker*. *894

The fourth Experiment also proves that the 792 reflecting Power is somuch greater, as the Mediums, which are separated by a reflecting Surface, differ more in Density; for Glass and Air differ more in Density than Glass and Water.

In this Experiment we also fee that Reflection 793 is made by the fame Power by which the Rays are refracted, which produces different Effects in different Circumftances.

A Ray, which goes out of a denfer into a rarer Medium, by the Attraction of the former, 625 is made to recede from the Perpendicular *; if the Obliquity of the incident Ray be increas'd, the Obliquity of the refracted Ray will also increafe, till it comes at laft to move in the very Surface which parts the Mediums. And this obtains, when the Sine of the Angle of Incidence is to the whole Sine, as the Sine of Incidence, in the denfer Medium, is to the Sine of Refraction in the rarer; for in that Cafe the Angle of Refraction is a right one. If the Obliquity of the incident Ray be more increas'd, it is plain that the Ray cannot penetrate into the rarer Medium : This is the Cafe in which the Light is wholly reflected; which Reflexion depends upon that Attraction by which the Rays are refracted. For when the Ray is mov'd thro' the Space of Attraction, it is bent towards the denfer Medium *; if it be in the denfer Medi- * 618 G 3 um,

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um, and fo bent, that, before it has gone thro' the whole Space of Attraction, the Tangent to the Curve be parallel to the Surface that feparates the Mediums, the Curve, being continued, turns back again; and therefore the Ray is reflected by the Attraction of the denfer Medium, and this Continuation of the Curve is fimilar and equal to the first Part, and makes the Angle of Reflexion equal to the Angle of Incidence; because the Light returns thro' the same Part of the Sphere of Attraction; and the fame attracting Force acts upon the Light in correspondent Points of the two Parts of the Curve. Thus a Projectile, in its Afcent and Defcent, defcribes fimilar Curves.

Book III.

Yet that all Reflexion does not depend upon that 794 Attraction, in the fame Manner, is evident; for in that Cafe, in which the Refraction is made, Part of the Light is reflected; for the Light does not wholly penetrate out of the rarer into a denfer Medium ; for even in that Cafe, in which the Attraction is the most oppos'd to the Reflexion that is poffible, yet fome Rays are reflected.

Yet it cannot be doubted but that, in every 795 Cale, Reflexion has Relation to the refracting Power. Where Light paffes without Refraction, there it is 796 617, not reflected *; but where the Refraction is greateft, there the Reflexion is alfo frongeft *; which is true, * 631. not only when Light, moving in a denfer Medium, flrikes against a rarer, as in the fourth Experiment; but the fame Thing is observ'd, when Light firikes against the denser Medium : Thus fuppoling the Light to move in Air, the Surface of Glafs reflects more ftrongly than that of Water; and that of a Diamond yet more ftrongly. If Glafs and a Diamond be immers'd in Water; the refracting Power is lefs in the Separation of those Bodies with Water, than where thole

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those Bodies touch the Air*. These Bodies also * 631 reflect Light less ftrongly in Water than in Air. From this Relation of the reflecting and refracting Powers, we deduce, that Light is driven 797 back at a certain Distance from the Bodies, in the fame Manner that the refracting Power does alfo act at some Distance from the Body: This Proposition is confirm'd from what has been demonftrated concerning Reflexion, which does not depend upon a Stroke made against the folid Parts of Bodies; and this is fully made out, if we confider, that polifb'd Bodies reflect the Light 798 regularly (which we observe in Looking-Glaffes) the there be a great many Scratches in their Surfaces : For as they are polifh'd with the Powder of Emery and Putty, tho their Parts are very finall, yet they leave very great Scratches on the Surface in Respect of the Particles of Light; whence in the Surface itfelf the Reflexion must needs be irregular; but if we conceive the Reflexion to be made at fome Diftance from the Surface, the Irregularities are diminish'd, and almost wholly taken off, as is eafily underftood by any one that confiders it with Attention. .

CHAP. XIV.

Of Plane Mirrors.

Plate XII. Fig. 4.] LET b c be the Surface of a Plane Mirror or 799 Looking-Glafs; A a radiant Point. Let the Plane of the Glafs be continued, and from the Radiant A let a Perpendicular AC fall upon it; if this Line be continued, and C a be made equal to CA, a will be the imaginary Focus of the reflected Rays that proceed from A. Let A b be the incident Ray; bf the reflected Ray; which continue beyond the Glafs; G 4 because

* 784

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the Angles of Incidence and Reflexion are equal to one another *, their Complements also, which are the Angles AbC, fbd are equal; to this is equal its oppofite and vertical a bC: The Triangles A bC, abC, which are rectangular, have the Side Cb common, and the Angles Cba, CbA, equal; therefore they agree in all Respects, and C A and Ca are equal to one another: Which Demonstration may be applied to all other Rays which flow from A, in whatever Plane, perpendicular to the Glass, they be conceiv'd to be. Therefore wherever the Eye is, if the reflected Rays come to it, they will enter the Eye as if they came from a; and the Point A will appear in that Place *; but the Appearance of that Point will have the fame Position in Respect of the Mirror. 800 behind it, as a radiant Point has before it.

If this be applied to all the Points of the Ob-801 ject, it will appear, that the Object will appear behind the Glass in the same Position that it has before the Glass.

CHAP. XV.

Of Spherical Mirrors.

- VERY fpherical Surface may be confi-802 H, der'd; as made of innumerable small Planes; and a Plane, which touches the Sphere in any Point, is as it were a Continuation of fuch a fmall Plane.
- 803 Spherical Mirrors are either Concave or Convex.

The first are made of Part of an hollow Sphere poliih'd.

The fecond are Parts of Spheres polifh'd on the Outfide.

A

A Ray coming upon any (pherical Mirror, together 804. with its reflected Ray, is in a Plane, which, being continued, goes thro' the Center of the Sphere *, for * 782 fuch a Plane is perpendicular to the Surface of the Sphere. A Line which is drawn thro' the 804. Center of the Sphere and Point of Incidence, being continued makes equal Angles with the incident and reflected Rays *; for that Line is perpendicular to * 784 the Surface, and those Angles are Angles of Incidence and Reflexion: Therefore the Ray that goes 806 thro' the Center, or which, being continued, wou'd go thro' the Center, when reflected returns upon it self.

Plate XII. Fig. 5.] Let bc be a Portion of the convex Mirror; A a radiant Point; let A b, A d, A c, be incident Rays ; the reflected Rays will be bf, dg, cb; if from the radiant Point A, a 807 Tangent be drawn to the Mirror, the reflected Ray will be a Continuation of the incident, or rather the Reflexion of the Rays terminated in the Point of Contact.

If b f, dg, cb, the Rays that are reflected from 808 the convex Mirror, be continued, with all the intermediate ones, by their Interfections they will form the Curve a a, which all these Raystouch, and the neighbouring Rays interfect in the Periphery of the Curve ; fo that they always enter the Eye as if they came from a Point of the Periphery; in which therefore the Point A does always appear *, as long as the reflected Rays can come to * 737 the Eye, and the Eye is mov'd in a Plane which goes thro' the Center of the Sphere ; but when the Eye is mov'd out of that Plane the radiant appears in another Curve, because there are fuch Curves in every Plane, which may be conceiv'd to pais thro' A and C.

Since all these Curves, and each of them 800 wholly are behind the Glass, all the Objects also appear behind the Surface of the Glafs.

The

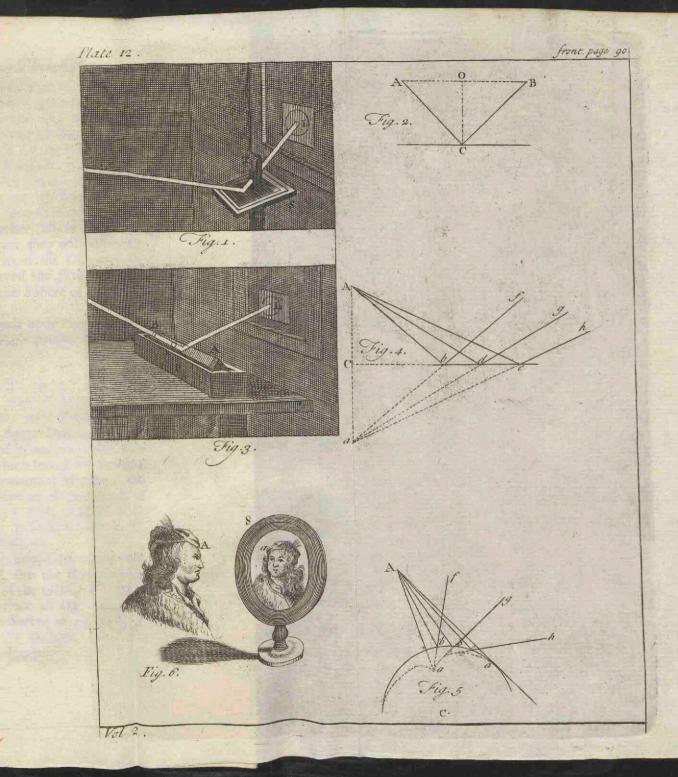
810 The Objects likewife appear erect. For if the Point A be mov'd about the Mirror, the whole Curve a a is carried with the fame Motion; which proves (in Refpect to the erect inverse Situation) that the Points of the Representation have the fame Relation to each other, as the Points of the Object itfelf.

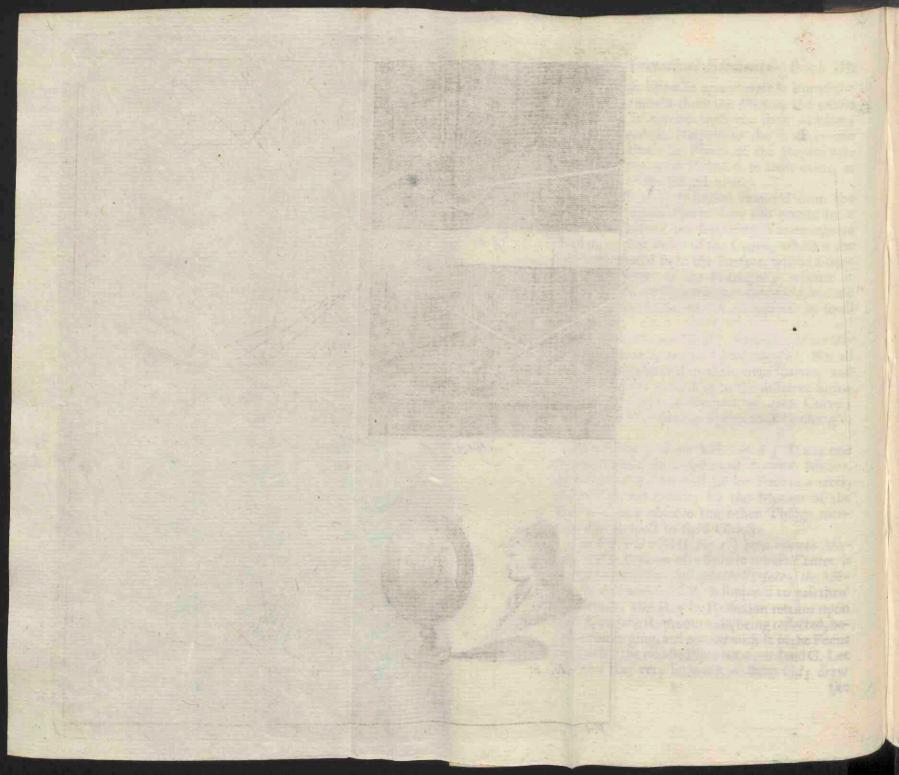
As the Point A is farther remov'd from the Glafs, the whole Curve does alfo recede by a contrary Motion; but fuppoling A at an infinite Diftance, that Point of the Curve, which is the fartheft remov'd from the Surface, will be diftant from a Quarter of the Diameter; whence it

- 811 follows, that the Objects appear diminish'd, because all the Representations are comprised in small Limits.
- 812 If the Eye be mov'd, the Appearance of the Object is also mov'd, and its Figure chang'd: For all the Points are mov'd in their own Curves, and that unequally, according to the different Situation of the Eye in Respect of each Curve; whence of Necessity the Figure must be chang'd.

Experiment 1! Plate XII. Fig. 6.] If any one fees his Face in a fpherical convex Mirror, ftanding at A, he will fee his Face at a creft, diminifh'd and unlike; by the Motion of the Eye one may observe the other Things mention'd in Respect to fix'd Objects.

813 Let b d (Plate XIII. Fig. 1.] be a concave Mirror, and a Portion of a Sphere whole Center is C; let parallel Rays fall upon the Surface of the Mirror, one of which, Cd, is fuppos'd to pass thro' the Center; this Ray by Reflexion returns upon
* 806 itfelf*, and the Rays next to it, being reflected, become converging, and concur with it in the Focus Fy which is the middle Point between d and C. Let A b be a Ray very little diftant from Cd; draw
2 the





the Semidiameter C b; the Angle of Incidence will be A bC, to which the Angle of Reflexion C 6 F* is equal, as also the alternate Angle * 805 b CF; therefore b FC is an Ifosceles Triangle, and the Sides F C and F b are equal; becaufe b d is very fmall, F d and F b do not fenfibly differ; therefore F C and F d are equal; which Demonstration will ferve for all the Rays that are but very little diftant from C d.

If parallel Rays are farther diftant from C d, they do not meet at F; yet they will all come together into a little Circle, if the Diameter of the Mirror does not exceed the fifth or fixth Part of the Diameter of the Sphere of which it is a Portion.

Burning Mirrors are made upon this Founda-814 tion, which collect the Sun's parallel Rays into a Focus.

Experiment 2. Plate XIII. Fig. 2.] Let there , be a concave Mirror S, made of Metal, or of Glass quickfilver'd behind ; let it stand upon the wooden Foot P, whofe upper Part is bor'd fo as to receive a Cylinder of Wood made fast to a transverse Piece A A, which serves for turning the Glass round with a horizontal Motion; and the Mirror itfelf must move upon two Ends of an Axis between the Pillars A B, A B, fo as to be inclin'd in any Angle, and the Screws B, B, will make it fast in any Polition.

Having expos'd the Burning-Glass to the Sun-Beams in fuch a Manner, that the Ray, which comes upon the Middle of the Glafs, is perpendicular to its Surface; fince all the others are Parallel to it, they are collected in a Focus, at a Quarter of a Diameter's Diffance from the Glais, and there burn violently.

If the Diameter of the Surface of the Mirror (as it is in mine) be of about 15 Inches, and the Focus is 18 Inches diffant from it, Wood will immediately be in a Flame, and thin Plates of Lead prefently melt.

If we confider the Rays that are at fome Diflance from Cd, and parallel to it, those of them that are nearest one another being reflected, will intersect before they come to Cd, and

- 815 in that Cafe, that is, where parallel incident Rays fall obliquely on the Glass, being a little dispers'd by Reflexion, they are collected in a Point.
- 816 If the Focus, in which parallel Rays are collected by a concave Mirror, becomes the radiant Point, the Rays, which are but little difperfed, are refle-
- * 813 Eted parallel to one another *.
- From these Properties of a concave Mirror
 we deduce the Method of representing Objects in a dark Place, much like what was before shewn in Respect of a Convex Lens*.
- 711 Plate XIII. Fig. 3.] Let there be an Hole F thro' the Wall; let a b be a concave Mirror, fo plac'd as to collect the parallel Rays that are perpendicular to the Wall at F: The Rays, com-
- ⁸¹⁶ ing from F in that Direction, are reflected *, and fuch are the Rays, which, being reflected from the external Objects, interfect one another at F.

Let A F be Rays coming from a Point of a diftant Object; these Rays are by the Mirror reflected perpendicularly to the Wall; and because Rays coming from a distant Point, and passing thro' a small Hole, may be looked upon as parallel, these Rays will, after Reflexion, be collected into one Point at *a*, at the Distance of 815 the Wall*, that is in its Surface; where therefore the Point will be represented. In the same Manner the Rays which come from a Point thro'

thro' B F, are collected at b; which, as it is true with Regard to all the Points of an Object, will give the Reprefentation of it upon the Wall; and if the Wall be white, and the Object enlighten'd by the Rays of the Sun, the Picture will appear in very lively Colours.

Experiment 3. Plate XIII. Fig. 4.7 In a dark Place cover the Backfide of the Window-fhut with white Paper, an Hole being made in the Middle of it little more than half an Inch Diameter, fo as to answer to an Hole behind it in the Window-fhut, over-against which, at a Distance no lefs than of 50 Foot, there are feveral Objects enlighten'd by the Sun : Let a concave Mirror whofe Surface is 15 Inches wide, and which collects parallel Rays at the Diftance of 18 Inches, be placed at that Diftance from the Window, in fuch Manner, that a Line paffing thro' the Center of the Hole, and the Center of the Surface of the Mirror, be perpendicular to the Plane of the Paper and the Surface of the Glafs. Then the Objects will be reprefented upon the Paper in a Circle concentric with the Hole, and whole Diameter is equal to the Diameter of the Mirror. You must join to the Hole, on the Outfide of the Room, an hollow, truncated Cone, to exclude the Light which does not come from the Objects to be represented.

Let be (Plate XIII. Fig. 5.) be a concave Mir-818 ror; C the Center of its Concavity; A a radiant Point farther diftant than C the Center of the Glafs; Ab, Ad, Ae incident Rays, whole refle-Eled ones b f, d g, e h, with their intermediate ones by their mutual Interfections form the Curve a a, which they touch; therefore the Point A appears in that Curve*; and if the Eye be mov'd in the * 808 Plane of the Curve, the Appearance will change Place

Place in that Curve. But in all the Planes which may be conceived to pass thro' C A, there is fuch a Curve, and they all concur in the Line

- 819 C A, namely, at the Point a. Therefore in that Point a, the reflected Rays are the most abundantly collected, which therefore is call'd the Focus of the Rays coming from A. On the contrary, A
- * 785 will be the Focus, fuppofing the Radiant at a*. In this Figure there is only Part of the Curve drawn, which is produced by one Part of the Line AC; fuch another Part must be conceiv'd on the other Side, and both join in the Focus of the radiant Point.
- 820 As the radiant Point recedes, the Curve comes nearer the Mirror.
- 821 As the Radiant comes forward, the Curve recedes from the Mirror, and moves towards the Radiant, till they concur at the Center C; in which, if the Radiant be plac'd, all the reflected
- * 806 Rays will coincide with the incident *, and the whole Curve will, as it were, be fwallow'd up in the Center.
 - 822 If the Radiant is yet brought nearer to the Mirror, fo as to be between the Center and the Gla/s, the Curve will recede farther, and be then beyond the Center; and in the Curve, that Point will recede most of all, in which all the Curves concur which are conceiv'd in feveral Planes, that is, the Focus of the radiant Point : And that Focus
- 823 will be at an infinite Diftance when the Radiant is diftant from the Mirror just the fourth Part of the
- * 813 Diameter of the Sphere.* Then also the Curve is extended in infinitum, and the two Parts which concur in the Focus of the Radiant are feparated; this feparated Part is feen at a a : (Plate XIII. Fig. 6.) If the Radiant be brought fill nearer, the 824 Curve Parts decline from one another, because the Rays, such as A b, and those nearer it, being re-

fleeted,

flected, do not touch the Curve, but become divergent; that is, those reflected Rays, being continued beyond the Glafs, will interfect one another, and form a new Curve behind the Glass, which has two Legs, one of which is feen at aa; they concur in the Line CA continued, namely, at a, and, receding from the Glass, are ftretched out in infinitum. And there is also, on each Side of the radiant Point, a Point in the Surface as d, which feparates the Rays that form the Curves a a and aa; and the Ray A d, being reflected in d g, touches neither of the Curves, if it be infinitely continued toward each Part g, g, tho it is continually coming nearer to each Curve. If the whole Sphere was compleated, in Respect of the opposite Part of the Sphere, the Radiant would be beyond the Center, and the reflected Rays would form the Curve which we have mention'd * 818 before, * by which the feparated Legs, as a a, would be join'd. These Things thus premifed, we proceed to explain the Phænomena of concave Mirrors.

If the Mirror be enlightened by a lucid Body, the 825 Rays which come from all the Points of the Object, being reflected, will form Curves, but are chiefly collected in the Faci of these Points; * therefore if * 819 these Foci are in the Surface of a white Plane, there 826 will be upon it a Representation of the lucid Body, as In the fecond Experiment of Chap. IX. And that Representation will be inverted ; for the Line, which joins the radiant Point with its Focus, goes thro' the Center of the Sphere; * in which there- * \$19 fore all fuch Lines interfect one another; and this Interfection is between the radiant Point and the Focus, * in which the Point is represented. * 818

But as the lucid Body is brought nearer to the Mirror, the Appearance recedes farther, * and in * 820 this Cafe becomes bigger.

Experi-

Experiment 4.] Hold a lighted Candle between the Mirror and the Center of the Sphere of which it is a Portion; yet fo that it may be more diftant from the Mirror than that Center: If then there be a white Plane perpendicular to the Line that paffes thro' the Candle and the Center of the Mirror, and this Plane be held beyond the Center, you will have upon it an inverted Reprefentation of the Candle; the proper Place will be found by moving the Plane forwards and backwards; as likewife by this Proportion, viz. As the Difference of the Diftances of the Candle, from the Center of the Sphere and from the Mirror, is to the fourth Part of the Diameter of the Sphere, fo is the Diftance of the Candle from the Glass, to the Diftance requir'd. As the Candle is brought nearer to the Speculum, the Plane must be mov'd farther off, and the Reprefentation will grow bigger.

827 Plate XIII. Fig. 5.] Objetts, placed beyond the Center, appear between the Glass and the Center, for
* 818 all the Points appear in a Curve as at a a; * the
828 Images of the objetts will also be diminish d and inverted: for they are reduced into a narrow Space; and as the Point A moves downwards, its Representation will be carried upwards; for the Line a a keeps the fame Situation in respect of A C a, as it is carried round the Center C.

Experiment 5. Plate XIII. Fig. 7.] Left the Representation of the Object shou'd be less vivid, the Mirror is to be included in a Box. If you have a Mirror, whose Surface is about 8 Inches wide, and which is a Portion of a Sphere of one Foot and a half in Diameter; shut it up in a Box P, in whose Fore-Part there is an Hole of

of about 6 Inches Diameter, and from which the Glafs is diftant above 6 or 7 Inches; and let this Opening be turn'd from the Light. Now if any Perfon, as A, beholds himfelf at the Diftance of about 2 Foot from the Glafs; his Face will appear inverted in the Box towards the Hole; and if the Beholder comes nearer, he will fee a Head coming out of the Hole.

The Representation of a Point, placed in the Cen-829 ter of a Sphere, coincides with the radiant Point itfelf, and is as it were fivallowed up by it *. * \$21

If the Eye be plac'd in that Genter, no Object can 830 be feen by it; for then only the Rays, that flow from the Eye, will be reflected to it*. *806

If the Object be between the Center and the Point, 831 in which parallel Rays are collected after Reflexion, the Object will also appear without the Glass, at a greater Distance from the Glass than the Object itfelf*; the Representation is inverted, which is pro-*822 ved in the fame Manner as in N°. 828; and magnified, because it is farther remov'd from the Center, than the Object itself is distant from it; for the Representation recedes from the Center in infinitum; whill the Object goes thro' the fourth Part of the Diameter of the Sphere.

If the Object be lefs diftant from the Mirror than 832, the fourth Part of the Diameter of the Sphere, according to the different Situation of the Eye, the Object appears either before or behind the Glass.

Plate XIII. Fig. 6.] If the Eye be io plac'd, that those Rays may come to it which form the Curve a a, as towards f, it will fee the Appearance of the Objects towards the Glass * magni- * 824. fied; because the Curves, as a, which belong to feveral Points, are diverging.

If those Rays come to the Eye which form 833 the Curve *a a*, the Object will appear without the Glass: And in both Cafes the Representation is Vol. II. H erect 3

 $ere \mathcal{E}t$; for as the Point A afcends or defcends, the Curves *a a* and *a a*, in which it is reprefented, are agitated with the fame Motion.

834 If the Eye be in the Point, in which the reflected Rays, that belong to each Curve, interfect one another, as in O, the Appearance of the Object will be double.

It is plain in every Cafe of the Appearance, that the Points have not the fame Relation to one another, as the Points of the Object; and

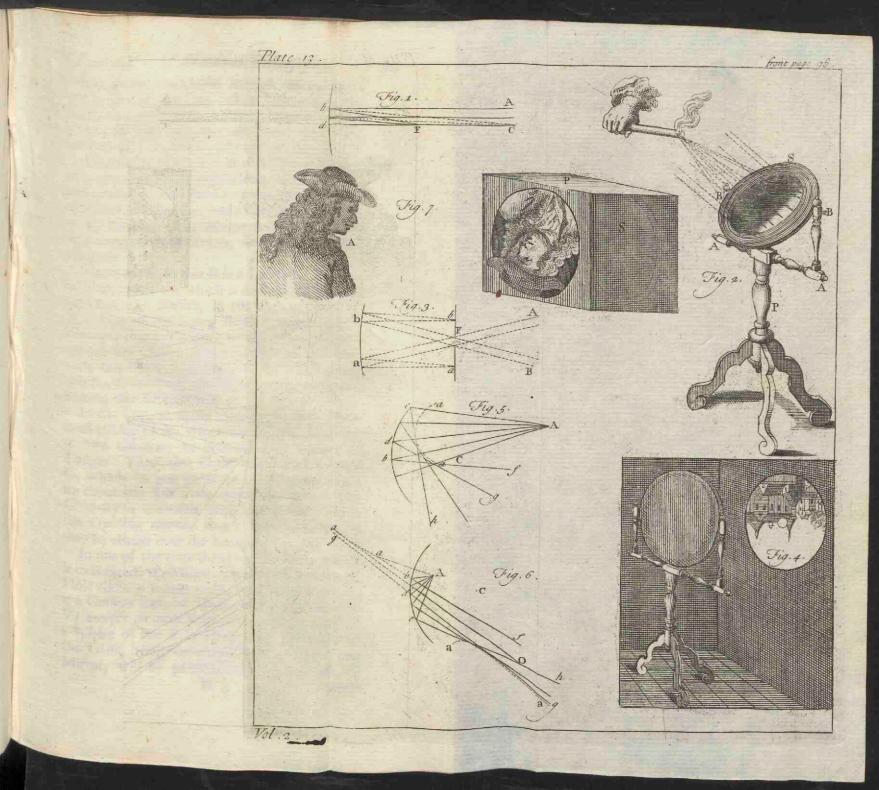
- 835 therefore that a Concave Mirror never reprefents the Objects exactly : But the most irregular Reprefentation of all is that which is in fuch Lines, as a a.
- 836 There are also cylindric Mirrors convex and concave; these in one Respect are plane, and in another Respect spherical; and therefore the Representation of the Objects is very irregular; which Irregularity, fince it depends upon a regular Figure, may be so determin'd, that Pictures may be drawn, which, tho truly irregular, will in such a Glass appear regular by Resterction, in a determinate Situation of the Eye.

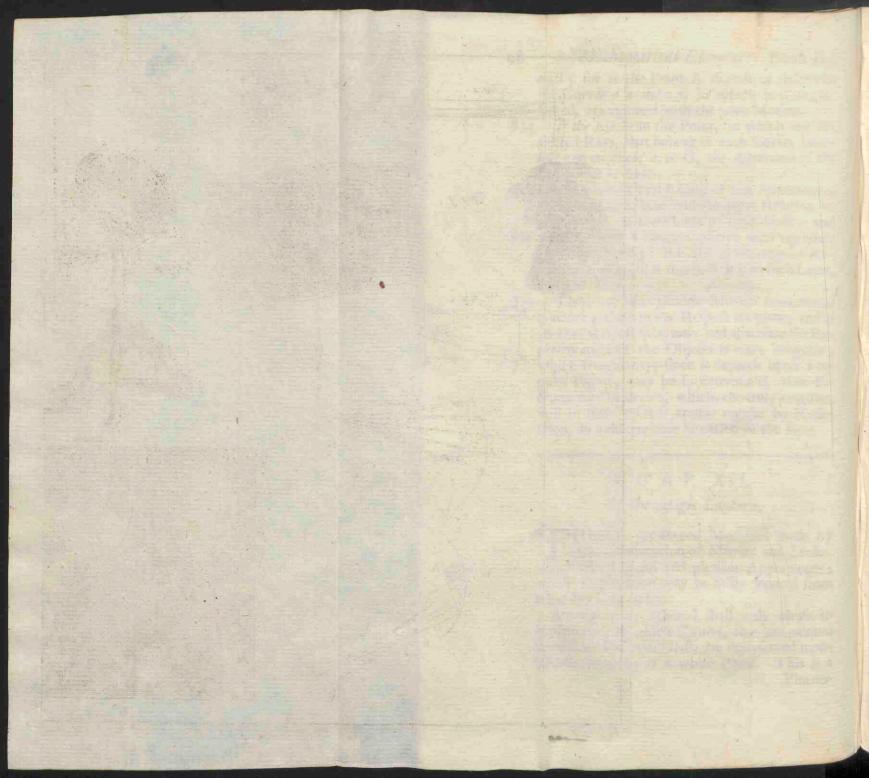
CHAP. XVI.

Of the Magic Lantern.

THERE are feveral Machines made by the Combination of Mirrors and Lenfes, which afford ufeful and pleafant Appearances; whofe Explanation may be eafily deduc'd from what has been faid.

Among many other I shall only chufe to explain one, in which Figures, that are painted upon small Pieces of Glass, are represented monstroufly large upon a white Plane. This is a Phæno-





Phænomenon wonderful enough to deferve a particular Explanation. The Inftrument that performs this is call'd a *Magic Lantern*, which Optic Writers have not altogether pafs'd by, but yet have not fufficiently explain'd.

Plate XIV. Fig. 1.] Let there be a wooden 837 Box about a Foot and an halflong, 14 Inches high, and as wide; there must be a concave Mirror S, of 8 Inches Diameter, and a Portion of a Sphere of 18 Inches: This Mirror is fix'd to a Foot which moves upon Rulers, along the Length of the Box.

There is also in this Box a Lamp L, fustain'd by a wooden Foot which is moveable long-wife between two Rulers, in the Side of the Box. The Pipe of the Lamp stands forward in such Manner, that the Center of the Flame is overagainst the Center of the Surface of the Mirror; this Flame is made up of four little Flames; which, by touching one another, make one square Flame, two Inches wide.

In the Top, or upper Plane of the Box, there is an oblong Hole, which has a Cover that flides to two Grooves, or between two Rulers or Ledges: Thro' this Cover paffes the Chimney C, which (as you fee in the third Figure) flands up about one half Foot above the Box. The Chimney is moveable with the Cover, whilft the Opening remains flut; that the Chimney may be always over the Lamp.

In one of the little Sides of the Box, which is overagainft the Mirror, there must be a round Hole about ς Inches wide; which must have in it a Convex Lens of Glais of the fame Bigness V, convex on both Sides, which are Portions of a Sphere of one Foot Diameter : The Axis of this Glais, being continued to the Surface of the Mirror, will be perpendicular to it, and fall H 2 upon

upon its Center, as likewife to the Plane of the Flame, thro' whole middle Point it also passes.

TOO

This Hole is flut and open'd by a Plane moveable in a Groove, which is mov'd by a Cylinder that flands out of the Box at E.

To this Hole without the Box answers the Tube T, whose Length and Diameter is of about 6 Inches, at the End of which there is a Ring, in which the second Tube t moves, of about 4 Inches Diameter, and 5 or 6 Inches long.

In the leffer Tube there are two Lenfes; the first in that End which is thrust into the Tube T, and it is of the fame Convexity as the Glass V, and three Inches and a half Diameter. The fecond Lens is three Inches from the first, and flatter, being terminated by Portions of a Sphere of four Foot Diameter. Between these Lenses, at the Distance of about one Inch from the second, there is plac'd a wooden circular Stop, or Aperture, which shuts up all the Tube, except an Hole of an Inch and a Quarter Diameter in the Middle of the Wood.

The Objects, that are to be reprefented, are to be painted upon a flat thin Piece of Glafs, which muft be mov'd without the Box overagainst the Glafs V, between it and the Tube T, the Picture being in an inverted Position. If these Pictures are round, they may be of ς Inches Diameter : That they may be moved easily, they are put into flat Boards, three in one Board. The Picture also may be painted upon long Glaffes, which may be fucceffively made to flide before the Glafs V.

Plate IV. Fig. 3.] This whole Box flands upon a Frame or Foot, made fo that it may be fix'd at different Heights. There are flat Pieces of Wood fix'd to the Box, at Bottom, which flide in Grooves in the Frame; each of them has a Slit

Slit in it : fo that the Box may be made fast at any defir'd Height, by the Help of Screws Join'd to the Frame, and moveable in the Slits.

The whole Machine is placed at the Diffance of 15, 20, or 30 Feet from a white Plane, which Diftance must be different according to the Bigness of the Plane ; for this Diftance may be equal to the Length of the Plane: The Box must be just at such an Height that the Tubes, in the Side of the Box, may be exactly opposite to the Middle of the Plane.

The Lamp being lighted, the Box must be fhut, and the Figures which are painted upon the Glass will be represented upon a white Plane. By moving backwards and forwards the Tube that has the two Lenfes in it, you will find the proper Polition of the Glaffes requir'd to give a diftinct Representation. As for the Disposition of the feveral Parts of the Machine, which immediately ferve for exhibiting this Appearance, we shall here more particularly explain.

Plate XIV. Fig. 2.] The Parts in this Figure 828 are fhewn feparately; SS is the Mirror, 11 the Flame which confifts of four Flames in the Line 11; V V is the Glafs V of the first Figure; O O is a Picture painted upon a flat thin Piece of Glass; aa the biggest Lens; dd the flattest Lens; bb the wooden Stop between the Lenfes; f the Aperture or Hole in the Middle of the wooden Circle.

These Things being dispos'd as has been already explain'd, and as may be feen in the Figure, the Rays which proceed from a Point of the Picture O O, by going thro' the Lens a a, become lefs diverging, and fall upon the Lens d d, as if they came from a Point more remote *; * 669 from this Lens they go out more converging*, * 699 and are collected upon the Surface of the white Plane,

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Plane, where they exhibit the Point of the Pi-* 711 cture that is painted on the Glafs *. This Picture is illuminated both by the Rays that proceed from the Flame 11, and by the Rays reflected by the Mirror S S.

For the Perfection of this Machine it is requir'd, 1/t, That the Figure OO be enlighten'd as much as poffible; 2dly, That it be equally enlighten'd in all its Points; 3dly, That all the Light, by which every Point of the Picture is enlighten'd, go thro' the Lenfes a a and d d to the white Plane, and ferve to make the Reprefentation; 4thly, That no other Light but that go out of the Box, left the Reprefentation fhou'd be lefs lively, by reafon of extraneous Light.

The first Requisite depends upon the Bigness of the Flame and of the Mirror, and of its Concavity; the more concave it is, the nearer it is to be brought to the Flame, and then the more Rays will be intercepted and reflected; yet Care must be taken that the Mirror (which may be made of very good Glass) be not too much heated.

When the Flame and Mirror are fo contriv'd, that the Picture is the most enlighten'd that it can possibly be, and every where equally enlighten'd, the Flame and Mirror must be fo plac'd, that the inverted Image of the Flame shall fall just upon the Picture*. Now, as the Representation of the Flame can be increas'd and diminish'd *, the Mirror and Flame must be fo dispos'd, that the Representation of the Flame shall cover the whole Picture upon the Glass, but so as not to exceed it : For then the Picture is as much enlighten'd by the reflected Light as it can be, and all its Points are equally illuminated ; the direct Light also does pretty near

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near equally fall upon all the Points of the Picture; this Light wou'd indeed be increas'd by bringing the Flame nearer; but the reflected Light wou'd be diminifhed, and the Diminution of this laft wou'd be greater than the Increase of the other.

The Glass V V ferves to inflect the Light, by which the Picture O O is illuminated, before it comes to the Picture; by which Inflection all the Light comes to the Lens *a a*, and ferves for the Representation on the white Plane.

All the Light, that is of Use for this Reprelentation, goes thro' the Hole f; and the Rays, coming from different Points, interfect one another there; wherefore the Picture upon the Glafs, which is placed inverted, is reprefented crect upon the white Plane; by the Ring bb all the Rays, which do not ferve to form the Reprefentation, are intercepted, left they fhou'd enter the Room, and make the Picture lefs diffinct. This Ring or Aperture also intercepts those Rays by which one Point is more enlighten'd than another, whereby the Light, which (from what has been faid) is equally enough diffus'd, is yet made more equal. But unless the Stop or Aperture b b be just where the Rays interfect, it does a great deal of Mischief.



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BOOK



BOOK III.

PART III.

Of Opacity and Colours,

C H A P. XVII. Of the Opacity of Bodies.

DEFINITION.

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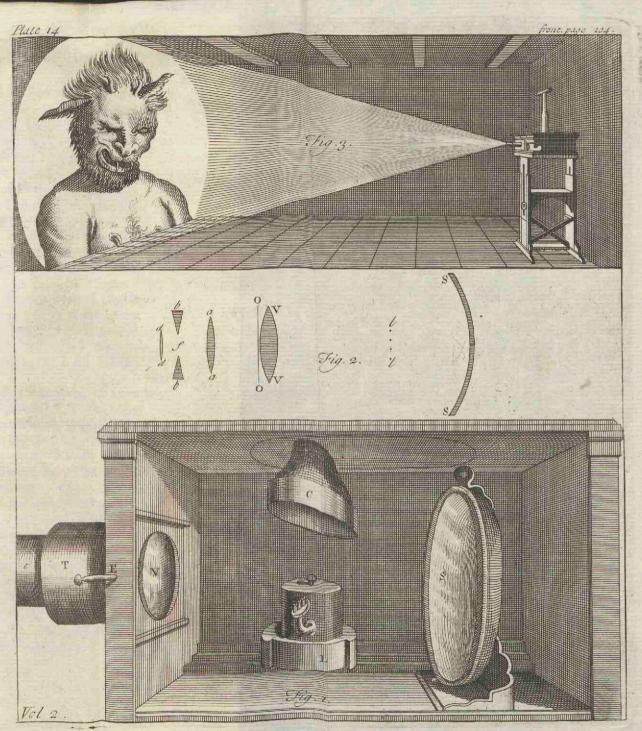


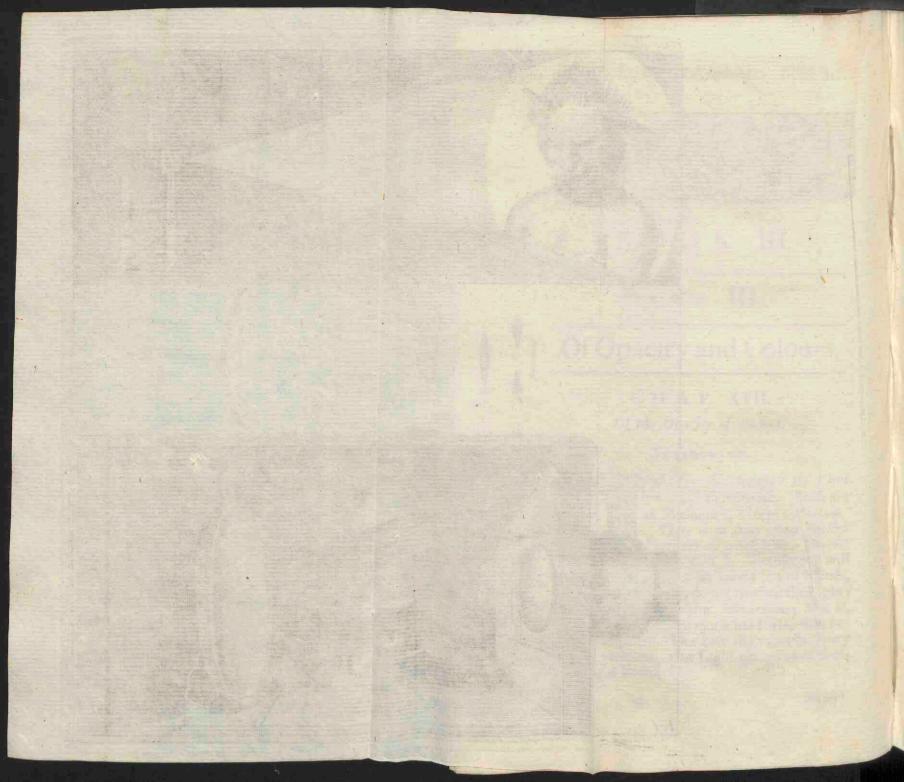
ODIES that transmit the Light are call'd Transparent. Such are all Mediums*, except a Vacuum. There is no Body whose smallest Parts are not transparent: No one who is used to Microscopes will

doubt of this. There are fome Parts of Metals, which, tho very finall, do not transmit the Light; But if they be diffolv'd in *Menstruums*; that is, if they be divided into much less Parts, they become transparent. One may also prove by a very casy Experiment, that Light can go thro' feveral opaque Bodies.

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Experis





Experiment 1.] In a dark Room, in which the Sun's Light comes in thro' an Hole, let that Hole be cover'd with a thin Plate of an opaque Body, and the Light will go thro' it; a Piece of Wood, of the Thicknels of the tenth Part of an Inch, does not intercept all the Light. But the perfect Transparency of Parts in opaque Bodies is not prov'd by this Experiment, for that Transparency is only in very fmall Parts.

Opacity does not (as is commonly imagin'd) 842 bappen in Bodies, becaufe the Way, thro' which the Light might pass, is stopp'dby Particles of Matter; for Light passes thro' all the finallest Parts of Bodies; neither is such an Interception of Light of any Use to cause Opacity : It is requir'd for Opacity that the Light shou'd be reflected and deflected from a right Line, for which there is only requir'd the Separation of two Mediums *.

Let us conceive a Body, confifting of very 796 fmall Parts, perfectly transparent (fuch as are the Particles of which Bodies confift *) and fepa- * 841 rated from one another by Pores; and that those Interffices are either void, or fill'd with a Medium whole Denfity is different from that of the Particles ; if Light enters fuch a Body, it will every Moment fall upon a Surface dividing Mediums differing in Denfity; therefore it will undergo innumerable Reflections and Refractions in that Body*, fo as not to be able to get thro' * 631 it. Therefore we fee that Opacity depends upon 796 the Pores; for if you fill the Pores with a Medium 843 of the same Density as the Particles of the Bodies themfelves, the Light will undergo no Reflection or Refraction in the Body, but pais directly thro' *, and the Body will be transparent. * 841 631

'Tho 697

Tho we cannot make Experiments, whereby to fill the Pores of the Body with a Medium exactly of the fame Denfity as the Particles, yet the following Experiments will clearly enough prove Sir *Ifaac Newton*'s Doctrine concerning Opacity.

Experiment 2.] Paper becomes more transparent, when moisten'd with Water; for it fills the Pores, and differs less in Density from the Particles than Air does. Oil has the same Effect.

Experiment 3.] Take a Piece of Glass two Inches thick; and take feveral Plates of the fame Sort of Glass laid upon one another, yet fo as not to be quite two Inches thick; and you will find that these will be less transparent than the folid Piece, because of the Airbetween the Plates, which does not get into the folid Piece where all the Parts cohere.

Experiment 4.] Take twelve Plates of the fame Glafs, as near as may be of the fame Thicknefs; let fix and fix of them be laid together; if you take the leaft transparent of those two Particles, and, having dipp'd it in Water, take it out again, it will become more transparent than the other; becaufe the Water, which in that Cafe fills the Interstices between the Planes, differs less in Denfity from the Glass than Air does.

What has been faid of Opacity is farther confirm'd, and put out of all Doubt by innumerable 844 Experiments, by which Bodies perfectly transparent become opaque, by the Separation of their Parts, without the Intervention of any opaque Body.

Experi-

Experiment 5.] Let any perfectly transparent Liquid, that may be chang'd into Froth, be thak'd, till it be full of Bubbles; it will immediately become opaque, by reafon of the Interflices that are fill'd with Air.

Experiment 6.] Turpentine and Water are transparent Bodies, but, when mix'd, they become opaque.

Experiment 7.] Water and Oyl, by being mix'd together, become opaque, tho, fingly, they are transparent.

Experiment 8.7 Tho Glass be transparent, yet if it be reduced to Powder it becomes o-Paque; as it also does when it is crack'd.

We clearly fee in all these Cases that Opacity is produc'd, because there is a Medium of different Density between the transparent Parts; which may also be observ'd in the Clouds, which are epaque on account of the Air interpos'd between the Particles of the Water.

If we add to this what is faid in the 22d Chapter following, concerning the Colours of thin Plates, we shall have new Experiments, by which alone it is fully prov'd that Bodies intercept the Light, because they confist of very thin Particles, encompass'd with a Medium differing from them in Denfity.

Some opaque Bodies reflect a little Light, and the reft of the Light, by innumerable Di-Visions which it undergoes in the Reflections and Refractions above-mention'd, is extinguish'd in the Body; fuch are black Bodies; if there were any Bodics perfectly black, they wou'd reflect no 845 Light; for all Bodies, when no Light falls upon them,

them, and therefore they reflect no Rays, appear black.

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Other opaque Bodies appear to have various Colours; fome transparent Bodies are also ting'd with Colours: And whence those arise we are now to examine.

CHAP. XVIII.

Concerning the different Refrangibility of the Sun's Rays.

B ODIES appear differently colour'd, tho' they be enlighten'd by the fame Rays of the Sun which are reflected by them. Befides these there are feveral Phænomena of Light, relating to Colours, not to be overlook'd.

- 846 In these we are to observe three Things. 1/l, The Rays are to be examin'd. 2*dly*, Their Reflection is to be observ'd. 3*dly*, We are to enquire into the Constitution of the Surfaces of Bodies differently colour'd.
- 847 As to the Rays, the first Property, to be taken Notice of here, is, that in the fame Circumstance all Sorts of Rays do not undergo the fame Refraction.

DEFINITION I.

848 The Rays, which undergo this different Refration, are faid to have a different Refrangibility, and those are faid to be most refrangible, which are most inflected by Refraction.

DEFINITION II.

849 Those Rays are said to be Homogeneous, which do not differ from one another in Refrangibility.

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DEFINITION III.

And those are call'd Heterogeneous, which, un- 850 der the same Circumstances, are not equally infletted by Refraction.

Plate XV. Fig. 1.] Between A B and C D let there be a Sun Beam made up of an innumerable Quantity of parallel Rays; these are not equally refracted; for if they fall obliquely on the Surface B D of the denfer Medium, fome of them are refracted between B E and D G, and are mov'd according to that Direction in the denfer Medium ; others are more inflected, and direct their Motion between BF and DH according to the Direction of those Lines; and indeed no Direction can be conceiv'd between the Mediums, along which the Raysdo not move in every Point between B and D: So that the leaft Beam whatever is, by Refraction, divided into an innumerable Quantity of Rays; because every Sun-Beam, however small, is heterogeneous, and confifts of an innumerable Quantity of Rays refrangible in all Degrees of Refrangibility.

The above-mention'd parallel Rays falling 851 upon a plane Surface, by Refraction, are mov'd between B E and D H; which Lines diverge from each other, and, being continued, are more and more feparated; fo that the Rays abovemention'd are difpers'd by Refraction. In No. 653. We have confider'd the Rays as homogeneal, as alfo every where in the Foregoing Part; the Difference of Refrangibility is finall enough in the Rays of the Sun, not to have been worth obferving in the foregoing Propositions. Befides, we were first to examine what happens in homogeneal Rays; and what must be chang'd in the Propositions, upon account of the different Refrangibility, will easily appear.

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That this Refrangibility of the Rays may be made visible, the above-mention'd Divergence must be increas'd; which will be perform'd, if the Rays above-mention'd fall upon the Surface E H, which terminates a denfer Medium, and feparates it from the rarer Medium, and which makes any Angle with the Surface B D, and is fo inclin'd to it, that the Rays, which are more refrangible, fall more obliquely upon it than those that are less refrangible; fo that the former, going into a rarer Medium, are more turn'd out of the Way, and diverge more from the others upon a double Account, that is, both upon the Account of a greater Refrangibility, and of a greater Inclination. The Rays between B E and D G, which are the least refrangible, being refracted a fecond Time, continue their Motion between EI and GL, the others between FM and HN: In which Cafe, if these Rays fall upon a Plane at the Diftance of 15 or 20 Foot, those that are most refrangible will be separated from those that are leaft refrangible, and the whole intermediate Space is enlighten'd with Rays endow'd with a mean Refrangibility.

Experiment 1. Plate XV. Fig. 2.] Let there be an Hole in the Plate of a Metal in the Windowfhut, of about a Quarter of an Inch, thro' which a Sun-Beam enters the darken'd Room : Let this Beam fall upon a triangular Prifm of Glafs A A, in fuch Manner, that it may be wholly reflected by the lower Surface (fee the Exper. of N°. 789.) By the two Refractions which the Light undergoes, Rays that are differently refrangible do not diverge, and, being reflected, come upon a fecond Prifm B B, which is like= wife of Glafs, and triangular, in the fame Manner, as if they had come directly from the Sun. These

These two Prisms are moveable about their Axes, as we before explain'd *; the first A A * 197 is laid upon a Piece of Wood fasten'd to it, yet fo as not to hinder it from turning about its Axis; this Piece of Wood is fasten'd to a Threelegg'd Staff with a Ball and Socket, fuch as is us'd in practical Geometry. The fecond Prism B B is laid on a Frame or Stand S, in whose opposite Sides there are feveral Slits that answer one another; by these the Prism is fustain'd at different Heights, but always horizontally, the End of the Axis being plac'd in correspondent Slits.

The Light, that comes to the Prifm B B, moves perpendicular to its Axis, and paffes thro', as is demonstrated in Fig. 1. in which B D and E H reprefent the Sides of the Prifm; the Light is alfo equally inclin'd to each Side; which will happen by moving the Prilm about its Axis; for the Light of the Sun (as you move the Prifm) will afcend, and then defcend again; and the Situation requir'd is that in which the Light is higheft of all : Now both the Prifins are to be fo dispos'd, that in this Cafe the Light may go horizontally out of the Prifm B B. This horizontal Beam, at the Diftance of 15 or 20 Foot, must fall upon the Board T which is cover'd with white Paper, and has fuch a Foot, that upon it, it may be rais'd and fix'd at different Heights. The Rays come diverging to the Paper, and upon it form the Oblong Image RV, terminated at the Sides by parallel Lines, but femicircular at R and V.

If the Rays of the Sun, that go thro' the round Hole, fall upon a Plane at a certain Diftance, they will form a bright round Spot, fo much the greater, as the Plane is more diftant from the Hole; which arifes from the Rays that come

come from the Sides of the Sun, which making an Angle with those that come from its Center to the Hole, and intersecting them in the Hole, gives the Image of the Sun upon the Plane.

If the Rays did not go thro' the Prism B B, and fall upon the Plane at the Diftance of the Board T, the Image of the Sun wou'd have its Diameter equal to the Breadth of the Image VR; which Breadth is not alter'd by the Re-873 fraction, because the Rays enter the Prism perpendicular to its Axis, and, in Respect of the Breadth of the Image, are not inclin'd to it. But as, in another Respect, the Image of the Sun is oblong, it follows plainly from thence, that all the Rays have not undergone the fame Refraction; for homogeneous Rays, tho refracted, will give a round Image of the Sun. The leaft refrangible Rays go to R, and the most refrangible to V; and the whole Image R V is terminated with Semicircles at R and V, becaufe the whole Image confifts of Circular Images. Now between R and V there are innumerable Quantities of circular Images, made by Rays of all poffible Refrangibilities; otherwife the Image R V wou'd not be terminated with rectilinear Sides.

854 In most Experiments, we have faid that the Light is let into the darkened Room thro' a Slit or an Hole, which we leave to the Contrivance of the Workman; our Method was the following.

Plate XV. Fig. 3.] We made an Hole 4 Inches fquare in the Window-fhut, which on three Sides had Ledges of Wood AB, BC, CD, for abbetted, as to make Grooves to keep in the fquare flat Piece of Wood QL, which is fix Inches long, and fix Inches wide. It may be drawn out, and

and feveral Sorts of them ferve for feveral Experiments. That which we use in the following Experiment has a Hollow and a Hole in the Middle behind to contain a convex Lens, which is the Object Glass of a Telescope of 16 or 20 Foot; the round Hole in the Middle mark'd f is above half an Inch in Diameter, thro' which the Rays of the Sun, paffing thro' the Glafs, enter the Room. In the Fore-part alfo the Board Q is hollow'd, but not in the Middle; this Hollow contains the brass round Plate L, which towards its Circumference has an Hole at f. which Hole is equal to the Hole in the Board Qitfelf, and agrees with it. Together with L there is another concentric Plate, which is lefs, and moveable about the Center; this Plate has feveral unequal Holes, which fucceffively come to f as the Plate is turn'd round; fo that you may at Pleasure let the Light enter a dark Place thro' a bigger or a leffer Hole ; which in many Experiments, that are made without the Lens above-mention'd, is of good Ufe. The Pin m, join'd with the Plate, ferves to turn it by.

Experiment 2.] Let in a Sun-Beam thro' the Lens above-mention'd, thro' an Hole a Quarter of an Inch wide, into the darken'd Room; by the Rays thus let in, at the Diffance where parallel Rays are collected, the Sun is very exactly reprefented, and its Image terminated with exactly defcrib'd Bounds. For the Rays that come from the feveral Points of the Sun, which, by reafon of its immenfe Diffance, may be look'd upon as parallel, are collected in one Point at that Diffance.

Plate XV. Fig. 2.] Now if with these Rays you make the Experiment above-mention'd, the Vol. II. I several feveral Images made by the homogeneal Rays, fuppofing the Board at a just Diftance, are exactly terminated; and therefore the oblong Image R V, which is made of all these Images, is likewife regularly terminated.

This Experiment will fucceed in the fame Manner, if the Rays pass thro' a Prism made of any Body which is denser than Air.

Experiment 3. Plate XV. Fig. 2.] Let a triangular Prifm be made of Wood and two glafs Planes, and fill'd with Water; fuch as is reprefented at B B (*Plate* XVII. Fig. 3.) If you make use of this Machine instead of the Prifm B B in this Fig. the Experiment will succeed in the same Manner; and, in passing thro' Water, heterogeneous Rays in the same Manner are separated by Refractions.

Experiment 4.] If any Person, standing 15 or 20 Foot off, looks at the Hole thro' which the Light is let in, it will appear round; but if before his Eyes he hold a triangular glass Prism, or the watery one mention'd in the last Experiment, fo that the Rays coming from the Hole (after such Refractions as the Light suffers in the foregoing Experiment) reach the Eyes, the Hole will appear oblong. The Situation of the Prism will be found, if holding it horizontally with one Edge (or the refracting Angle) upwards, it be a little mov'd about its Axis, so as to make the Image of the Hole to ascend and descend; and the Prism be held fast in that Position in which the Hole appears most depres'd.

This Experiment, as well as the foregoing ones, proves the different Refrangibility of the Rays; for by the homogeneal Rays of each Refraction, the Hole appears in the imaginary Foci of the Rays coming from the feveral Points of

of the Hole*, which Image is round; the Rays, * 737 which undergo a different Refraction, enter the Eyes in different Directions; and you have feveral Images, all which form the oblong Image, which is really feen.

But that this different Refrangibility does not de-855 pend upon the refracting Medium, but the different Conflitution of the Rays themsfelves, is prov'd, because those Rays, which in one Cafe undergo the greatest Refraction, are in any Refraction turn'd out of the Way more than any.

Experiment 5. Plate XV. Fig. 4.] All Things being difpos'd, as in the first Experiment, at any Diftance from the Prifm B B; let the oblong Image of the Sun fall upon the Vertical Prifm C C, which is also moveable about its Axis; as the Rays go thro' it, let it be turn'd about its Axis, and left fix'd where the Rays are leaft of all turn'd out of the Way by Refraction thro' the Prifm : In that Cafe the Rays are refracted in the fame Manner thro' this Prifm as thro' the first, but they are not dispos'd the fame Way, for that wou'd form a fquare Image. But here the Image, keeping the fame Breadth, is inclin'd at R V, the Rays at V being mostly turn'd out of the Way, as in the Refraction thro' the first Prifin BB.

The Demonstration, before given *, of the con-856 frant and fettled Ratio between the Sines of the An-*639 gles of Incidence and Refraction, may be referr'd to any homogeneous Rays; but confidering the different Refrangibility of the Rays, this Proportion varies, as it follows from the Experiments of this Chapter.

But the Refrangibility in all the different Sorts 857 of Rays is every Way unchangeable, as will appear by the Experiments to be mention'd hereafter.

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CHAP. XIX.

Concerning the Colours of the Rays, and their Unchangeablenefs.

858 HE different Refrangibility of the Raysgoes along with the different Colours; and every Sort of Rays, as they are more or les infleted by Refraction, have a particular Colour of their own, and which is wholly unchangeable.

In Refpect of the Colours, the fame Thing is to be obferv'd, as has been taken Notice of in regard to other Senfations *; the Colours are Ideas which have nothing common with the Rays, by which they are excited: Therefore we mult define what we mean by colour'd Rays and colour'd Objects.

DEFINITION I.

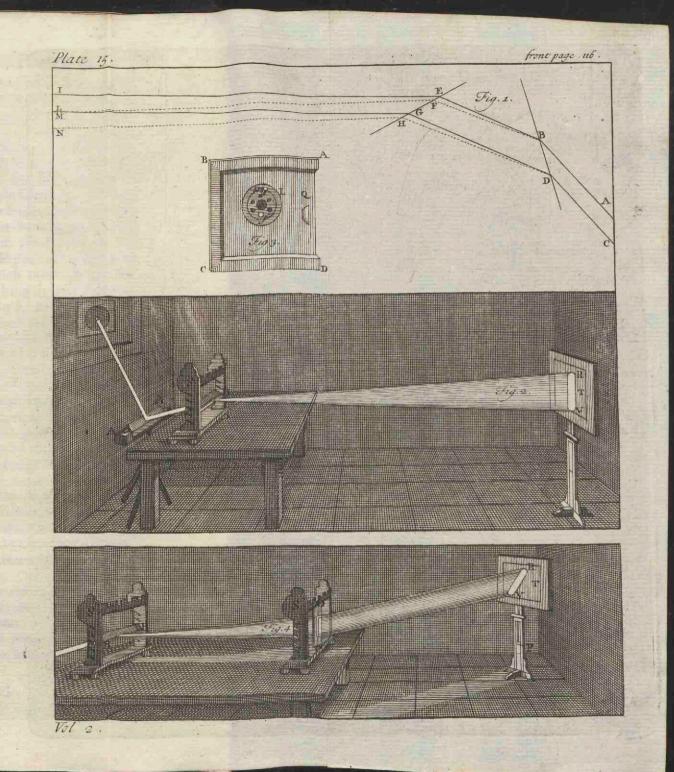
859 An Object is faid to be of fuch a Colour, whole Idea is excited in the Mind by the Rays reflected from the Object.

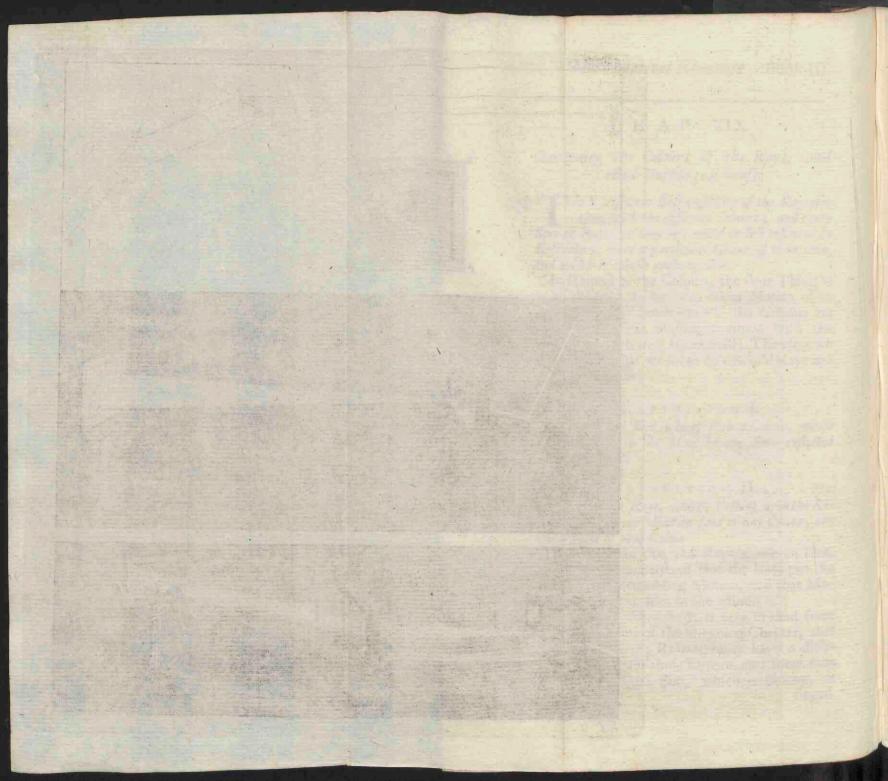
DEFINITION II.

860 Homogeneous Rays, which, striking upon the Retina, excite in our Mind an Idea of any Colour, are call'd Rays of that Colour.

We have faid that the Rays excite an Idea, by which we underftand that the Rays put the Fibres into a trembling Motion, and that Motion occafions an Idea in the Mind.

Plate XVI. Fig. 1.] It is very evident from the Experiments of the foregoing Chapter, that Rays of different Refrangibility have a different Colour; for those Experiments shew that the Image of the Sun, which is oblong, is tinged





tinged with different Colours. The Rays, which 861 are least turn'd out of the Way by Refraction, are red; the rest of the Colours are in the following Order, namely Orange, Yellow, Green, Blue, Purple, and Violet, of which laft Colours are the Rays that are most refracted *. The oblong Image of the Sun above-mention'd, as has been faid *, is * 853 made up of an innumerable Quantity of round Images; if their Diameters be diminish'd (which is done by intercepting the Sun's Rays, fo that only those that come from the Center of the Sun shall pass thro' the Prism) the Centers of the particular Images which make up the oblong Image will not be chang'd; and therefore the Length of the Image a b, between parallel Lines, is not chang'd : And this alone wou'd remain the fame, if the Breadth of the Image, fhou'd be infinitely diminish'd : And therefore this Length alone is to be confider'd in determining the Limits of the Colours in the Image itfelf; these Colours are represented in this Figure by the Letters a, b, c, d, e, f, g, b, and the Number which is fet down against each Colour denotes the Space taken up by it in the Image, the whole Lenth of the Image being divided into 360 Parts.

If the Breadth of the oblong Image of the Sun be 862. diminifb'd, the heterogeneous Colours are more feparated in the Image; becaufe there are fewer particular Images confounded in the feveral Points, where Rays of different Refrangibilities differ little from each other.

The Colour of any Ray, as alfo its Refrangibility, 863 cannot be changed by any Refractions, or Reflexions, or mingling of the Rays together.

In this Chapter I shall speak of Reflexion and Refraction, and of the Mixture of Rays in the following one.

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That,

864 That the Refrangibility is not chang'd, by Refraction, is prov'd by the *fth* Experiment of the foregoing Chapter; which may also be referr'd to Colour; but is more clearly evinc'd by the following Experiment. Concerning which it is to be observ'd (as may be also faid of what follows) that the Experiments are to be made with Prisms of clear Glass, free from Veins; for they occasion the Light to move irregularly; and the Rays are not duly scattered by the Refractions.

Experiment 1. Plate XVI. Fig. 2.7 Every Thing being as in the first Experiment of the foregoing Chapter, you must make the Experiment with a Sun-Beam, going thro' an Hole of half an Inch Diameter: The Frame or Stand S, in which the Prifm is, must be fuch, that the little flat Board t, to be us'd in this Experiment, may move between its Sides ; this Board has an Hole F of s of an Inch Bore, thro' which the Light refracted by the Prifm is transmitted, whereby the Rays in the oblong Image are better feparated from one another, the Rays being mostly intercepted; this oblong Image R V, at the Diffance of 10 or 12 Feet from the Prifm, falls upon the Board t of a Stand like the former; in which Board alfo there is a fmall Hole F like that in the first ; thro' this the Rays passupon a second Prifm laid upon this Stand, and are refracted in the fame Manner as in the first. By moving the first Prism a little about its Axis, the Image R V afcends or defcends, whereby the different Rays are fucceffively transmitted thro' the Hole; in every Cafe, the Rays refracted thro' the fecond Prifm, and ftriking upon the flat Board T which is cover'd with white Paper, are not dispers'd at H i but the Image is round, if the Rays fall perpendicularly upon the Paper, and alfo of the fame Colour

Colour as the Rays falling upon the fecond Prism. Yet the Image H is so much the more lifted up, as the Rays, by Refraction thro' the first Prism, are more turn'd out of the Way; that is, those, that are most refrangible in the one Case, do also fuffer the greatest Refraction in the other.

It will appear also by the following Experi-865 ments, that the Refrangibility and Colour are not chang'd by Reflexion.

Experiment 2.] The Rays which, for Example, make the red Part of the oblong colour'd Image, whatever Body they are reflected from, are always red; that is, all Bodies appear red in that Light: In the violet Light they are violet; green in the green; and so on in the other Colours.

This will appear by trying it with Vermillion, Orpiment, blue Bice, or Cloth of any Colour, Bc.

Experiment 3. Plate XVI. Fig. 3.] Let the Light enter the dark Room thro' two Holes of a Quarter of an Inch Diameter each, in the round Plate moveable in the Window; let thefe Beams be about two Inches afunder, and reflected by the plane Mirror S.

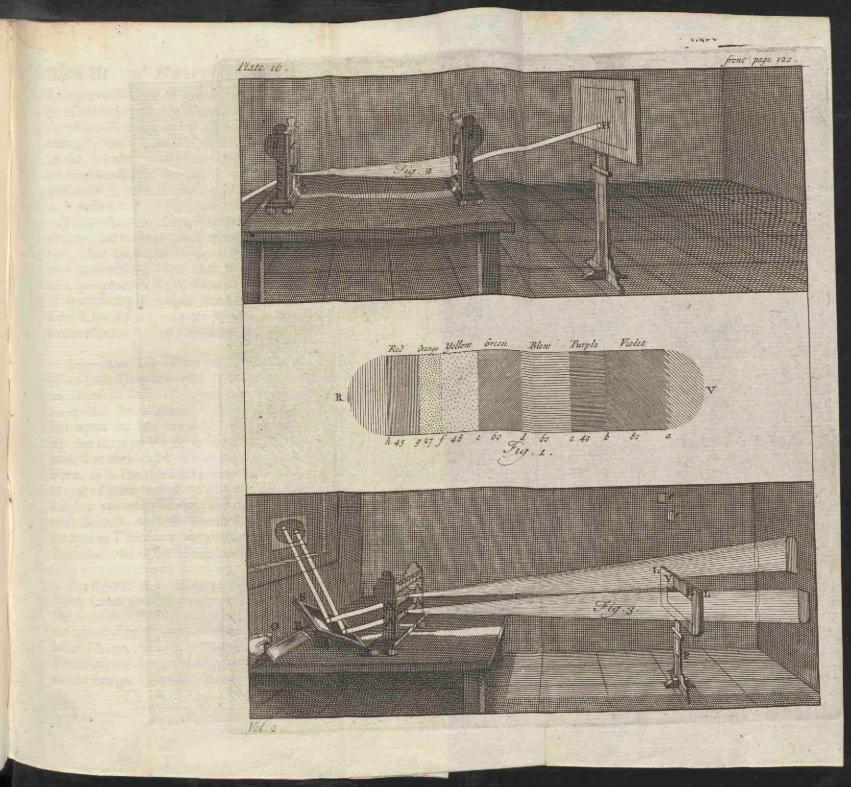
The Plate and Mirror muft be fo difpos'd that the two Beams may fall upon two Prifms A A, which are laid horizontally upon the fame Frame; fo that the oblong Image, produc'd by the Refraction of those Prifms, may touch one another at their Sides. Let one Prifm be turn'd a little about its Axis, that the Red of one Image may be just on the Side of the Violet of the other : Let these Colours, and none of the reft, be intercepted by a wooden Ruler that has a white Paper pasted upon it, fo that the I 4 Red

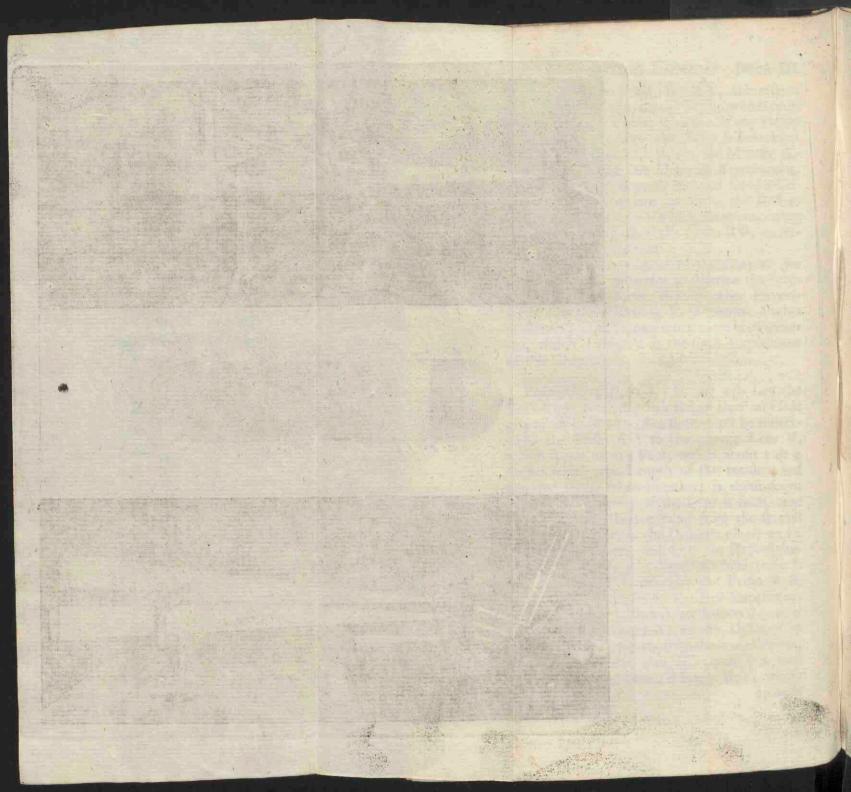
Red be at R, and the Violet at V, the reft of each Image falling upon the Wall, which mult be cover'd with a black Cloth. If any Perfon ftand at O, and thro' the Prifin B B looks at those Colours, R and V, (in the Manner defcrib'd concerning the Hole, in *Experiment* 4. of the foregoing Chapter) he will see the Colours separated from one another, the Red at r, and the Violet at v; which therefore, being reflected in going thro' the Prism B B, undergoes a greater Refraction.

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In the first Experiment of this Chapter we gave a Method, whereby to separate the heterogeneous Rays better than in other Experi-866 ments; in the following Experiments, Lights of divers Colours become much more homogeneous, which is requir'd in the fixth Experiment of this Chapter.

Experiment 4. Plate XVII. Fig. 1.] Let the Sun's Light enter the dark Room thro' an Hole of 4 of an Inch ; the Sun Beam must be reflected by the Prifm A A to the convex Lens V. which stands upon a Foot, and is about 3 or 4. Inches wide ; the Length of the incident and reflected Rays, taken together, is about feven Feet; the Convexity of the Lens is fuch, that the Foci of the Rays coming from the feveral Parts of the Hole, at the Diftance of 10 or 12. Feet from the Lens, will form the Reprefenta-* 711 tion of the Hole, if a Paper be held there *. Just beyond the Lens place the Prifm B B, whereby the Rays (as in the first Experiment of the foregoing Chapter) are difpers'd ; now placing the flat Board T, at the Diftance at which the Rays, coming from the feveral Points, are collected, there will fall upon it a well terminated oblong colour'd Image R V, whole Length





Length will have a greater Proportion to its Breadth than in other Experiments; and which therefore is made up of Rays more homogeneous*; and fo much the lefs mix'd, as the Hole * 862 in the Window is the lefs. By moving forward or backward the Board T, one may find the Diffance at which the Image is most diffinct and terminated without the Penumbra.

Now that the Rays, feparated in this Experiment, are homogeneous enough, is prov'd from this, that they can be no more difpers'd by a new Refraction : And that this is the Property of homogeneous Rays, follows from their equal Refrangibility ; which has been already confirm'd by an Experiment *, and is more fully * 864 demonstrated by the following one.

Experiment 6.] Take two Circles of Paper of an Inch Diameter each, and let Light fall upon them in fuch a Manner, that the one may have the homogeneous Rays of any Colour fall upon it, and the other may be enlighten'd by the Rays of the Sun; if both these Papers be look'd at thro' the Prism at the Distance of fome Feet, as in the third Experiment of this Chapter, the Circle, enlighten'd by the heterogeneous Light, appears oblong, and ting'd with different Colours, as in the fourth Experiment of the foregoing Chapter; but neither the Colour nor Figure of the other Circle is chang'd.

Experiment 6.] Upon a white Paper draw black Lines parallel to one another, and about the fixteenth Part of an Inch wide; let thele be enlighten'd by throwing upon the Paper the oblong Image, which is deferib'd in the fourth Experiment, fo that the Lines may lie long-wife in the Image. Befides this you muft have a convex

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vex Lens, about five or fix Inches wide, fuftain'd upon a Foot, fuch as are reprefented at V (Plate XIX. Fig. 2.) which collects the red Rays that come from a radiant Point, fix Foot diftant from the Glafs, at an equal Diffance on the other Side. If this Lens be plac'd at the Difance of fix Feet from the Image above-mention'd, the Parts of the black Lines that are enlighten'd by the Red falling upon the Paper, by means of the Rays which are collected by the Lens, will be exactly reprefented in the red Image at the Diffance of fix Feet; but you must bring the Paper forward about three Inches and a Half, to make those Parts in the Lines which are enlighten'd by the Purple, to appear diftinct in that Part of the Image which is of the fame Colour : The intermediate Colours give Images at intermediate Diffances; the Violet is fo weak, that the Threads cannot be reprefented in it.

This Experiment does also confirm, that the Colour of reflected Rays is not chang'd by a new Refraction thro' the Lens; as also that the most refrangible Rays are more inflected than the others in passing thro' the Lens.

867 This laft Experiment does also prove, that the different Refrangibility of the Rays is the Caufe that kinders the Perfection of Telefcopes. For the Foci of the Points, that are equally diffant, fall at different Diffances from the Lens, according to their different Colours; whence all the Reprefentations of the Points are unequally diffant from the Eye-Glafs; and therefore they cannot be all perfectly feen thro' it.

868 Concerning the Reflexion of the Rays it is to be observ'd, that those Rays are more easily refletted totally, which have the greatest Refrangibility; for the greater is the Refraction of the Rays, the Book III. of Natural Philosophy. 123 the lefs Obliquity is required to make them be totally reflected*. *793

Experiment 7. Take a Prifm, plac'd as before *, in the Experiment where it was observ'd, * 790 that by moving the Prifm about its Axis, the Rays that first went thro', when their Inclination is increas'd, become wholly reflected; but if the Prism be gently mov'd in this Case, we shall perceive that the violet Rays are the first which will be wholly reflected, then the purple Rays, and all the rest in the same Order, as they are in the oblong Image of the Sun, so often mention'd; which appears, if the reflected Rays be separated by the Refraction of the Prism.

CHAP. XX.

Of the Mixture of Colours, where we shall Speak of Whitenefs.

THAT the Refrangibility and Colour of 869 the Rays are not chang'd, by the Mixture of Rays of different Refrangibility, has already been faid*, and now we must prove it by Expe- * 863 riments.

Experiment 1. Plate XVI. Fig. 3.] This Experiment must be made in the fame Manner as the third of the foregoing Chapter; here you must make the Red R, and the Violet V be confounded and mix'd together, by throwing them upon the fame Part of the Ruler L L, which thereby will appear of a purple Colour in that Place. But if a Perfon looks at them thro' a Prifm, the Colours will appear feparated; and therefore neither the Colour nor Refrangibility are chang'd by this Confusion of Colours.

Expe-

Experiment 2. Plate XVII. Fig. 2.] If the oblong colour'd Image of the Sun (of which Memtion was made in the first Experiment of the 18th Chap.) falls at RV upon the convex Lens L. mention'd in Experiment 4, of the foregoing Chap. which must be placed at fix or feven Foot from the Prifm B B, the divergent Rays will converge by the Refraction of the Lens, and interfect one another at the Diffance of 9 or 10 Feet at A; if the Board T be placed at a greater Diftance, the Rays, which diverge again after the Interfection, will fall upon it difpers'd; and you will again have an oblong colour'd Image; but the Colours by reafon of the Interfection at A, will appear in a contrary Order, and will not be chang'd by having been mix'd together at A.

Experiment 3. Plate XVII. Fig. 2.] Every Thing ftill remaining as in the former Experiment, if with a black Paper you intercept fome of the Rays of the Image R V, which changes the Mixture (which this Way may be varied in any Manner) the Colours of the other Rays, that are again feparated, are no Way chang'd.

870 If the Rays of the Sun, as they come to us, are wholly reflected by any Body, that Body appears White; but these Rays are an Heap, or Parcel
* 847 of Rays of various Colours *; whence we de* 858 duce, that a Mixture of different Colours makes a
871 Whitenefs; for if the Colours which are observed in the oblong Image of the Sun, fo often mention'd, be mix'd and confounded together, in the fame Proportion as they are in that Image, a Whitenefs will be produced; which also proves, that in that Respect the Rays are unchangeable. The Rays that come from the Sun appear white;

white; but, if they are feparated, their Colours are difcover'd; and, if they be mix'd again, the Whitenefs will be reftor'd.

Experiment 4. Plate XVII. Fig. 2.] Things being difpos'd as in the two foregoing Experiments; let the Board T be plac'd at A, in the very Place where all the Rays of the Image R V are confounded together; there will be a Whitenefs at A; if with a black Paper you intercept the Red of the Image R V, the White will vanish, and the Colour at A be bluish; but if you intercept the violet and blue Rays, the White becomes reddifh.

Experiment 5. Plate XVII. Fig. 3.] Take three triangular Prifms, made of Wood and Plates of Glass which contain Water, as BB, DD, DD; the Plates of Glass in each of them make an Angle of about 70 Degrees: The Length of the Plate is of fix or feven Inches, and their Breadth f three Inches; these Plates are fix'd in another Manner in the Prifms D D, D D, than in the Prifm BB, fo as to make their Bafes bigger, that is, hefe Prifms are fhorter and deeper than the ther. Let the Sun's Rays be refracted thro' the rifm B B, as in the 3d Experiment of the 18th Chap. And let the oblong Image of the Sun at the Diftance of three or four Foot fall upon the Surtace of the Prism D D, plac'd parallel to the Surface of the Prifm BB, out of which the Rays go. In the fecond Prifm the Rays undergo a contrary Refraction than in the first; because of the Parallelifin above-mention'd : And be-Caufe the Edge of the Angle, form'd by the Glass Plates in the Prifm BB, is turn'd upwards, and that of D D downwards; therefore the first Refraction is deftroy'd by the fecond, and the Rays go out of the Prifin DD parallel to one another,

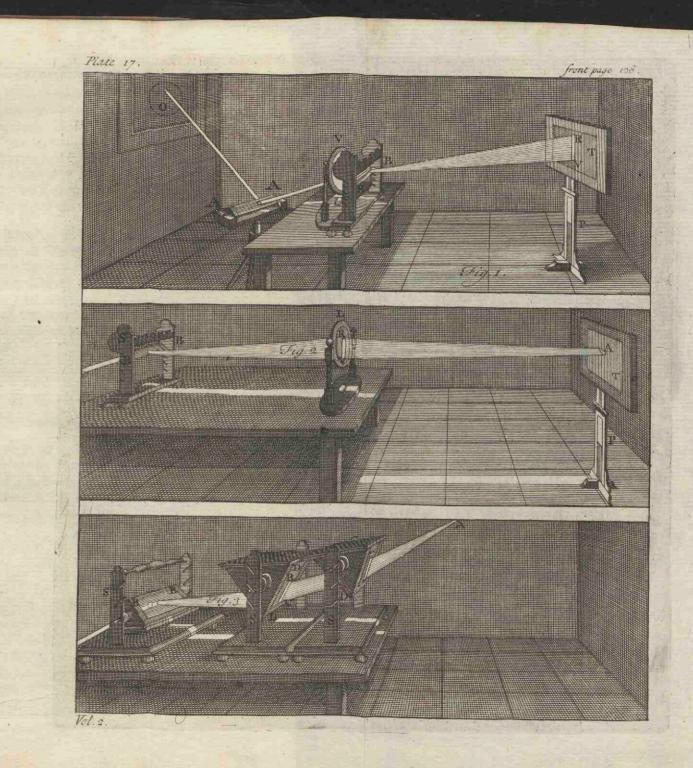
another, and fall upon R V ; for if the Prifms be brought together, fo that the parallel Surfaces may touch one another, the Light will pass thro' a Medium terminated with parallel Planes, (which is form'd by the two Prifms join'd together) thro' which Light, of any Degree of Refrangibility, will pass without Change of Di-* 628 rection*. Now the Prifms are feparated, that the heterogeneous Rays may be feparated, before they again become parallel; if these colour'd Rays fall upon the third Prifm DD, and in paffing thro' it undergo a Refraction like that which they have undergone in paffing thro' the first or fecond Prism; the Rays that go out at rv converge, on account of unequal Refractions in Rays of different Colours, and concur at A, in which Place alfo Whitenefs will be produc'd, as in the foregoing Experiment.

> Experiment 6.] If an oblong colour'd Image of the Sun be made, after the Manner deferib'd in Experiment 1. of the 18th Chap. and a Perfon flanding at the Diffance of the Prifm, that refracts the Light, looks at the Image thro' the faid Prifm, as was done in refpect of the Hole in the fourth Experiment of the fame Chap. he will fee a round and white Image; the fecond Refraction defiroying the first; for thereby the Rays, being again mix'd, enter the Eye, and in this Cafe make the Object appear white.

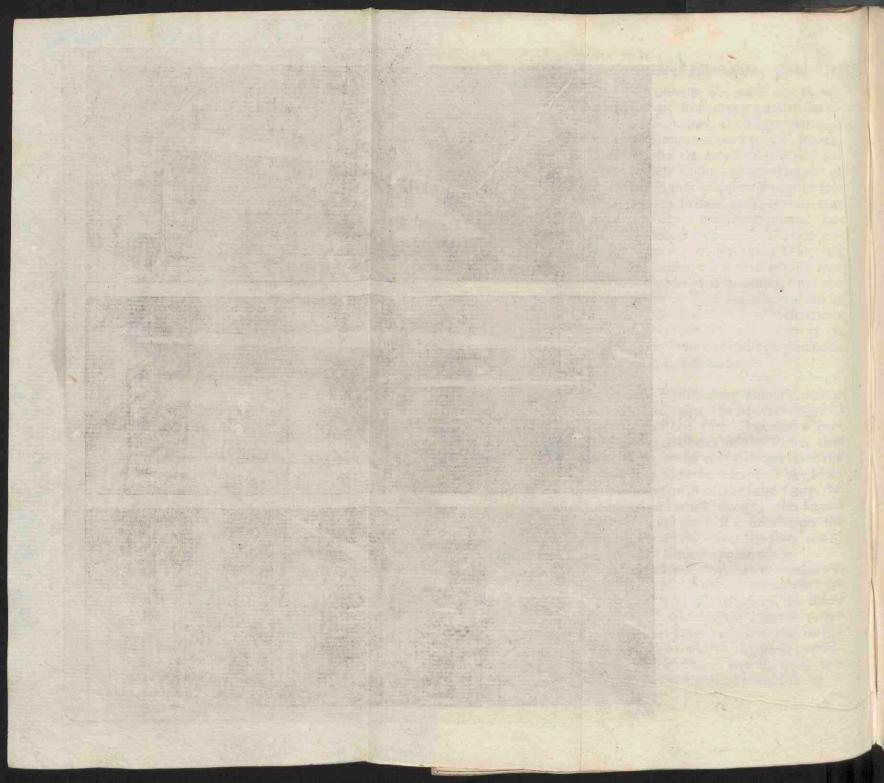
872 The Mixture of all the Colours, which are observed in the oblong Image of the Sun, is not necessary for producing White: The Whiteness of the Sun's Rays is a little inclin'd to yellow; and therefore if Part of the yellow Rays be taken out of the Mixture, the White will be the more perfect. From the Mixture of four or five Colours, in a just Proportion, White will be produc'd.

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Primary, that is, homogeneous, Colours, being 873 mix'd, produce innumerable Colours, different from the primary or homogeneous ones. From a Mixture of other Colours one may produce a Colour like that which is homogeneous; but when there can be no Difference observ'd by the naked Eye between an homogeneous and a mix'd Colour, one may perceive a sensible Difference thro' a Prism.

Experiment 7.] Thro' a Prifin look at any 874. fmall Objects, fuch as Letters upon Paper, Flies, and other fuch things; if they be exposed to the common open Light, they will appear confused; but if they be enlightened by the homogeneous Light of the fourth Experiment of this Chapter, they will appear diffinct when feen thro' the Prifin.

CHAP. XXI.

Of the Rainbow.

AVING made an End of confidering what relates to the Rays, whereby Bodies are enlighten'd; before we leave this Subject, we must explain a Phænomenon, which is too remarkable and common to be past by in Silence.

The Iris or Rainbow is what every Body has often feen; we must explain what is the Cause of it, having first laid down some Things for that Purpose.

Plate XVIII. Fig. 1.] Let there be a dense Me- 875 dium encompass d with a rarer, terminated by the Circle BDFH. Let homogeneous Rays that are parallel to one another fall uponit, and let AB be one of these Rays; let the Semidiameter CB be drawn and continued to N; it will be perpendicular

cular to the Surface separating the Mediums: A B N therefore is the Angle of Incidence ; This Angle is equal to the opposite vertical Angle CBL, whole Sine is CL, perpendicular to BL and going thro' the Center; the Ray is refracted * 624 towards the Perpendicular*, and C B M is the Angle of Refraction, whole Sine is CM drawn from C perpendicular to BD: There is the fame * 639 Proportion between fuch Lines as CL and CM*, for every Ray, as there is for A B. The Ray BD does in Part penetrate into the rarer Medium in the Direction DE, and is in Part reflected along DF; and makes the Angle of Reflexion * 805 CDF equal to the Angle of Incidence * B DC; whence BD and DF are equal. The Ray DF does also in Part go out of the denfer Medium along FG, and is in Part reflected along FH; which in the fame Manner docs in Part go thro' HI, and is in Part reflected : But this Reflexion, and other farther Reflexions and Refractions we shall not confider; they are too weak, on account of the feveral Divisions that the Light has undergone.

Plate XVIII. Fig. 1.] The Ray FG, which after one Reflexion goes out of a denfer Medium, makes the Angle GPA with the incident Ray AB, which varies in different incident Rays; therefore, tho these Rays shou'd be parallel, they will be featter'd when they go out, after one Reflexion, as may be feen by the fecond Figure.

The Ray EE, which, being continued, goes thro' the Center C, is not turn'd out of its Way * 806 either by Reflexion or Refraction *. 629

As you recede from this Ray, the returning Ray is continually lefs and lefs inclin'd to the incident one. So the Ray DD, which goes out of a denfer Medium along d d, and returns along the Line, makes a greater Angle with d d, than

than the intermediate Rays between D D and E E do make with their returning Rays, which go out of the denfer Medium.

Let B B be a Ray, in respect of which this 876 Inclination is the least of all, that is, which makes the Angle G P A (Fig. 1.) the biggest of all. Beyond B B the returning Rays are more inclin'd to the incident ones; thus A A returns along aa.

From this Difperfion of the returning Rays re-877 ceding from the denfer Medium, they become continually weaker and weaker, and their Colour cannot be perceived throughout the whole Space which they fill, tho the Colour of the incident Rays be vivid. The Colour in the returning Rays is only fenfible, where the neighbouring Rays are parallel, and those next to them diverge out very little; fo that at a great Diffance they are denfe enough to be perceiv'd. These only are faid to be efficacious, and will be fuch, when the incident Rays which are near one another, being refracted, concur in the very Point of Reflexion.

Plate XVIII. Fig. 3.] Let A B, a b, beneighbouring Rays parallel to one another, falling upon a circular Surface that terminates a denfer Medium; if these being refracted along B D, bd, do concur in D the Point of Reflexion, the reflected ones D F, D f, will make the fame Angles with one another as the incident ones D B, Db; therefore the refracted Rays F G, fg, will be pa- * 627 rallel * and efficacious*. In this Case the follow- * 877 ing Method will ferve to determine the Angle made by the incident Ray with the returning one; that is, the Angle A P G, which here is the greateft of all.

Let us call the *Ratio* of I to R, that which 878 is found between the Sines of the Angle of Incidence and Refraction, when the Light goes Vol. II. K out out of a rarer Medium, by which a denfer is encompals'd, into a denfer contain'd in the Circle itfelf. Therefore, having drawn Cm perpendicular to bD, and the Arc mn with the Center C and Semidiameter Cm,

l, R :: CL, CM :: Cl, Cm :: CL-Cl=Ll, CM-Cm=Mn.

Draw B o perpendicular to B L, and alfo B pperpendicular to B D; and draw bp, fo that it may make a right Angle with B p; laftly, join together by Lines the Points B, C, and M, m, the Triangles B bo, BCL are fimilar; for they are rectangular, and the Angles o B b and CBL (the Difference of each of which from a right Angle is the Angle o BC,) are equal.

It may be prov'd in the fame Manner, that the Triangle BMC and Bbp are fimilar; the Triangle Mmn, which is rectangular at n, is alfo fimilar to this, for the Sides M n B p, which are perpendicular to the Line B D, are parallel; as alfo Mm and Bb, becaufe the Lines BD, b D, are bifected at M and m into equal Parts. Therefore alfo B b is the Double of M m, and Bp the Double of m n.

Hence we deduce

BC, BL :: Bb, Bo. BC, BM :: Bb, Bp.

Therefore

BL, BM :: Bo = Ll, Bp = 2 Mn :: I, 2 R :: CL, 2 CM, by comparing these Proportions with the aforesaid Proportion.

Now fince the Squares of proportional Quantities are themfelves proportional, you will have

BL⁹, CL⁹ :: BM⁹, 4CM⁹. Whence we deduce

 $\begin{array}{c} BL^{q} + CL^{q} :: BC^{q}, BL^{q} :: BM^{q} + 4CM^{q} \\ = BC^{q} + 3CM^{q}, BM^{q} = BC^{q} - CM^{q} \\ BL \times LC^{q} - CM^{q}, BM^{q} = BC^{q} - CM^{q} \\ \end{array}$

By fubtracting the first and second Term from the third and fourth, (which does not change the Proportion) you will have

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BC9, BL9, :: 3 CM9, LC9-CM9, :: 3 Rq, Iq-Rq; for there is the fame Ratio between CM and LC as between R and I.

If therefore you know the Ratio between R and I, you have the Ratio between the Semidiameter BC, and the Line BL, which is the Sine of the Angle BCL, which Angle therefore is given ; and by Confequence you have the Arc BN, and alfo F G, for they are equal.

Having the Sine BL, you have also the Sine of the Angle B C M; because (as we have shewn before)

BL, BM : 1 I, 2 R. Therefore the Arc BD is determin'd, to which DF is equal.

Hence we may eafily find the Arcs NH and 870 BF; if the former be taken out of the latter. and the Remainder be divided into two equal Parts, we shall have, as is well known, the Meafure of the Angle APG.

When the Ratio between I and R varies, the Angle APG is changed, which therefore becomes different, according to the different Refrangibility of the Rays.

If the above faid Surface be enlightened by hetero- 880 geneous Rays, as they flow from the Sun, the efficacious Rays of different Colours do not make equal Angles with the incident ones, and so by the Help of this Refraction the Colours are separated.

Plate XVIII. Fig. 4. Experiment 1.] Let the Sun's Light enter the dark Chamber, thro'a Slit in the moveable round Plate O, and being horizontally reflected from the Looking-glafs S, let It pass thro' a Slit in the little Board or Stand T; 25

630 as has been before explain'd. Take a Phiol exactly cylindric, made of clear Glafs, and full of Water. Let the Beam or Ray at f g fall upon the Surface of the Phiol, it will be refracted in the Water towards bi, and there reflected, and will go out of the Phiol at 1 m. The Phiol may be eafily fo placed, as to make these Rays efficacious; and, because of the Breadth of the incident Ray or Beam, efficacious Rays of all Colours will go out of the Phiol in the fame Time; for they are very little diftant from one another in their Incidence. If these efficacious Rays fall upon a white Paper, at the Diftance of 4 or 7 Feet from the Phiol, they will produce vertical Fascia or Pillars, of various homogeneous Colours, arifing from the efficacious Rays of each Colour; if allo, the Eye be placed any where at the Distance of some Feet from the Phiol, as at N, in these efficacious Rays, it will see in the Phiol that Colour whofe Rays enter the Eye, and by a fucceflive Motion of the Eye, it will perceive all * 861 the afore-mention'd Colours.*

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881 Plate XVIII. Fig. 5.] But as to those Rays, which after a double Refraction in a denser Medium emerge, they will be efficacious, if they are parallel after the first Reflexion; for then F H, f h, are in the same Manner inclin'd to Hb as BD, bd, to Bb; and therefore supposing the incident Rays AB, ab, to be parallel, the emer-* 927 ging Rays H I, bi, will also be parallel *.

In this Cafe dD is Half of the Difference between the Arcs DE and df, or DB and db, but their Difference is Bb - Dd; if therefore this laft be fubtracted from the other, there will remain the Double of the Arc Dd, whole Triple therefore is Bb. If by Lines the Points D, d, and B, b, be join'd, the Triangles B E b and DEd will be fimilar, as is known; which there-2

fore obtains, tho' the Arcs Bb, Dd, be fo very fmall, that they may be taken for right Lines.

There is therefore that Ratio between E D and E b, which obtains between these Arcs, that is, ED is the third Part of E b, or EB; because we suppose the Arc B b to be extremely small. MD therefore is divided into two equal Parts at E; and ME is a third Part of EB.

Now if (as in Fig. 3.) the Triangles B o b, B p b, and M m n, be form'd, M m will be the third Part of B b, and B p the Triple of M n; now if, mutatis mutandis, we apply to this Figure what was demonstrated with respect to Fig. 3.* * 878 because in this B p is equal to 3 M n, whose Square is 9 M n, we shall have

BC9, BL9:: 8 R9, 19-R9. From which Proportion, as was faid of Fig. 3. the Arc BN may be difcover'd to which HG is equal; and becaufe in this Cafe

BL, BM :: I, 3 R,

You have also the Arc BD, to which (because the Angles of Reflexion are equal to the Angles of Incidence) * DF and F H are equal. *805

From which Things being found, we may 882 eafily deduce the Arcs G F D N, and B H, whole half Difference is the Measure of the Angle H P B, which is form'd by the emerging Ray with the incident one; which Angle in this Case is the least of all that are like it, and is different according to the different Refrangibility of the Rays. Whence also in this Case the effi-883 cacious Rays of various Colours, fupposing the incident ones parallel, are feparated after a double Reflexion.

Plate XVIII. Fig. 4. Experiment 2.] This Experiment is perform'd after the fame Manner as the former, only the Situation of the Phiol must K 3 be

be a little chang'd, that after two Reflexions in the Phiol, the Rays may come to the Eye or Paper towards N.

What has been hitherto explain'd, may be ap-884 plied to the Rainbow; for which Phænomenon it is requir'd, that Drops of Water (bould be fuspended in the Air; that the Spectator should be placed with his Back towards the Sun, between it and the Drops; and that there should be a dark Cloud behind the Drops, that the Colours may be more fenfible, which are fearcely perceiv'd, if vivid Light enters the Eve at the fame Time.

If these Things being suppos'd, we conceive each Drop to be cut by Planes paffing thro' the Sun and the Eye of the Spectator ; what has been demonstrated of a Medium, terminated # 878, with a circular Surface, * may be applied to each 879, of these Sections.

880, Now here we fpeak of Rays that penetrate

- 881, out of Air into Water. In red Rays, that is, ⁸⁸², but of this have leaft of all refrangible, the *Ra*tio between the Sine of the Angle of Incidence and the Sine of the Angle of Refraction, that is, between I and R, is that of 108 to 81; or which is the fame, of 4 to 3; with which Numbers, if the Computation be made, the Angle
- * 879 APG (Fig. 3.) will be 42 Deg. 2 Min. *, and the Angle API (Fig. 4.) will be 50 Deg. 57
- *882 Min. * If we speak of the violet Rays, I and R are to one another as 109 to 81; which Numbers give the Angles APG (Fig. 3.) of 40 Deg. * 879 17 Min*, and API (Fig. 4.) of 54 Deg.7 Min.*

Plate XIX. Fig. 1.] Let the Drops be fuppos'd to be diffus'd in the Air, and enlighten'd by the Sun's Rays, which are parallel to one another, and to the Line OF, that passes thro' the Eye of the Spectator. Let the Lines e O, EO, bO, BO, be conceiv'd to be drawn; and let 3

let the Quantity of the following Angles be thus, (viz.) eOF of 40 Deg. 17 Min. EOF 42 Deg. 2 Min. b O F 50 Deg. 57 Min. BOF 54 Deg. 7 Min. These fame Lines with the incident Rays de, DE, ab, AB, form Angles that are respectively equal to the aforemention'd ; therefore, if the Drops be conceiv'd at e, E, b, B. the violet efficacious Rays, after one Reflexion in the Drop e, enter the Eye ; and the red efficacious Rays come to the Eye from the Drop E; in like manner after one Reflexion, the other intermediate Colours are observ'd between e and E, in the above-mention'd Order*. Aftertwo * 861 Reflexions in the Drop, the efficacious red Rays come to the Eye from the Drop b; and the efficacious violet ones from the Drop B; the intermediate Colours appear between these Drops after the fame Manner as between E, e; but they are difpos'd in a contrary Order, and by reafon of their double Reflexion are alfo weaker.

Let us conceive a Line at O e, to be revolv'd about a fix'd Line OF, preferving the Angle e OF, and to form a Cone or Part of a conic Surface : In every Situation the Line eO will, with the Rays of the Sun that are parallel to one another, and to the Line OF, form an Angle of 40 Deg. 17 Min. If therefore the Drops 885 be diffus'd near Part of the Surface of this Cone at the fame or different Diffances, the Eye will fee a violet Arc or Bow : The fame may be faid of the other Colours; and therefore Drops being fuspended in the Air, it fees an Arc or Bow of the Breadth e E, ting'd with the homogeneous Colours before-mention'd *, that are dispos'd in the * 861 fame Order as in the Experiments with Prifins; because the heterogeneous Rays are separated as much in the Drops as in the Prifm*, * 852 K 4 By 880

Mathematical Elements Book III. 886 By the fame Way of reafoning it is plain there will be a broader Arc or Bow furrounding the first, in which the same Colours will appear in a contrary Order, and weaker.

Experiment 3.] Hang up a black Cloth in the Light of the Sun; and let a Spectator look at it standing between the Sun and the Cloth ; then, if Water be made to fall in fmall Drops between the Spectator and the Cloth, the Spectator will fee a Rainbow, at least the inner one.

CHAP. XXII. Of the Colours of thin Plates.

N E pais on to the Colours of natural Bodies; and before we go any farther, we think it proper to examine thin Plates. Whoever attentively has observ'd very thin Glass, or Bubbles made of Water thicken'd with Soap, must have perceiv'd feveral Colours in them. Rays of Light, by Help of a thin transparent Plate, are separated from one another, and ac-887 cording to the different Thickness of the Plate, the Rays of some Colours are transmitted, and those of others are reflected; and the same very thin Plate is of another Colour, when feen by the transmitted Rays, than when seen by the reflected ones.

Experiment 1. Plate XVIII. Fig. 6.] Take two Object-Glaffes belonging to long Telescopes, AB and CD; and let one of them be laid upon the other ; then let them be prefs'd together hard, and in the Middle, where the Glaffes touch one another, you will fee a transparent Spot, which is incompafs'd with colour'd Circles. If the Light, reflected by the Air that is between

between the Glaffes, comes to the Eye at O, there will appear a black Spot; and the Colours, which, as you recede from the Center, are fo difpofed, that by reafon of the fame Colours coming over again, they may be referr'd to feveral Orders, will be as follows; BLACK, Blue, White, Yellow, Red: VIOLET, Blue, Green, Yellow, Red: PURPLE, Blue, Green, Yellow, Red: GREEN, Red; which Colours are alfo incompafs'd with other Colours; but, as you recede from the Center, grow continually weaker and weaker.

If the Light paffes thro' the Glass to the Eye at O, the transparent Spot which transmits all the Rays is White; and by that Series, as you recede from the Center, there will appear Colours, which are also referr'd to feveral Orders contrary to those above-mention'd: WHITE, yellowish Red, Black Violet, Blue: WHITE, Yellow, Red, Violet, Blue: GREEN, Yellow, Red, bluish Green: RED, bluish Green, which are also incompass'd with weaker Colours.

Experiment 2.] Blow up foap'd Water into a Bubble, to form a thin Plate of Water. Let this be cover'd with a very clear Glass, left, by the Agitation of the Air, the Colours, to be obferv'd in that Bubble, fhou'd be confounded by the Motion of the Water. Such a Bubble, becaufe the Water continually runs down every Way, is very thin at Top, and the Thicknefs in going down is continually increas'd; and for the fame Reafon the Thicknefs of the Whole is continually diminifh'd. Before the Bubble breaks, the Top of it becomes fo thin, as to transfinit the whole Light and appear Black. If in that Cafe the Bubble be observ'd by reflected Light, when

when it is enlighten'd by the Reflexion of a whitifh Sky, and the extraneous Light is intercepted by placing any black Body behind the Bubble; the black Spot above-mention'd will appear, and be encompass'd with the fame colour'd Circles, difpos'd in the fame Order as about the black Spot in the former Experiment. As the Water defcends, the Rings are continually dilated till the Bubble burfts.

When the extreme Circumference of the Bubble appears red by the reflected Rays, if it be look'd at fo as to be feen by the transmitted Rays, it will appear blue; and generally the Colours produc'd by transmitted and reflected Rays, in the fame Manner as in the foregoing Experiment, are opposite to one another.

888 By comparing these Experiments, it follows, that if we increase the Thickness of a very thin Plate, its Colour will be chang'd, and there will be the fame Changes fuccessively, and in the fame Order, whether the Plate be form'd out of a rarer or a denser Mediam. For in the Plate of Air, between the Glasses, and the watery one in the Bubble, whose Thickness increases as it goes farther from the 889 Middle, the Colours will be in the fame Order. Yet in a denser Plate a less Thickness is required than in a rarer, to have it ting'd with the fame Colour.

Experiment 3.] Every thing being difpos'd as in Experiment 1. if you wet the Edges of the Glaffes a little on one Side, the Water will by degrees infinuate itfelf between the Glaffes 3 and there will be obferv'd the fame colour'd Circles as in the Air, neither will their Order be changed, but the Circles will be lefs; when the Water is got as far as the Center, all the Portions of the Circles in the Water will be feparated from the Portions of the Circles in the Air, and be all reduc'd into a lefs Space. The

The Colour of a Plate depends upon its Thicknefs* 890 and Denfity, and not upon the encompassing Mediam. * 887 * 889

Experiment 4.] If you take a Piece of Ifingglafs or Talk, fo thin as to have it colour'd, the Colours will not be chang'd by wetting it; that is, if inflead of Air the Plate is incompafs'd with Water.

The Colour of the fame Plate is fo much the more Sor vivid, as its Denfity differs more from the Denfity of the circumambient Medium. This is prov'd by the foregoing Experiment, in which the Colours of the Plate, when wet, are more languid than those of the fame Plate incompass'd with Air. In the third Experiment also the Colours are lefs vivid than in the second; in both there is a Plate of Water; but in the fecond Experiment it is inclos'd with Air, and in the other with Glass; bur Water and Glass differ less in Denfity than Air and Water.

If Mediums equally differ in Denfity, the Colours 892. will be more vivid, if the denfer be incompafs'd with the rarer : For a very thin Glafs Plate, which is colour'd on account of its Thinnefs, being incompafs'd with Air, the Colours will be more vivid than in Experiment 1. in which a Plate of Air is incompafs'd with Glafs.

A Plate of the fame Density, incompass' d with the 893 fame Medium, will reflect fo much the more Light as it is thinner. Yet if the Thickness be too much diminish'd, it does not reflect the Light. All this is plain from the foregoing Experiments; in the three first, the colour'd Circles which are the least, and which are also the thinness, reflect Light best of any; but in the Center, where the Spot is the thinness of all, there is no fensible Reflexion; as this clearly appears in the fecoud: In the first there is also a very thin Plate of

of Air, which does not reflect Light; for the transparent Spot in the Center is much bigger than those Parts of the Surfaces of the Glasses, which immediately touch by the yielding Inwards of the compress'd Parts.

895 If there are Plates of the fame Medium, whole Thickneffes are in an Arithmetical Progression of the Natural Numbers 1, 2, 3, 4, 5, 6, 7, &c. If the thinness of them all reflects any homogeneous Rays, the fecond will transmit the fame, the third again will reflect them, and the Rays will be alternately reflected and transmitted; that is, the Plates whole Thicknesses in the above-mention'd Progression answer to the odd Numbers 1, 3, 5, 7, &c. reflect the Rays which the others transmit, whole Thicknesses answer to the even Numbers 2, 4, 6, 8, &c.

This Property of the Plates obtains in refpect of any Sort of homogeneous Rays, with this Difference, that different Thickneffes are requir'd for different Colours, as has been faid be-

- * \$87 fore *; the thinneft of all for reflecting Violet, and they must be thickest for reflecting Red; if the Thicknesses are intermediate, the Rays of an intermediate Refrangibility are reflected, that
- 896 is, as the Refrangibility of the Ray increases, the Thickness of the Plate, that reflects it, is diminish'd.

Experiment 5.] Let the Experiment be made in a dark Room to produce the oblong Image of the Sun upon a Paper, fuch as is mention'd in Experiment 1. Chap. 18. Take two Object-Glaffes of long Telescopes (fuch as were mention'd in the first Experiment of this Chapter) let them be prefs'd together, and fo dispos'd, that every fingle Colour of the Image abovemention'd may be succeffively seen in them as in a Looking-glass; that is, that the Glaffes may be

be fucceffively enlighten'd by feveral homogeneous Rays; which may be done, by gently moving about its Axis the Prifm that feparates the Rays to make the oblong Image. The Rings. mention'd in the first Experiment, appear, but more in Number, and only of one Colour ; by Reafon of the Unchangeableness in homogeneous Rays: * In the Interflices of those Rings the * 863 Rays are transmitted, as plainly appears by holding a Paper behind, upon which the transmitted Rays will come; the Rings are leaft of any, when they are violet ; then are fucceffively dilated, confidering the following Colours, quite to the Red. The Rings being of any Colour, if you measure exactly the Diameters of the Circles that may be conceiv'd to be in the Middle of the Breadth of the Rings, the Squares of their Diameters will be to one another as the odd Numbers 1, 3, 5, &c. and measuring in the fame Manner the Diameters of the Circles in the Middle of every one of the Interflices between the Rings, the Squares of their Diameters will be as the even Numbers 2, 4, 6, &c. Now in using fpherical Glaffes, the Thickneffes of the Plate of Air, in the Circle above-mention'd, are as the even and odd Numbers.

DEFINITION.

An homogeneous Colour, in a Plate of any Medi- 897 um, is faid to be of the first Order, if the Plate be the thinnest of all those that reflect that Colour ; in a Plate whole Thickness is triple, it is faid to be of the fecond Order, &c.

A Colour of the first Order is the most vivid of 898 any; and fuccesfively in the following Orders, in the fecond, third, &c. it is lefs and lefs vivid. * 893

When a Plate of Air is inlighten'd with heterogeneous Rays, as that between the Glaffes of

of the Telescopes, or any *Plate* like it of any other Substance, as in *Experiment* 1. and 2. feveral of the Rings seen in the last Experiment are confounded together, and that Colour is seen

- 899 which is made of their Mixture; for the fame Thicknefs of a Plate is often requir'd, for reflecting different Colours of various Orders: So a Plate, which reflects the Violet of the third Order, does also reflect the Red of the fecond Order, as may be deduc'd from the last Experiment, if you attentively confider it: Therefore in the first and second Experiment, the third violet Ring is confounded with the outward Edge of the second red Ring, and purple Colour is produc'd; yet all the Red of the second Order is not absorb'd, because the red Ring is wider than the violet one.
- 900 The more the Thickness of a Plate is increas'd the more Colours it reflects, and different ones, of different Orders. The violet Plate of the tenth Order falls in with the blue one of the ninth Order, and the yellow one of the eighth Order, and lastly with the red one of the feventh Order; and the Colour of the Plate is made up of a Mixture of those Colours.
- 901 If in the first and fecond Experiments a Spe-Etator look obliquely upon the Plates, such as that made of Air, and that which is made of Water, the Rings will be dilated as they are seen more obliquely, that is, in that Motion of the Eye the Colour of the Plate in a determinate Place is chang'd; yet the Dilatation is greater in the first
- SO2 Experiment; which proves, that the Colour is more chang'd by the Obliquity of the Rays, if the Plane be incompass'd with a denser Medium, than if it be inclos'd by a rarer Medium.

Plate XVIII. Fig. 7.] The Demonstration of which Proposition is easily deduc'd from the Laws

Laws of Refraction. Let L and / be thin Plates : the laft incompass'd with a denfer, and the first with a rarer Medium ; let them both be of the fame Thickness : If the Rays A B, ab, equally inclin'd to the Plates, fall upon them, there will be at L a Refraction towards the Perpendicular; * on the contrary, at / the Rays are refracted * 624 from the Perpendicular; * and tho BD and b d * 625 are equal, bc is longer than BC; and therefore there is a greater Change in the Motion of the Light in the Plate I than in the Plate L. The 902 Density of the Plate L being increas'd, the Medium with which it is incompass' dremaining the same there will be a less Difference between BC and B D. and therefore a lefs Change of Colour ; and if the 904 refracting Power of the Plate be so increas'd, that the refracted Rays (whatever be the Obliquity of the incident ones) shall not fensibly differ, there will be no fensible Difference in the Colour of the Plate, whatever Situation the Eye is plac'd in.

Hence we may eafily deduce, that the Colour 905 of some Plates will vary by changing the Position of the Eye, and that the Colour of others is permanent.

CHAP. XVIII.

Concerning the Colours of Natural Bodies.

W HAT relates to the Colours of all Sorts of Bodies may be eafily deduc'd, from what has hitherto been explain'd.

We have fhewn that the Rays of Light have Colours peculiar to themfelves and unchangeable, fo as not to be chang'd by Reflexion. * * 863

Therefore the Rays refletted from Bodies, ac- 906 cording as a greater or lefs Refrangibility is proper to the Colour of the Body it/elf, have a greater or lefs Refrangibility.

Experi-

Experiment 1.] In the Middle of a black Paper fix two fquare Pieces of Ribbon, the one red and the other violet, which muft be join'd fo as to touch one another at their Sides in the fame Manner as the red and violet Colours in the third Experiment of *Chap.* 19. The black Paper muft be fo plac'd, that the Ribbons may be well inlighten'd by the Light that comes into the Room thro' the Window : If a Spectator looks at thefe Ribbons thro' a Prilm, as in the Experiment above-mention'd, the Colours will appear feparated in the fame Manner as in that Experiment.

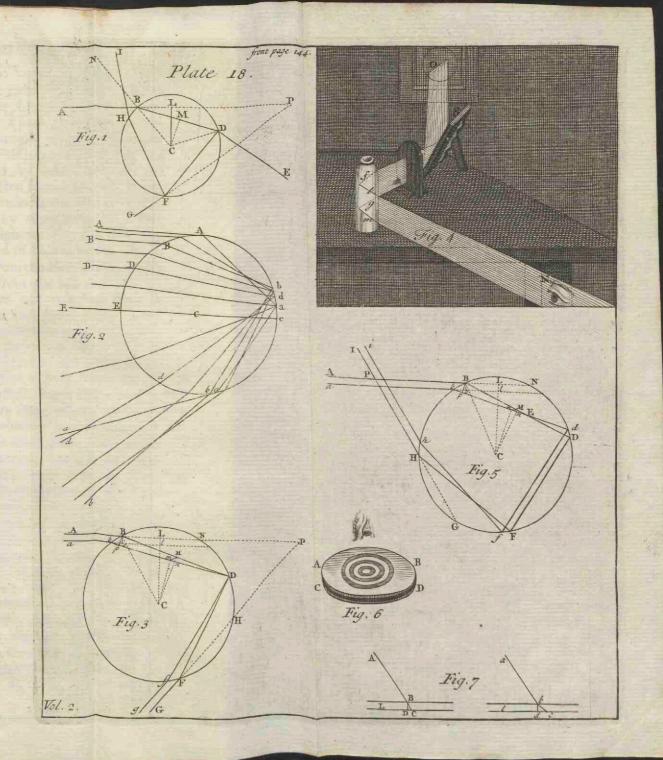
Experiment 2. Plate XIX. Fig. 2.] Place the two Ribbons mention'd in the former Experiment at R and V; let the laft be violet, the first red; let them be inlighten'd by the Flame of a Candle; at fix Feet Diftance place the convex Lens V; (of which Mention has been made in Exper. 6. Cap. 19.) at the Diftance of about fix Feet, you will have the Reprefentation of the Ribbon R upon a white Paper at r; at a lefs Diftance you will have an exact Reprefentation of the other at v. You may determine when the Reprefentations are exact, by binding black Threads upon the Surface of the Ribbons; for thefe Threads appear diffinctly, where the Reprefentation is exact.

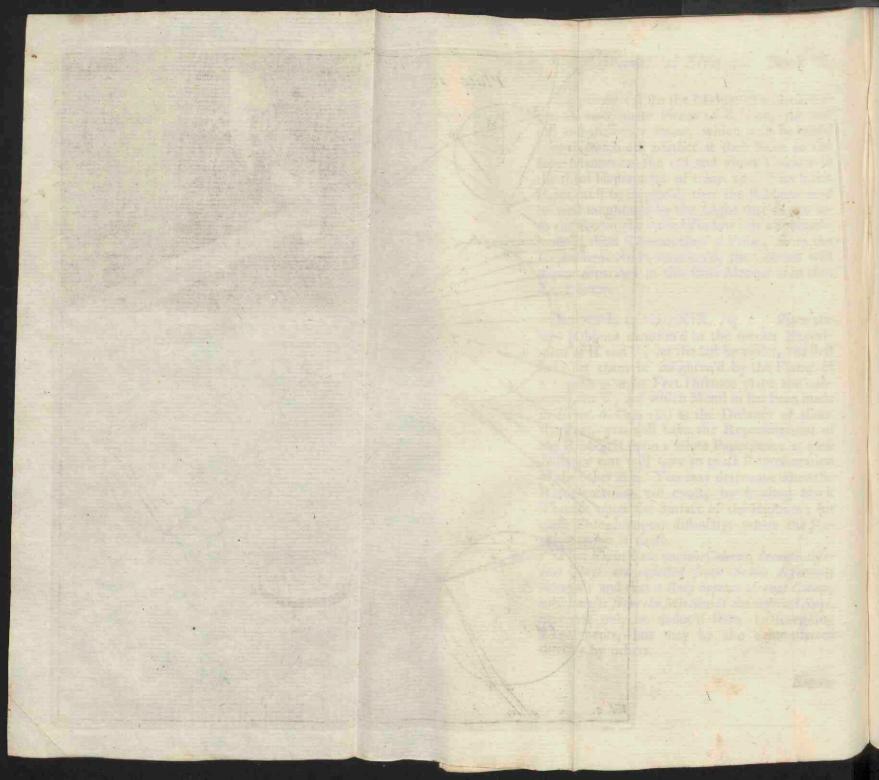
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That Bodies have various Colours, because different Rays are reflected from Bodies differently colour'd; and that a Body appears of that Colour, which arises from the Mixture of the reflected Rays, may not only be deduc'd from the foregoing Experiments, but may be also demonstrated directly by others.

Experi-





Experiment 3.] Take two Bodies of any Kind, the one red, the other blue; let them be enlightened fucceffively, in a dark Place, by the Colours of the coloured Image, which is made by the Refraction of the Prifm; every one of the Colours will be reflected by both; but the red Rays are copioufly reflected from the red Body, whilft the blue Body reflects but few of them, which plainly appears by comparing both Bodies, when they are enlightened by the red; the contrary may be observed in blue Colours, which are copioufly reflected from the blue Body, whilft only a few are reflected from the red one.

The Rays, which are not reflected from a Body, penetrate into it, and there fuffer innumerable Reflexions and Refractions, as we have explained between N° 842, and 843; till at length they unite themfelves to the Particles of the Body itfelf.* Therefore a Body grows bot fo *546 much the fooner, as it reflects Light lefs copiou/ly, 908 For which Reafon a white Body, which reflects al-909 moft all the Rays with which it is enlightened, * *870 heats the floweft, which a black Body, into which almoft all the Rays penetrate, because only few are reflected, * acquires Heat fooner than any other. *845

To determine that Conftitution of the Surfaces of Bodies, upon which their Colour depends, we muft take Notice of the fmalleft Particles, of which thefe Surfaces are made up: Thefe Particles are transparent *, and are fepa-*g41 rated by a Medium of different Denfity from the Particles themfelves; they are also thin, *843 otherwife the Surface would as it were be covered by a transparent Body, * and the Colour *843 would depend upon the Particles under thefe. Therefore in the Surface of every coloured Body there are innumerable fmall thin Plates; but by Vol. II. L leffen-

leffening the Plate, keeping the fame Thicknefs, its Properties, as to the Reflexion of Light, are not changed; for the leaft Plate, in refpect of the Rays of Light, is very large: Wherefore, what has been demonstrated in the preceding *Chap*. may be applied to these Plates in the Surfaces of Bodies. From whence we deduce the following Conclusions.

910 The Colour of a Body depends upon the Thickness *890, and Density of the Pants of the Body, which are in 911 the Surface, between the Pores of the Body.*

911 the Surface, between the Pores of the Body.* * 898, The Colour is fo much the more vivid and homo-899, geneous, as the Parts are thinner.*

912 Cæteris paribus, the aforefaid Parts are of the *896, greatest Thickness when the Body is red; and of the 861 least, when violet.*

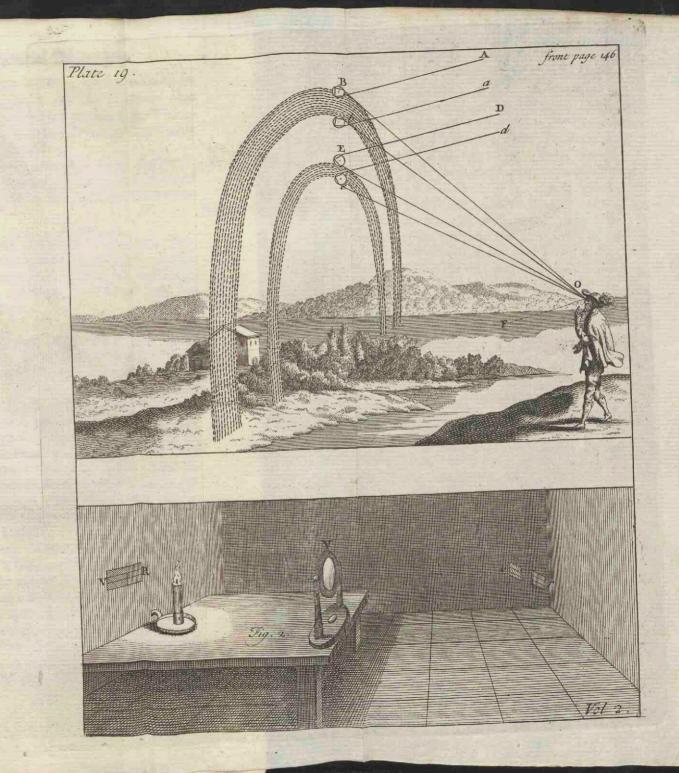
913 The Parts of Bodies are much denfer than the *902, Medium contained in their Interflices.**

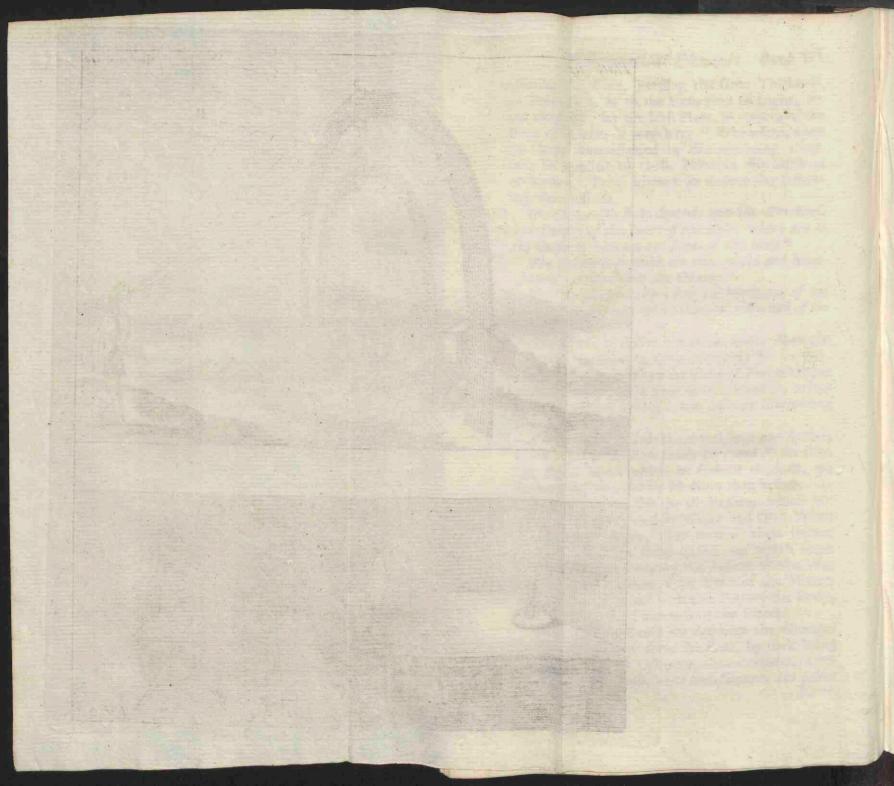
904. This Denfity is lefs in the Tails of Peacocks, and 914 in fome Silks, and in general in all Bodies, whofe 901 Calour varies according to the different Situation of 923 the Eye.* 915 The Colour of a Deduction of Compared Lower

⁹¹⁵ The Colour of a Body is more obfcure and darker, when a denfer Medium enters its Pores;* for then the Parts, upon which its Colour depends, are furrounded by a denfer Medium than before.

We experience this in all Bodies, which are thoroughly penetrated by Water and Oil : When the Bodies are dry, they recover their former Colour, unlefs in fome Cafes, in which fome Parts are carried away by the Action of the Water, or Oil, or when fome Parts of the Water, or Oil, are fo united with the Parts of the Body, as to change the Thicknefs of the Plates.

From fuch a Caufe are deduced the Changes in the Colours of fome Liquors, by their being 916 mixed with others. Often the faline Particles, fwimming in one Liquid, unite themfelves to the faline Parti-





Particles finimizing in another; or the united Particles are separated by the Actions of others that are superadded; on all which Accounts the Thickness of the Particles is changed, and together with it the Colour of the Liquids.* *888

Sometimes the Colour of a Liquid is different, 917 when seen by reflected Rays, from what it is, when seen by transmitted ones: We have shewn before whence this arifes.*

Experiment 4.] An Infusion of Lignum Nephriticum, that is not too much tinged, appears blue by reflected Rays, but yellow, if the Phial, which contains the Infusion, be placed between the Light and the Eye.

Experiment 5.] If you pour Spirit of Vinegar into the Infufion of Lignum Nephriticum, it will appear yellow in any Polition whatever.

In this Cafe the Thicknels of the Particles is changed, and the Rays, that were transmitted through every one of them, are now intercepted; but though the Liquor is placed between the Eye and the Light, it is feen by reflected Rays; for we may eafily conceive that fuch Rays come to the Eye by the various Reflexions which the Light undergoes in the Liquid. But this Colour only is fensible, because the Rays cannot penetrate directly through the Liquid.

From this we may deduce the Reafon, why a 918 coloured Liquid in a Glafs, of the Figure of an inverted Cone, if it be placed between the Eye and the Light, appears of a different Colour, in different Parts of the Veffel; in the lower Part all the Rays which are transmitted thro'the Particles, are not intercepted; then they are more and more intercepted, according as there is a greater Quantity of the Liquid between the Eye and the Light; till at Vol. II. L2 length

length they come to be all intercepted; and only those Rays which are reflected by the Particles penetrate the Liquid; in which Cafe the Colour coincides with the Colour of the Liquid seen by the reflected Rays.

919 Clouds often appear very beautifully coloured; they confift of aqueous Particles, between which Air is interfperfed; therefore, according to the va-920 rious Thicknefs of those aqueous Particles, the Cloud "388 will be of a different Colour.*

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The End of the Third Book.

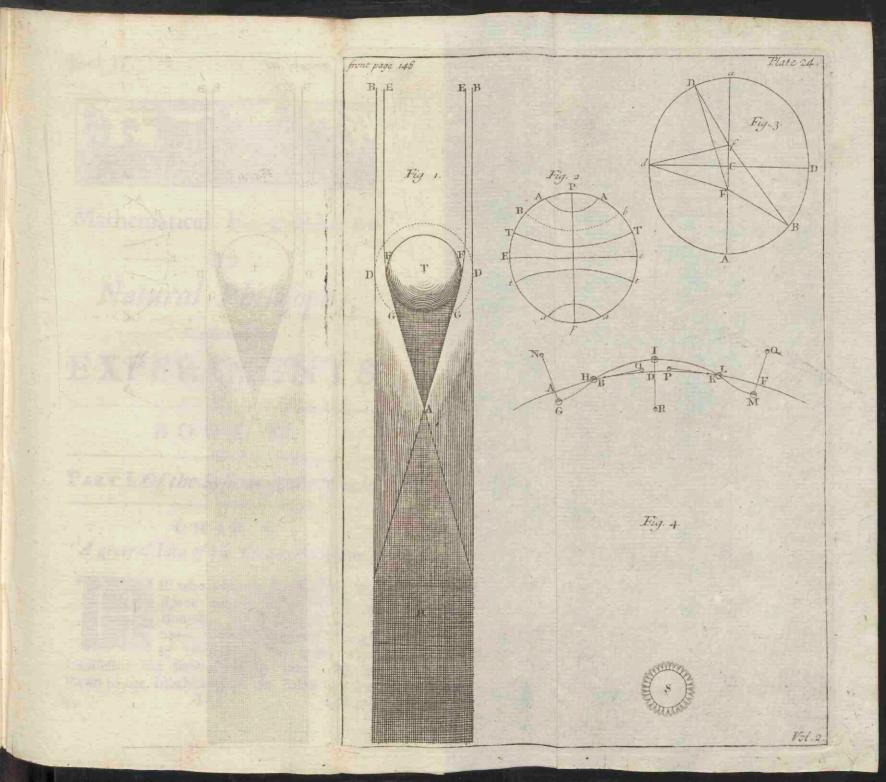
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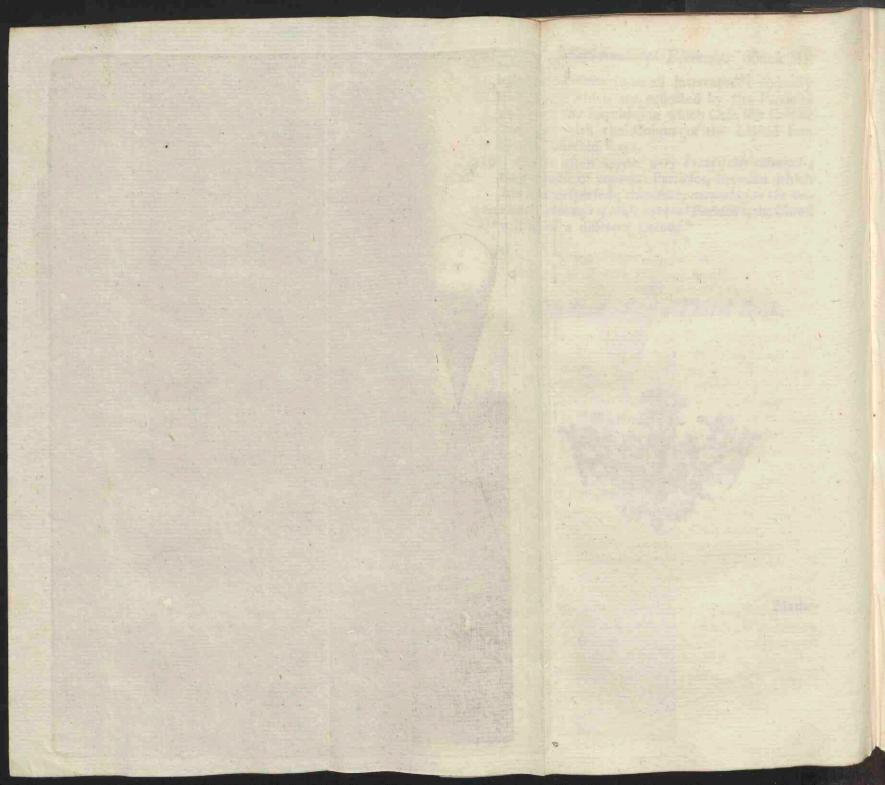
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Mathematical ELEMENTS

OF

Natural Philosophy,

Confirmed by

EXPERIMENTS

BOOK IV.

PART I. Of the System of the World.

CHAP. I.

A general Idea of the Planetary System.



E who attentively confiders that *15 Space can be terminated by no Bounds, will fcarce be able to deny, that the Supreme Almighty Intelligence has every where manifested the fame Wildom which He has shewn to the Inhabitants of the Earth in a small L 3 Compafs.

Compass. What I here call a finall Compass, immensity exceeds our Comprehension; yet it is but finall, compared with infinite Space.

921 Our Earth, with fixteen other Bedies (for we have no Knowledge of any more (moves in a determinate Space; neither do thefe Bodies recede from, or come nearer to one another, beyond their fet Bounds; and their Motions are performed according to unchangeable Laws.

DEFINITION I.

922 This Collection of Seventeen Bodies is called the Planetary System.

The whole Art of Aftronomy is almost employed concerning thefe alone; and thefe will be chiefly my Subject in this Work : The other Bodies that conflitute the Univerfe, are too far diftant from us, for their Motions (if they are moved) to fall under our Obfervations : Of thefe only the lucid Bodies can be perceived by us; and of thofe only the more remarkable ones, and which are lefs diftant from us than the reft; even most of fuch as are feen by the Telefcope are invisible to the naked Eye.

DEFINITION II.

923 All these Bodies are called fixed Stars.

They are called fixed, because, as far as can be perceived, they keep the fame Position, with respect to one another : We must take Notice of fomething peculiar concerning these hereafter.

924 But as to the Planetary System : In this we have faid there are feventeen Bodies, which are all spherical : One only shines by its own Light, the rest are opaque, and are visible only by the Light which they borrow from that.

825 The Sun is that lucid Body, and far the greatest of all in the Planetary System; it is quiescent in the middle Book IV. of Natural Philosophy. 151 dle of it, at least is agitated by a very small Motion.

DEFINITION III.

The other fixteen are called Planets.

Thefe are divided into two Claffes; fix are called 926 the Primary Planets; ten are called the Secondary Planets. When we fpeak of the Planets, without any Diffinction, we always underftand the Primary ones.

The Primary Planets move round the Sun, and are 927 carried at different Diffances from it, in Curves that return into themfelves.

A Secondary Planet revolves round a Primary 928 one, and accompanies it in its Motion round the Sun.

The Planets in their Motions describe Ellip-929 tic Lines, not much different from Circles. And all those Lines are fixed; at least, there is but a fmall Change to be observed in a long Time, in their Situation.

Plate XXIV. Fig. 3.] An Elliptic Line is 930 formed, if a Thread, whole Extremities are fixed in the two Points F and f; is moved, remaining firetched; as is to be feen in this Figure, in which the Thread is reprefented at F df, F E f, F B f. The Points F, f, are called the Foci; the Line A a, which paffes through them, and is terminated on each Side by the Circumference of the Ellipfe, is called the greater Axis; and it is the greateft Line that can be drawn in the Ellipfis, The middle Point C of the Axis is the Center of the Ellipfe ; and the leffer Axis D d falls perpendicularly through this Point upon the greater Axis.

The Orbits of all the Primary Planets are in 931 fuch a Position, that one of their Foci falls in with the Center of the Sun; let the Ellipse A D ad L 4 represent

reprefent the Orbit of a Planet, F will be the Center of the Sun.

DEFINITION IV.

- 932 The Diftance between the Center of the Sun and the Center of the Orbit, is called the Eccentricity of the Planet: as F C.
- 933 In every Revolution the Planet approaches once to the Sun, and once recedes from it; and at its greatest Diftance is at a, the Extremity of the greater Axis of the Orbit, and at its least Diftance, in the opposite Extremity A.

DEFINITION V.

934 That Diftance of the Planet from the Sun, is called the mean Diftance, which is equally different from the greatest and the least. At this Diftance the Planet is in the Extremities D, d, of the smaller Axis.

DEFINITION VI.

935 The Point of the Orbit, in which the Planet is at its greatest Distance from the Sun, is called the Aphelium; as a.

DEFINITION VII.

936 The Point of the Orbit, in which the Planetis at its leaft Diftance from the Sun, is called the Perihelium.

DEFINITION VIII.

937 These Points are commonly called the Auges or Apfides.

DEFINITION IX.

938 The Line which joins the Apfides, that is, the greater Axis of the Orbit, is called, the Line of the Apfides.

Every

Every Orbit is in a Plane which paffes through 939 the Center of the Sun.

DEFINITION X.

The Plane of the Orbit of the Earth is called 940 the Plane of the Ecliptic.

This Plane is to be continued every Way; and Aftronomers confider the Position of the Planes of the other Orbits, with respect to this.

DEFINITION XI.

The Points, in which the other Orbits cut the 941 Plane of the Ecliptic, are called the Nodes.

DEFINITION XII.

The Line which joins the Nodes of any Orbit, 942 that is, the common Section of the Plane of the Orbit with the Plane of the Ecliptic, is called the Line of Nodes.

A Planet is not carried with an equal Celerity in 943 all the Points of its Orbit; the less it is distant from 944 the Sun, the swifter is its Motion; and the Times, in which the several Arcs of its Orbit are run thro', are to one another, as the Area's formed by Lines drawn from the Planet to the Center of the Sun: The Arcs A B and a E are run through in Times, which are to one another, as the Area's of the mixed Triangles A F B, a F E.

All the Planets are carried the fame Way, and 945 their Motion in their Orbits is contrary to that Motion which we observe daily in all the Cœleftial Bodies, by which in one Day they feem to be carried round the Earth; of which hereafter.

DEFINITION XIII.

A Motion, fuch as is that of the Planets in their 946 Orbits, is faid to be in Confequentia, and direct.

DEFI-

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DEFINITION XIV.

947 The contrary Motion is called a Motion in Antecedentia; and fometimes Retrograde.

948 The more diftant the Planets are from the Sun, the flower they move in their Orbits; fo that the periodical Times of the most diftant ones are greater, both because they have a greater Orbit to run through, and a flower Motion.

DEFINITION XV.

- 949 The Line which passes through the Center of the Planet, and about which it moves, is called the Axis of the Planet.
 - 950 The Planets, at leaft most of them, and the Sum itfelf move round their Axis: There are Two, of which, in this refpect, Astronomers have been able to make no Observations, but which, in all Probability, have this Motion.
 - 951 This Motion agrees or confpires with the Motions of the Planets in their Orbits, that is, it is in Confequentia.
 - 952 The Axes themfelves are moved by a parallel Motion, fo that all the Points of the Axis of a Planet defcribe equal and fimilar Lines.

DEFINITION XVI.

- 953 The Extremities of the Axis are called the Poles of the Planet.
- 954 Plate XX. Fig. 1.] Aftronomers compare together accurately enough the Diftances of the Planets from the Sun, to give us an Idea of the whole Syftem. The Dimensions of the Orbits are represented in this Scheme, in which the Points N N shew the Nodes of each Orbit.

955 Neverthelefs we cannot compare the Dimensions of this System with any that we know upon the Surface of the Earth; for no Astronomers will affert,

that

that the Obfervations made concerning fuch a Comparison, are free from Error.

But, that the feveral Parts of the Syftem may 956 be compared together, we fuppole the mean Diftance of the Earth from the Sun divided into 1000 equal Parts, which we make use of in meafuring the other Diftances.

The Sun , as was faid before, is agitated by 957 a fmall Motion in the middle of the System; it moves round its Axis in the Space of 25 Days: And its Axis is inclined to the Plane of the Ecliptic, making an Angle of 87 Degr. 30 Min.

Mercury rightary is the leaft diftant from the Sun of 958any of the Planets; its mean Diftance from theSun is 387, its Eccentricity 80, the Inclinationof its Orbit, that is, the Angle, formed by thePlane of its Orbit with the Plane of the Ecliptic,is 6 Degr. 52 Min. it performs its Revolutionround the Sun in 87 Days, 23 Hours.

The next is Venus 2, whole Diftance from the 959 Sun is 723, its Eccentricity 5, the Inclination of its Orbit 3 Deg. 23 Min. It performs its Periodical Motion in 224 Days, 17 Hours; and its Motion round its Axis is performed in 23 Hours.

The third Planet in order from the Sun is 960 our Earth ⊕; its mean Diftance from the Sun is 1000, its Eccentricity 169: It is moved in the Plane of the Ecliptic; its Periodical Time is 365 Days, 5 Hours, 51 Min. and the Motion about its Axis is performed in 23 Hours, 56 Min. 4 Sec. Its Axis makes an Angle with the Plane of the Ecliptic of 66 Degr. 31 Min.

The mean Diftance of Mars & from the Sun 961 is 1524, its Eccentricity 141, the Inclination of its Orbit I Degr. 52 Min. Its Periodical Time 686 Days, 23 Hours. Its Revolution about

about its Axis is performed in 24 Hours, 40 Min.

962 Jupiter 4, the biggest of all the Planets, is diftant from the Sun, at a mean Diftance 5201, its Eccentricity 250. The Inclination of its Orbit 1 Degr. 20 Min. The Periodical Time 4332 Days, 12 Hours ; and its Revolution about its Axis in 9 Hours and 56 Min.

963 The mean Diftance of Saturn h, the most diftant Planet from the Sun, is 9538; its Eccentricity 547; the Inclination of its Orbit 2 Degrees 30 Min. The Periodical Time 10759 Days, 7 Hours. It is encompaffed with a Ring which does not touch the Planet, but never leaves it. This Ring is not visible without a Telescope.

The mean Diftance being given, if you add the Eccentricity, you will have the greatest Diftance; but if you fubftract the Eccentricity from the mean Diftance, you will have the leaft *836 Diftance.*

- 964 The three Planets, Mars, Jupiter, and Saturn, which are more diftant from the Sun than the Earth, are called the Superior Planets; Venus and Mercury are called the Inferior ones.
- 965 Of the Primary Planets, three are accompanied by Secondary ones.

Five Planets called Satellites, move about Saturn ; four about Jupiter ; one about the Earth (viz.) the Moon.

The fecondary Planets, except the Moon, are not visible to the naked Eye.

966 The Satellites, by Lines drawn to the Center of the Primary Planets, describe Area's about them preportional to the Times; as has been faid of the Primary Planets, with refpect to the Center of the *044 Sun.*

The Moon moves about the Earth in an Ellipfe, 967 one of whole Foci is in the Center of the Earth, 968, from which the mean Diftance of the Moon is 60 969 .Semidiameters of the Earth and a Half: Its Eccentricity is liable to Change ; the mean one is of 3 Semidiameters and a Third. The Plane of its Orbit forms an Angle with the Plane of the Ecliptic of 5 Degr. but this Inclination is not con- 970 Stant, or always the fame. In the Motion of the Moon round the Earth, neither the Line of the Nodes, nor the Line of the Apfides, is carried in a parallel Motion ; but the Line of the Nodes Weftward, or in Antecedentia; the Lines of the Apfides Eaftward, or in Confequentia ; the first performsits Revolution in about 9 Years, the fecond in 19. The Periodical Time of the Moon's Motion about the Earth is 27 Days and about 7 Hours ; and it is turned about its own Axis exactly in the fame Time.

Plate XX. Fig. 2.] The first or inmost of the 971 Satellites of Jupiter is diftant from Jupiter's Center 25 Diameters of Jupiter : It is moved round Jupiter in one Day, 18 Hours, 28 Min.

The Diftance of the Second is 4 and an Half Diameters of Jupiter : Its Periodical Time is 13 Hours, 18 Min.

The Diftance of the Third is 75 Diameters; its Periodical Time 7 Days, 4 Hours.

The Fourth is diftant 123 Diameters: It performs its Motion in 16 Days, 18 Hours, 5 Min.

Plate XX. Fig. 3.] The first or inmost Satel-972 lite of Saturn is diftant from Saturn's Center 30 of a Diameter of the Ring: Its Periodical Time is I Day, 21 Hours, 18 Min.

The Diftance of the Second is 11 Diameter of the Ring : Its Periodical Time is 2 Days, 17 Hours, 41 Min. The

The Diftance of the Third is 13 Diameter of the Ring; its Periodical Motion is 4 Days, 13 Hours, 47 Min.

The Diftance of the Fourth is 4 Diameters of the Ring; its Periodical Time 15 Days, 22 Hours, 41 Min.

The Diffance of the Fifth is 12 Diameters of the Ring; its Periodical Time is 79 Days, 7 Hours, 53 Min.

973 Concerning the Motion of thefe, as alfo of the Satellites of *Jupiter* about their Axes, we can hitherto determine nothing certain from Aftronomical Obfervations.

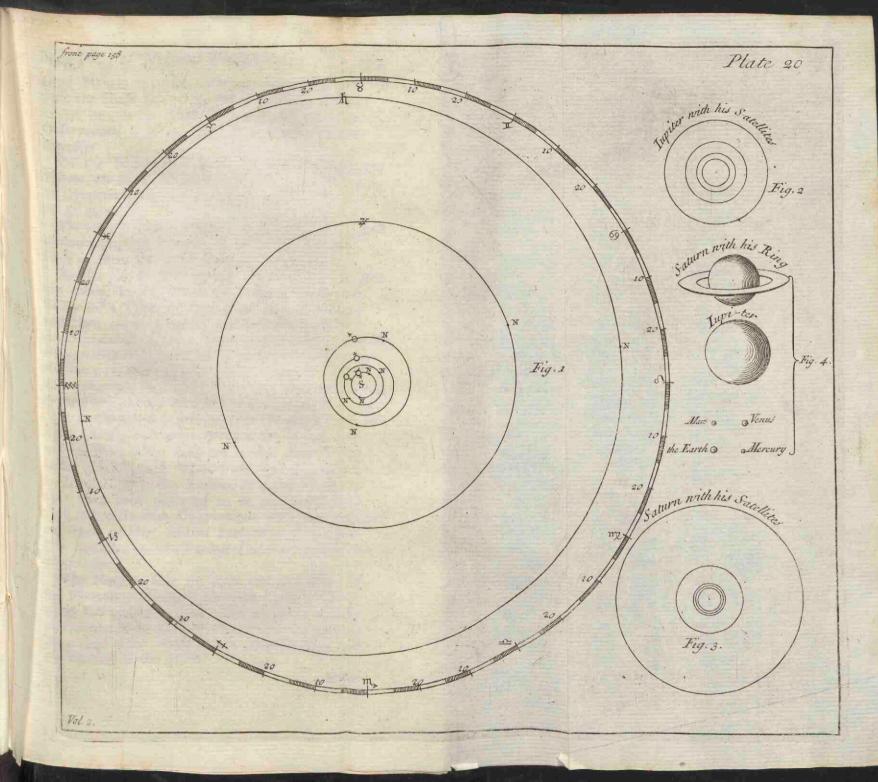
If we take Notice of the Diftances and Periodical Times of the Planets, we shall find that the following Rule holds good in our System, whereever feveral Bodies are moved round the fame Point, that is *about the* Sun, Saturn, and Jupiter;

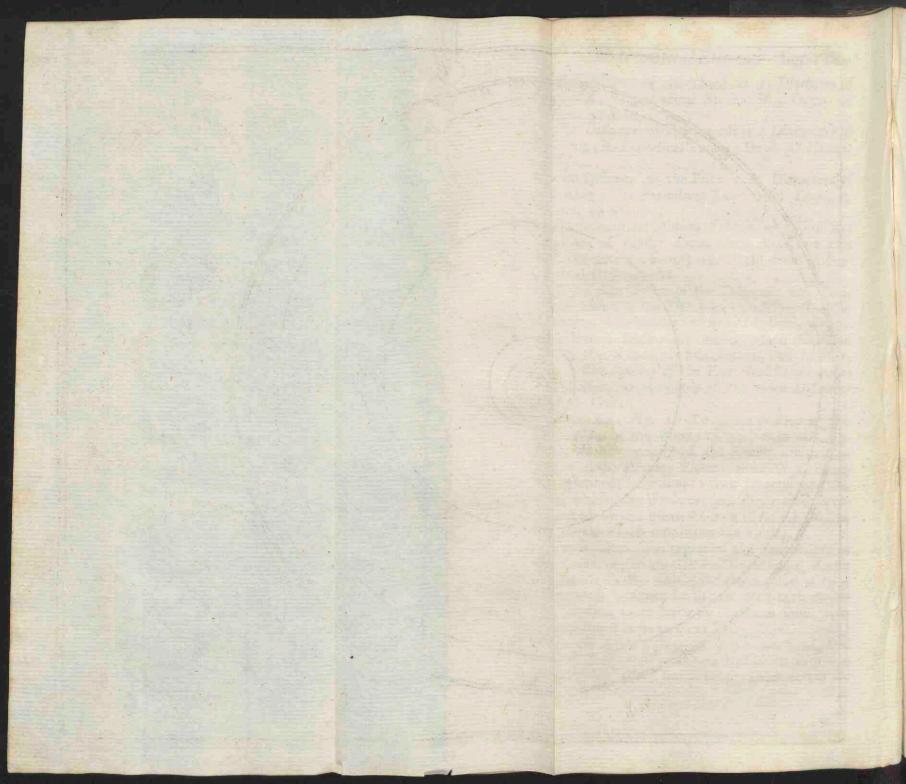
- 974 (viz.) The Squares of the Periodical Times are to one another, as the Cubes of the mean Diffances from the Center.
- 975 Plate XX. Fig. 4.] To give a Notion of the Dimensions of the Bodies themselves in our System, we have contrived the fourth Figure, in which all the Primary Planets, and Saturn's Ring, are defcribed, according to their Dimensions; the Sun, whose Magnitude exceeds all the rest, is represented by the greatest Circle (Fig. 1.) that is, the Circle which terminates the Figure.

These Dimensions represent the Proportions of the Bodies one to another exactly enough, if you except the Earth, which, for the Reason already *955 mentioned, * cannot be so compared with the o-

ther Bodies, as to leave no Room to doubt what Proportion it bears to them.

976 Yet the Earth's Diameter may be measured, and contains 3.400669Perches, each of which is equal to 12 Rhynland Feet; but altho' the Diameters of the other





other Planets may be compared together, and with the Sun's Diameter, yet it can't be determined how many Feet they contain, till after Obfervations shall be made at a proper Time hereafter.

Of the Bodies that make up the Planetary 977 Syftem, the Moon only can be compared to the Earth; its Diameter being to the Diameter of the Moon, as 40 to 11.

The other fecondary Planets are not meafured by 978 Aftronomers, but it can't be doubted, but that fome of them are bigger than the Earth.

Befides the Bodies already mentioned, there 979 are others in the Planetary Syftem, which are vifible for a Time, as they come near the Sun, and then recede from it, and become invifible; they are called Comets. They appear most com-980 monly with Tails, and the Tail is always turned from the Sun; in their Motion they deferibe Area's, by Lines drawn to the Center of the Sun, proportional *944 to the Times, as has been faid of the Planets.* 966

As to Comets, it is probable that they move in el. 981 liptic Orbits, that are very eccentric; fo that they are invifible, when they are in that Part of the Orbit which is most diftant from the Sun; which is deduced from the Periods of fome of them, that have been observed to be pretty regular: And it is plain from Observations, that fome bave 982 defcribed in their Motion Portions of very eccentric Eclipfes, in one of whose Foci was the Center of the Sun.

The Notion, that we have hitherto given of the Planetary Syftem, is founded upon Aftronomical Obfervations; and what we have already faid admits of no Difpute among Aftronomers, if we except what relates to the Elliptic Line and the Motion of the Earth.

Some

Some affirm, that the Orbits of the Planet's are not Elliptic, but that in their Motion they defcribe another Oval : *Kepler* has deduced from *Tycho Brabé*'s Obfervations, that thefe Lines are Elliptic ; and we fhall fhew in the following Part, that no other Curves can be defcribed by the Planets.

Those that fay the Earth is at reft, have no Aftronomical or Physical Argument for a Foundation of their Opinion; that is, don't reafon from *Pbænomena*: Neglecting the Simplicity of the System, and the Analogy of the Motions, they affert, that their Opinion is not contrary to Observations; in which they err, as we shall shew in the following Part.

CHAP. II.

Concerning the Apparent Motion.

WHOEVER, after having read the former Chapter, looks upon the Heavens, will fcarcely believe, that he beholds the Syftem which is explained there; and a more exact Confideration of the heavenly Motions will en-983 creafe his Doubt. No Wonder, fince we can obferve very little in the Heavens but false Appearances.

The common Observer of the Heavens is a Spectator, who, thinking himself to be at reft, is carried about by various Motions, and beholds Bodies, concerning whose Distance and Magnitude he makes false Judgments. The true System of the World was unknown for many Ages, even to the most exact Observers of the Heavens.

But

But we must explain how all Things, which 984 are observed in the heavenly Bodies, have Place in the System that has been explained, in respect of a Spectator upon the Earth; that is, we shall deduce the Appearances from the true Motions. Which cannot be done, unless we first lay down fome general Things concerning the apparent Motion in general.

It is certain, that we have no Art, by which we can difcover the true Motion, only the relative Motion can be perceived by the Senfes; and it is that only concerning which we treated in the former Chapter. Who can reafonably affirm or deny, that all the Bodies, which are known to us, are not carried in a common Motion through the immenfe Spaces ?

The relative Motion to be diftinguished from the 985 apparent one; for the apparent Motion is the Change which appears to be in the Situation of the Bodies, and depends upon the Change of the Picture in the Bottom of the Eye; for Objects have the fame apparent Relation to one another, as their Reprefentations have in the Eye; for they are feen as they are painted in the Eye; * *716 and the Change in that Picture, from the Motion of the Bodies, most commonly differs from the Change of the Relation between the Bodies themfelves; as follows from the Formation of that Picture.

The Heavens are nothing but an immense Space, 986 which cannot be seen, and would appear black, * *845 if innumerable Rays of Light, flowing from the heavenly Bodies, did not continually penetrate our Atmosphere. Most of them come to us from the Bodies in right Lines, yet a great many fuffer various Reflexions in the Atmosphere, and enlighten the whole Atmosphere; which is the Reason that, in the Day, Bodies are enlight-Vol. II. M ened,

ened, even without the Reflexion of the Clouds, to which the Rays of the Sun cannot come directly.

These Rays are heterogeneous and white; for there are Bodies enlightened by these Rays which appear white: And these Bodies seen through Prisms, at the Extremities, are tinged with Colours, which does not happen in an homogeneous *874 Colour; * also a Circle of white Paper of Half an Inch Diameter, being put upon black Cloth, if it is enlightened by these Rays, it will appear ob-*861 long through the Prism; * and the same Colours, which are observed in the Rays of the Sun, are feen here in the same Manner : All which Things would not happen, if the Air, as many think, was a blue Liquid; that is, through which only the Sun's blue Rays, at least mostly such, pass.

987 When we look at the black Sky, the white Rays, beforementioned, enter our Eyes, whence the blue Colour of the Sky arifes. Becaufe we are accuftomed to fee a Colour where there is a coloured Object, we also refer the Colour of the Heavens to an Object; but fince this is feen equally towards all

988 Parts, we conceive a concave spherical Surface, in whose Center we are placed; we imagine this Surface to be opaque, and therefore distant from us beyond all visible Bodies.

When a Body is between a Plane and the Eye, of whofe Diftance we cannot judge, the Body appears to us to be applied to the Plane, whatfoever the Diftance is between that and the Plane; for there is no Reafon why the Parts of the Plane, which are painted at the Sides of the Image of the Body in the Eye, fhould not appear at the fame Diftance with the Body.

989 Thence also all the celestial Bodies (of which the least distant from us, (viz.) the Moon) is yet fo removed that we can give no Judgment of

162 .

of its Diftance *, are referred to that imaginary *732 Sphere above-mentioned; and they all appear equally diftant, and jeem to move in the Surface of a concave Sphere. So the Moon appears to be amongft the fixed Stars, although its Diftance bears fearcely any fensible Proportion to the Diftance of Saturn; and the Diftance even of Saturn itself is nothing, compared with the immense Diftance of the fixed Stars. It is no Wonder therefore, that the common People know nothing of the Magnitude of the celeftial Bodies, and the Immensity of the Heavens.

We fee from what has been faid, how the Motion of any Body being given, and the Motion of the Earth being known, the apparent Motion may be determined.

We have faid that a Sphere is imagined beyond the fixed Stars, in whofe Center is the Spectator: * The Orbit of the Earth is fo fmall, *988 in refpect of the Diameter of this Sphere, that the Center of the Sphere is not fenfibly changed by the Alteration of the Place of the Spectator, whilf the is carried along with the Earth. Wherefore in all the Points of the Earth's Surface, and at 990 any Time, the Inbabitants of the Earth imagine the fame Sphere, to which they refer the beavenly Bodies; and which hereafter we fhall call the Sphere of the fixed Stars.

These Things being laid down, if we conceive 991 a Line to be drawn through the Earth, and a Body, which, being continued beyond the Body, cuts the afore faid Sphere; we have a Point, to which the above-mentioned Body is referred, and which is the apparent Place of that Body.

Whilft the Body, or the Earth, or both are moved, this Line is moved alfo, and the apparent 992 Motion is the Line, which is described among fi the fixed Stars by the Extremity of the Line above-Vol. II. M2 mentioned.

mentioned, which goes through the Earth and the Body, whole apparent Motion is observed.

993 Therefore the same Appearances will follow from the Earth's being moved out of its Place, as if the Body had been moved, and the same also may be deduced from the Motion of both.

994 But if the Body and the Earth be fo moved, that the Line which passes through these Bodies be carried in a parallel Motion, the Body will seem to be at rest amongst the fixed Stars; because in this Case, the Space, gone through by the End of the Line amongst the fixed Stars, cannot exceed the Space gone through by the Earth; but the Line that is equal to the whole Space which the Earth can go through, at so great a Distance as the fixed Stars, is not fensible to us.

995 From the Motion of the Earth round its Axis there is also produced an apparent Motion, which will be eafily deduced in its proper Place from the Foundation laid down in this Chapter.

That the apparent Motion differs from the relative, and is varied by the Motion of the Spectator, is what Sailors every Day observe.

CHAP. III.

Of the Phanomena or Appearances of the Sun from the Motion of the Earth in its Orbit.

the apparent Motion of the Sun depends upon the Celerity of the angular Motion of the Earth, with refpect to the Center of the Sun; which Motion encreafes upon a double Account; on account of the Diftance from the Sun being leffened, and the Celerity of the Earth being encreafed: Both which Caufes always concur; **944 wherefore the Inequality of the apparent Motion of 997 the Sun is fenfible. In a whole Revolution of the 998, Earth, the Sun alfo feems to run thro' a whole Circle.

DEFINITION I.

This apparent Way of the Sun is called the Eclip-999 tic Line. It is the Section of the Sphere of the fixed Stars with the Plane of the Ecliptic, fupposed to be continued to this Sphere.

This Way is divided into 12 equal Parts, each of which contains 30 Degr. Thefe Parts are called the Signs, and are diftinguished by these Names; Aries γ , Taurus O, Gemini II, Cancer S, Leo St, Virgo W, Libra S, Scorpins M, Sagittarius S, Capricorn VS, Aquarius S, Pisces H. Whence these Parts have their Names, we shall explain when we treat of the fixed Stars.

The Sun is longer a going thro' the fix first Signs 1000 than the fix last, and the Difference is nine Days.

Although a Circle has neither Beginning nor 1001 End, yet when feveral Points muft be determined in it, fome Point muft be taken as the Beginning; this, in the Ecliptic Line, is the first Point of Aries; we shall shew how it may be determined hereafter. It is not fixed to one Place amongst the fixed Stars; therefore the Orbits of the 1002 Planets, which alter so little, that they may be looked upon as unchangeable, * don't preferve the *919 fame Situation, in respect of this Point.

DEFI-

DEFINITION II.

1003 The Distance of the Sun from the first Point of Aries, measured in confequentia, is called the Sun's Longitude.

The Longitudes of the other beavenly Bodies are meafured after the fame Manner in the Ecliptic;
1005 they are referred to this Line, by conceiving a great Circle to pafs through the Body, and cut the Ecliptic perpendicularly; for the Point, in which the Ecliptic is cut by this Circle, determines the Longitude of the Body.

DEFINITION III.

1006 The Diftance of a heavenly Body from the Ecliptic is called its Latitude. It is measured by an Arc of a great Circle, perpendicular to the Ecliptic, intercepted between the Body and the Ecliptic.

DEFINITION IV.

1007 If we imagine a Line to go thro' the Center of the Sphere of the fixed Stars, and perpendicular to the Ecliptic, the Points, in which this cuts the abovementioned Sphere, are called the Poles of the Ecliptic.

DEFINITION V.

1008 The Zodiac is a Zone which is imagined in the Heavens, which the Ecliptic Line cuts into two equal Parts, and which, on either Side, is terminated by a Circle parallel to the Ecliptic Line, and eight Degrees diftant from it. On Account of the finall Inclinations of the Orbits of the Planets, and the Moon to the Plane of the Ecliptic, no Bodies of 1009 the Planetary Syftem appear without the Zodiac.

DEFINITION VI.

Those of them that have the same Longitude are 1010 faid to be in Conjunction.

DEFINITION VII.

Those whose Longitudes differ 180 Degr. are faid 1011 to be in Opposition.

CHAP. IV.

Of the Phænomena of the Inferior Planets, arifing from the Earth's and their own Motions in their Orbits.

Plate XXI. LET S be the Sun, A V B v the Fig. 2.] U Orbit of an inferior Planet; let T be the Earth in its Orbit, and a v b Part of the Sphere of the fixed Stars; the apparent Place of the Sun is v.*

If from the Earth there be drawn to the Orbit of the Planet the Tangents T A a, T B b, it is evident that the Planet, in its apparent Motion, is never removed farther from the Sun, than the Diftance v a, v b; and that the Planet accompanies it in its apparent Motion round the Earth.

DEFINITION I.

The apparent Diftance of the Planet from the 1012 Sun is called its Elongation; v a or v b is the greateft Elongation: This varies upon two Ac-1013 counts; (viz.) becaufe the Earth and the Planet revolve in elliptic Lines.*

The Planet performs its Motion fooner than the 1014 Earth ;* therefore, in its Motion, it passes be-*948 tween the Earth and the Sun, and then moves be-

M4

yond

yond the Sun, in respect of the Earth : So that it is in Conjunction with the Sun in two Manners, but never in Opposition.

That we may have an Idea of the apparent Motion of the Planet, we muft conceive the Lines T B b, T S v, T A a, to move along with the Earth; fo that the Points A, V, B, and v, whilf the Earth performs its Revolution, may run thro' the Orbit of the Planet; but the Planet, which moves fwifter, paffes fucceffively through thefe Points over and over.

1015 When it is carried in its Orbit from V to D, it feems to move amongst the fixed Stars from v to d: In this Case, the apparent Motion is in An-

1016 tecedentia, and the Planet is retrograde. In Dit is faid to be stationary ; because it appears, for fome Time, in the fame Place among t the fixed Stars : This obtains, when the Orbit of the Planet, in the Place in which the Planet is, is fo inclined to the Orbit of the Earth, in the Place in which the Earth is, that, if the Line t d be drawn parallel to the Line T D, and at a fmall Diftance from it, D d be to T t as the Celerity of the Planet in its Orbit, to the Celerity of the Earth ; thefe *53 Lines are run through in the fame Time ; * and the Line, which is drawn through the Earth and the Planet, is carried in a parallel Motion; for which Reafon the apparent Place of the Planet is *991 not changed. *

Between d and B, the Orbit of the Planet is more inclined to the Orbit of the Earth; wherefore the Extremity of the Line paffing through the Earth and the Planet (although the Planet 1017 moves fwifter than the Earth) is carried in Confequentia; towards which Part also the apparent Mo-*992 tion of the Planet is directed. * Yet fince the apparent Motion of the Sun exceeds the apparent Motion of the Planet, the Elongation is increased, which becomes

becomes greateft, when the Planet is at B. Whilft the Planet goes through the Arc B v, its apparent Motion is alfo in Confequentia, and exceeds the apparent Motion of the Sun to which it is coming, and then goes beyond it, receding from it, till it comes to A. Between A and E the Motion in Confequentia is continued; but the Sun, whofe apparent Motion in this Cafe is fwifter, as has been explained concerning the Arc d B, comes towards the Planet, and the Elongation is diminifhed. At E, in the fame Manner 1018 as at D, the Planet is flationary, between E and V it is again retrograde.

The Orbit of the Planet is inclined to the Plane of the Ecliptic, ; * therefore it does not *958. feem to move in the Ecliptic Line, but fometimes 959 lefs, fometimes more diftant from it, and appears to be carried in an irregular Curve, which fometimes cuts the Ecliptic.

Plate XXI. Fig. 3.] Let NVN be the Orbit of the Planet, whole Nodes are NN; let S be the Sun; Tt the Orbit of the Earth in the Plane of the Ecliptic; the Earth T; the Planet V. If VA be imagined to pass through the Planet, and to be perpendicular to the Plane of the Ecliptick, the Angle VTA, or rather the Arc which measures it, is the Latitude of the Planet: * This is called the Geocentric Latitude, to *1006 diffinguish it from the Latitude of the Planet feen from the Sun, which is called the Heliocentric Latitude, and is in that Case the Angle VSA. Here we speak of the Geocentric Latitude, because we examine the Phænomena as they appear from the Earth.

When a Planet appears in the Node, it appears 1019 in the Ecliptic Line; and the Curve, which is deferibed by the Planet, by its apparent Motion in the Zodiac, cuts the Ecliptic Line; as the Planet 1120 recedes

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recedes from the Node, its Latitude is encreased, which is also different, according to the Situation of the Earth; fo the Planet remaining at V, the Latitude is greater if the Earth be at T, than if it was at t. Now, if the Earth remaining in its Place, we imagine the Planet to be carried from V to v, the Angle v T B will be lefs than the Angle V T A upon a double Account, from the Planet coming nearer the Node, and the Spectator being moved farther off.

Now, if we confider that both the Earth and the Planet are continually moved, we fhall eafily conceive that the Latitude is changed every Moment from each Caufe, which fometimes act contrariwife, and fometimes confpire in encreafing and diminifhing the Latitude; whence it neceffarily follows, that the Apparent Motion is performed in an irregular Curve, which, as was faid before, cuts the Ecliptic as often as the Planet paffes the Node, that is, twice in each of its Revolutions: This Curve alfo does not recede from the Ecliptic, on either Side, beyond certain Limits in the Zodiac.

We difcover alfo fome remarkable Phænomena of the Inferior Planets by means of the Telefcope, which are owing to their Opacity.

Plate XXI. Fig. 4.] Let S be the Sun, T the Earth, A, B, C, v, D, E, F, V, an inferior Planet, ex. gr. Venus in its Orbit. This Planet fhines with Light borrowed from the Sun, and that Hemifphere only which is turned to the Sun is enlightned, the other Hemifphere is invifible: Therefore that Part only of the enlightened Hemifphere, which is turned to the Earth, can be feen from it; in V the Planet cannot be feen; in v it would appear round, if the Sun's Rays did not hinder it from being feen,

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Going

Going from v the Planet continually decreafes; 1021 at D it has the Figure d; at e and f it is drawn as it appears at E and F, and continues to decreafe, till it vanishes at V, and then again encreafes fucceflively, changing its Figure, 'till the whole enlightened Hemisphere be turned towards Earth.

When the Node is at V, or near it, the Planet 1022 appears in the very Disk of the Sun, and as it were applied to it, and is observed as a black Spot which moves on the Sun's Surface : In this Cafe, properly speaking, we don't see the Planet, but we perceive where it intercepts the Sun's Rays.

The less distant the Planet is from the Earth, the 1023 greater it appears, * and the more lucid; but as it *733 comes nearer the Earth, the lucid Part that is visible is less; so that on one Account the Light encreases, and on another it is diminiscut and there is a Distance at which the reflected Light is greatest.

CHAP. V.

Concerning the Phænomena of the Superior Planets, arifing from their Motions and the Motion of the Earth in their respective Orbits.

T HE Apparent Motions of the Superior Planets do in many Things agree with what has been explained in refpect of the Inferior Planets, and in many Things difagree. They do not 1024 always accompany the Sun, but are often observed in Opposition; but in their Opposition (as has been faid of the Inferior Planets) they do not always seem to be carried in Confequentia, but often 1025 appear stationary, and often retrograde.

Plate

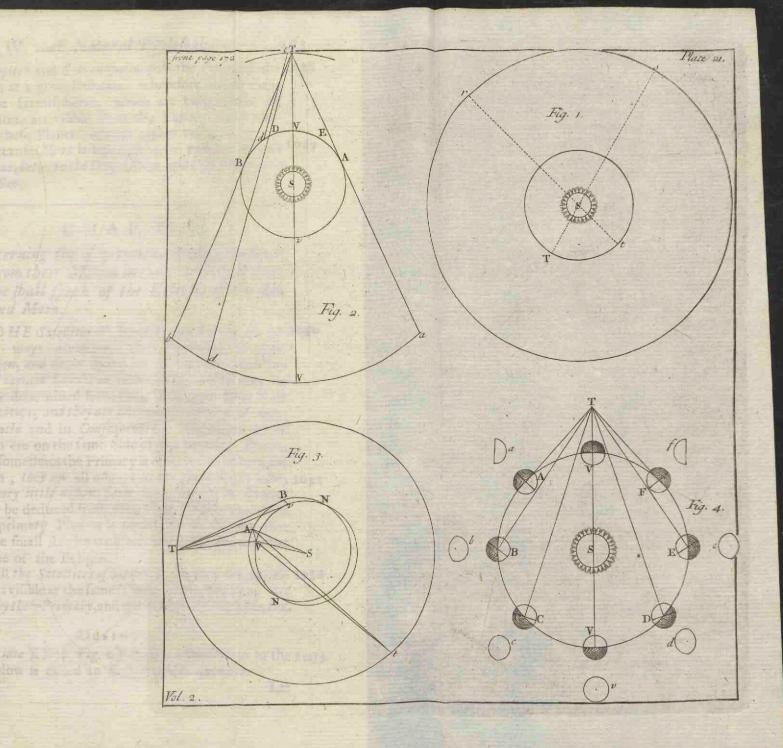
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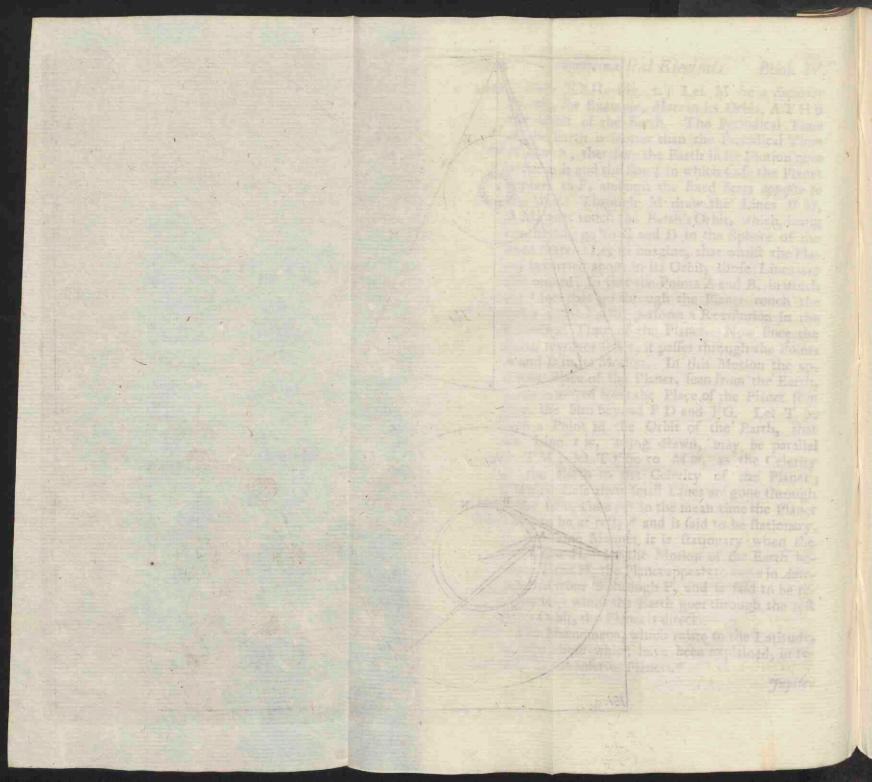
1026 Plate XXII. Fig. 1.] Let M be a Superior Planet, for Example, Mars in its Orbit, ATHB the Orbit of the Earth. The Periodical Time of the Earth is shorter than the Periodical Time "948 of Mars *; therefore the Earth in its Motion goes between it and the Sun ; in which Cafe the Planet appears at F, amongst the fixed Stars opposite to the Sun. Through M draw the Lines B M. A M, that touch the Earth's Orbit, which, being continued, go to G and D in the Sphere of the fixed Stars. Let us imagine, that whilft the Planet is carried about in its Orbit, those Lines are alfo moved; fo that the Points A and B, in which the Lines that go through the Planet touch the Orbit of the Earth, perform a Revolution in the Periodical Time of the Planet. Now fince the Earth revolves fafter, it paffes through the Points A and B in its Motion. In this Motion the apparent Place of the Planet, feen from the Earth, is not removed from the Place of the Planet feen from the Sun beyond F D and FG. Let T be fuch a Point in the Orbit of the Earth, that the Line t m, being drawn, may be parallel to TM; let Tt be to Mm, as the Celerity of the Earth to the Celerity of the Planet; in which Cafe thefe fmall Lines are gone through *53 in the fame Time; * in the mean time the Planet *994 feems to be at reft, * and is faid to be stationary. In the fame Manner it is stationary when the Earth is at H. In the Motion of the Earth between T and H, the Planet appears to move in Antecedentia from E through F, and is faid to be retrograde ; whilft the Earth goes through the reft of its Orbit, the Planet is direct.

1027. The Phænomena, which relate to the Latitude, are like those which have been explained, in re-*1008 fpect to the inferior Planets.*

PLAS

Jupiter





Jupiter and Saturn encompass the Orbit of the 1028 Earth at a great Diftance; wherefore almost their whole Hemispheres, which are enlightened by the Sun, are visible from the Earth; and therefore these Planets always appear round.

Becaufe Mars is lefs diftant, it appears a little 1029 gibbous, between the Conjunction and Opposition with the Sun.

CHAP. VI.

Concerning the Phænomena of the Satellites, from their Motion in their Orbits. Where we shall speak of the Eclipses of the Sun and Moon.

THE Satellites of Jupiter and Saturn do al-1030 ways accompany their Primaries in their Motion, and never appear to recede from them beyond certain Limits on either Side, which may be eafily determined from their Diftances from their Primaries; and they are alternately carried in Antecedentia and in Confequentia. Sometimes all of them are on the fame Side of the primary Planet, and fometimes the Primary is obferved to be between them; they are all always in the fame Right Line, 1031 or very little diftant from it. All which Things may be deduced from this, That the Motion about the primary Planets is performed in Planets that make fmall Angles with one another, and with the Plane of the Ecliptic.

All the Satellites of Saturn or Jupiter are not al- 1032 ways visible at the fame 'Time; fometimes they are bid by their Primary, and often immers' d in its Shadow.

DEFINITION I.

Plate XXII. Fig. 2.] Such an Immersion in the 1033 Shadow is called an Eclipse of the Satellite.

Let

Let S be the Sun, T t the Orbit of the Earth, I *Jupiter*, Mm the Orbit of a Secondary of *Jupiter*. Whilft the Secondary moves from M to m, it undergoes an Eclipfe; and, not being enlightened by the Sun, becomes invisible. If the Earth be at T, the Emersion into the Shadow is easily observed; on the contrary, the Emerfion is more fensible, if the Earth be placed at t.

1034 Amongst the Bodies that accompany Saturn, *963 we have faid that there is a Ring; * concerning

which it is to be obferved, that an Obferver upon the Earth never fees it wider than it is reprefented in the 4th Fig. of Plate XX. and that fometimes it is invifible; namely, when the Plane of the Ring being continued, goes through the Earth; for the Thicknefs of the Ring is not fenfible. The Ring is alfo invifible, when its Plane continued paffes between the Earth and the Sun; for then the enlightened Surface of the Ring is turned from the Earth: In each Cafe Saturn appears round, yet in the laft Cafe, by reafon of the Rays that are intercepted by the Ring, there appears a black Belt upon the Surface of the Planet, like that which is occafioned by the Shadow of the Ring.

The Phænomena of the Earth's Satellite, namely of the Moon, are very remarkable in refpect to us, and therefore particularly to be explained.

 1035 It is very often in Conjunction with the Sun, and as often in Oppofition to it, but not at every Revolution of the Moon in its Orbit; for whilst the Moon, after one entire Revolution of 27 Days and 7 Hours, returns again to the Place
 1036 amongst the fixed Stars, in which it was in Con-*960 junction with the Sun, the Sun is gone from that 996 Place, and is about 27 Degr. diftant from it :*

therefore the neighbouring Conjunctions are twentynine Days and a Half distant from one another.

DEFINITION II.

The Lunar Periodical Month is the Time of one 1037 Revolution of the Moon in its Orbit.

DEFINITION III.

The Moon's Synodical Month, or a Lunation, 1038 is the Time that the Moon spends between the two next Conjunctions with the Sun.

The Moon is invisible in its Conjunction with the 1039 Sun, because the enlightened Hemisphere is turned from the Earth. Let T (Plate XXII. Fig. 3.) be the Earth, N the Moon between the Sun and the Earth, the enlightened Hemisphere will be m li, which cannot be seen from the Earth.

Whilf the Moon is carried, in its Orbit, from the 1040 Conjunction to the Opposition, the enlightened Part, which is directed towards the Sun, does continually become more and more visible to the Inhabitants of the Earth; and in the Points A, B, C, the Moon does fucceffively acquire the Figures a, b, c.

At P, in its Opposition with the Sun, it appears 1041 round; then going through D, E, F, it decreases, as it is represented at d, e, f.

DEFINITION IV.

The Conjunction of the Moon with the Sun is 1042 called the New-Moon.

After the Conjunction, the Moon is as it were renewed.

DEFINITION V.

The Opposition of the Moon with the Sun is called 1043 the Full-Moon, because the whole Moon appears enlightened.

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Book IV.

DEFINITION VI.

1044 The Conjunction and Opposition of a Satellite with the Sun, are called by the common Name Syzygies.

1045 At A and F, the dark Part of the Moon is a little enlightened by the Rays that are reflected from the Earth; and therefore it is feen by a Spectator to whom the Sun is not vifible, that is, in the first Case, after the Setting of the Sun, and in the fecond, before its Rife.

DEFINITION VII.

1046 When the Light of the Sun is intercepted by the Moon, fo that, in respect of any Observer upon the Earth, the Sun is partly or wholly covered, the Sun is said to undergo an Eclipse.

Properly speaking, this is an Eclipse of the Earth, on whose Surface the Shadow or *Penumbra* of the Moon falls.

DEFINITION VIII.

1047 An Eclipfe of the Moon is the Obscuration of the Moon by the Shadow of the Earth.

- 1048 The Eclipse of the Sun is never observed, except at the Time of the New-Moon.
- 1049 The Moon is never eclipfed but at the Time of the Full-Moon.
- 1050 Yet the Luminaries are not eclipfed at every one of the Syzygies, because the Moon does not move
 - *969 in the Plane of the Ecliptic, * in which the Sun and Earth always are; wherefore, upon Account of the Moon's Latitude, its Shadow, at the New-Moon, often does not touch the Earth; and itfelf, at the Full-Moon, paffes befide the Shadow of the Earth.

1051 But when the Moon has no Latitude, or but very little, that is, when it is in the Node, or near it, at its Syzygies, an Eclipfe is observed; in that Plate

Cafe the Moon appears to be in the Ecliptic, or very near it; and this it is that has given the Name to that Line.

Plate XXIII. Fig. 1.] That what relates to the Eclipfe of the Moon may appear the more plainly, let OO be the Way of the Moon, R R the Plane of the Ecliptic; the Center of the *904, Earth's Shadow is always in it; *N is the Node 939 of the Moon's Orbit.

If the Center of the Earth's Shadow be at A, the Moon that goes by at F will not be darkened.

If the Moon be lefs diffant from the Node at 1052 the Full-Moon, as at G, the Shadow of the Earth is at B, and the Moon is darkened in part; this is called a partial Eclipse.

If, fuppofing the Shadow at D, the Moon be 1053 Full, the Moon will be wholly darkned at 1, it runs into the Shadow at L, and goes out of it at H; and the Eclipfe is faid to be Total.

The Eclipfe is faid to be Central, when the Cen-1054 ter of the Moon goes through the Center of the Shadow, which only happens in the very Node N.

We have hitherto fpoken of the Shadow of the Earth; because, when we mention the Earth, we understand its Atmosphere which is joined to it, of which we have spoken elsewhere: * The Shadow *418 of the Atmosphere is properly considered in Lunar 1055 Eclips; for the Shadow of the Earth itself does not reach the Moon.

Plate XXIV. Fig. 1.] Let T be the Earth, the Atmosphere about it FDGGDF. The Sun's Rays BD, BD, touching the Atmosphere; these go streight on, and terminate the Shadow of the Atmosphere, out of which if the Moon be, it is immediately enlightened by the Sun's Rays, but it is not enlightened in the same Manner all the while it is between BD and BD.

VOL. II.

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The

The Rays which enter the Atmosphere obliquely are
Tefracted; * and, as they come towards the Earth, they continually penetrate into a Medium which
tey continually penetrate into a Medium which
defer and denfer, * and therefore are every
Moment inflected and move into Curves. So the Rays E F, E F, penetrate the Atmosphere in the Curves F G, F G, that touch the Earth. All the Light between E F, E F, is intercepted by the Earth, and the Rays G A, G A, terminate the Earth's Shadow.

The Light between E F and B D, being refracted by the Atmosphere, is fcattered between G A and B D continued, and beyond A, the Point of the Earth's Shadow, the Lights that 1057 come from all Parts are confounded, but are continually weaker and weaker the farther from the Earth : So that the Shadow of the Atmosphere is not a perfect Shadow, but a weak Light, whereby the Moon is visible in an Eclipfe.

1058 The Shadow of the Atmosphere is Conical, becaufe the Sun's Diameter is greater than the Diameter of the Atmosphere, which is fearce bigger than that of the Earth; and this Cone does not reach quite to Mars, as appears from immediate Obfervations; but the Shadow of the Diameter, in the Place where it is cut by the Moon's Orbit, is fearce one Fourth lefs than the Diameter of the Earth.

With the fame Reafoning that we have proved, that the Moon comes into the Shadow of the Atmolphere, when the Moon in its full is in the Node or near it; it is alfo proved, that the Moon's Shadow falls upon the Earth at the New-Moon, when the Moon is in the Node or near the Node; 1059 therefore in that Cafe the Sun undergoes an Eclipfe; concerning which, feveral Things are to be observed.

. If has I A massed at h.

Plate

Plate XXII. Fig. 4.] Let S be the Sun, T the Moon; let the Shadow of it fall upon any Plane G H. This Shadow is encompafied with a Penumbra, for beyond L and E that Plane is enlightened by one entire Hemifphere of the Sun; as you go from L to H, and from E to G, the Light is continually diminifhed, and near G and H the Rays come to the Plane only from a fmall Part of the Sun's Surface.

DEFINITION IX.

This diminished Light, which encompasses the 1060 Shadow G H every Way, is called the Penumbra.

In the Eclipse of the Moon, the Shadow of the 1061 Earth is encompassed with the like Penumbra, but this is only sensible near the Shadow, and therefore has but a small Breadth; but if an Observer be placed upon a Plane upon which the Shadow falls, he may observe the whole Penumbra; as is 1062 the Case in the Eclipse of the Sun. An Observer I or F can only see the Semidiameter of the Sun, the rest of the Diameter is hid by the Moon; and going from L towards H, the Sun is continuelly more and more hid by the Moon, till it becomes wholly invisible in the Shadow itself.

Hence it follows, that there is a Solar Eclipfe, 1063 though the Shadow of the Moon does not touch the Earth, provided the Penumbra comes to its Surface. And alfo, that the Eclipfe is not observed in all the 1064 Places in which the Sun is visible; and that it is 1065 different, according as the Shadow or a different Part of the Penumbra goes through the Place, in the Places in which it is observed.

But the Eclipse of the Moon is every where the 1066 fame, where the Moon is visible, during the Eclipse.

But when the Shadow it felf of the Moon falls upon 1067 the Earth, the Sun's Eclipse is said to be Total; if only the Penumbra reaches the Earth, it is faid to be Vol. II. N 2 Partial;

Partial; and this is what happens when we confider an Eclipfe in general.

- 1068 But, as to particular Places, it is faid to be Total in those Places where the Shadow passes; Central, in those where the Center of the Shadow pasfes, that is, where the Center of the Moon covers the Sun's Center; and lastly, Partial, where the Penumbra only goes by; and this is drawn in Fig. 6.
- 1069 Plate XXII. Fig. 4.] The wider the Shadow G H is, the more Places is the Eclipse of the Sun Total in, and the longer is the Sun wholly obscured. But the Breadth of the Shadow is different according to the different Diffances of the Moon from the Earth, and of the Earth from the Sun.
- 1070 If there be an Eclipse of the Moon, supposing the Earth in the Perihelion, and the Moon in the Apogaum, that is, at the greatest Distance from the Earth, the Shadow of the Moon does not reach the Earth, and the Moon does not cover the whole Sun; such an one is called an Annular Eclipse, and is represented in Fig. 5.

CHAP. VII.

Of the Phænomena arising from the Motion of the Sun, the Planets, and the Moon, about their Axes.

1071 T HE Sun's Motion about its Axis is fensible, by observing the Spots, which are observed very often upon the Sun's Surface. These Spots feem to change their Figure and Situation every Day, and fometimes to move fwifter, fometimes flower; all which Things may be easily deduced from the Motion of a Spherical Surface; and the Sun, which, if it was not moved by fuch a Motion, would only once in a Year fucceffively turn

turn its whole Surface to the Earth, now fhews it to the Inhabitants of the Earth in lefs than the Space of one Month.

Such like Phænomena arife from the Rotation of 1072 Jupiter, Mars, and Venus, about their Axis, which Motions become (enfible, by observing the Spots in the Surface of the Planets.

Whilft the Earth is moved round its Axis, the Obferver, who is carried round, imagines himfelf to be at reft, and all the heavenly Bodies to be in Motion.

DEFINITION I.

The Points, in which the Axis of the Earth, be- 1073 ing continued both Ways, touches the Sphere of the fixed Stars, are called the Poles of the World.

DEFINITION II.

The apparent Motion, arifing from the Motion of 1074 the Earth about its Axis, is called The Diurnal Motion.

DEFINITION III.

A Plane is conceived to pafs through the Center 1075 of the Earth, perpendicular to its Axis, and continued every Way, and the Circle, in which it cuts the Sphere of the fixed Stars, is called The Coeleftial Æquator.

In the Motion of the Earth round the Sun, the 1076 Æquator is moved ; but fince the Plane of this Circle is carried by a parallel Motion, the Caleftial Aquator is not moved.*

*994

DEFINITION IV.

Circles, whose Planes go through the Axis of the 1077 Earth, are called Meridians.

They all pass through the Poles of the World, and 1078 are perpendicular to the Æquator.

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DEFINITION V.

1079 The Arc of any Meridian, which is intercepted between the Æquator and a Star, is called the Declination of that Star.

Plate XXIII. Fig. 2.] Let an Obferver be upon the Earth T, who directs his Sight along TA; after a little Time, when the Line 'T A shall be be carried by the Motion of the Earth to T a, if the Spectator directs his Sight through the fame Line, the Body A will appear to have been carried through the Arc a A; but when the Line has returned to its former Situation TA, the Body will feem to have performed one whole Revolution. But if he directs his Sight along the Axis of the Earth produced, becaufe that is at reft, the Body, which is feen in the Axis, will appear not to 1080 have moved; therefore in the Poles of the World the "1075 diurnal Motion is not observed. * But that Bodies, which are near the Poles, are moved round them is plain; and that the Body by its diurnal Motion defcribes fo much a greater Circle round the immoveable Pole, as it is farther dittant from it. 1081 Therefore the whole Sphere of the fixed Stars feems to revolve about the Axis of the Earth continued, in Antecedentia, in that Time in which the Earth really turns about its Axis. Therefore the diurnal Motion is common to all the coeleftial Bodies, except fo far as it is diffurbed by the Motions

The Æquator is equally diftant from both Poles, and divides the Heavens into two Hemifpheres, whofe middle Points are the Poles, which therefore are equally diftant from the feveral 1082 Points of the Æquator; therefore the heavenly Bodies which are in the Æquator by their diurnal Motion feem to defcribe the Æquator itfelf, the greateft Circle of all that can be defcribed by

above-mentioned.

by the diurnal Motion; the other Bodies describe 1083 Circles parallel to the Æquator.

The Axis of the Earth is inclined to the Plane of the Ecliptic in an Angle of 66 Degr. 31 Min. * '960 The Poles of the World therefore are diftant from 1084 the Poles of the Ecliptic 23 Degr. 29 Min. and the Plane of the Ecliptic of 23 Degr. 29 Min. Both Planes pafs through the Center of the Earth; but fince this may be looked upon as the Center of the fixed Stars, * it follows, that the Æquator and 1085 the Ecliptic Line are great Circles, which are inclined *988 to each other, and cut one another in two opposite 990 Points, in the Beginning of Aries, and the Beginning of Libra; which Points, in the Way of the Sun, are determined by these Interfections. * *1001

When the Sun is in those Points, it seems to de-1086 foribe the Aquator by its diurnal Motion; * when it 1087 is carried about in the Ecliptic by its apparent Mo-^{*1082} tion, it continually recedes more and more from the Aquator, and its Declination is increased, and it describes less Circles every Day; * till it comes to *1083 its greatest Distance from the Equator, which is 23 1088 Degr. 29 Min. * then it comes back to the Aquator *1084 again, and goes beyond it also 23 Degr. 29 Min. advancing towards the opposite Pole.

DEFINITION VI.

Those Circles, described by the Sun in its diurnal 1089 Motion, which are most distant from the Æquator, that is, 23 Degr. 19 Min. are called the Tropics.

One touches the Ecliptic Line in the first Degree of *Cancer*, and is called the Tropic of *Cancer*; the other is called the Tropic of *Capricorn*, and passes through the first Point of the Sign *Capricorn*, and there touches the Ecliptic Line.

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DEFINITION VII.

1090 The Pole of the World, which is next to the Tropic of Cancer, is called the Arctic Pole, and also the North Pole; the Opposite is called the Antarctic, and also the South Pole.

DEFINITION VIII.

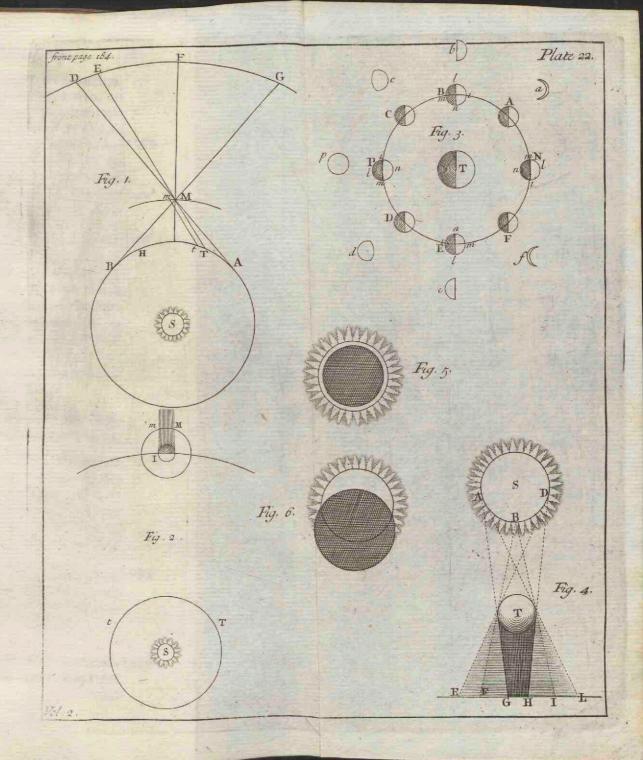
1091 The Circles that are described in the diurnal Motion by the Poles of the Ecliptic, that is, by the Points which are distant from the Poles of the World 23 Degr. 29 Min. are called the Polar Circles.

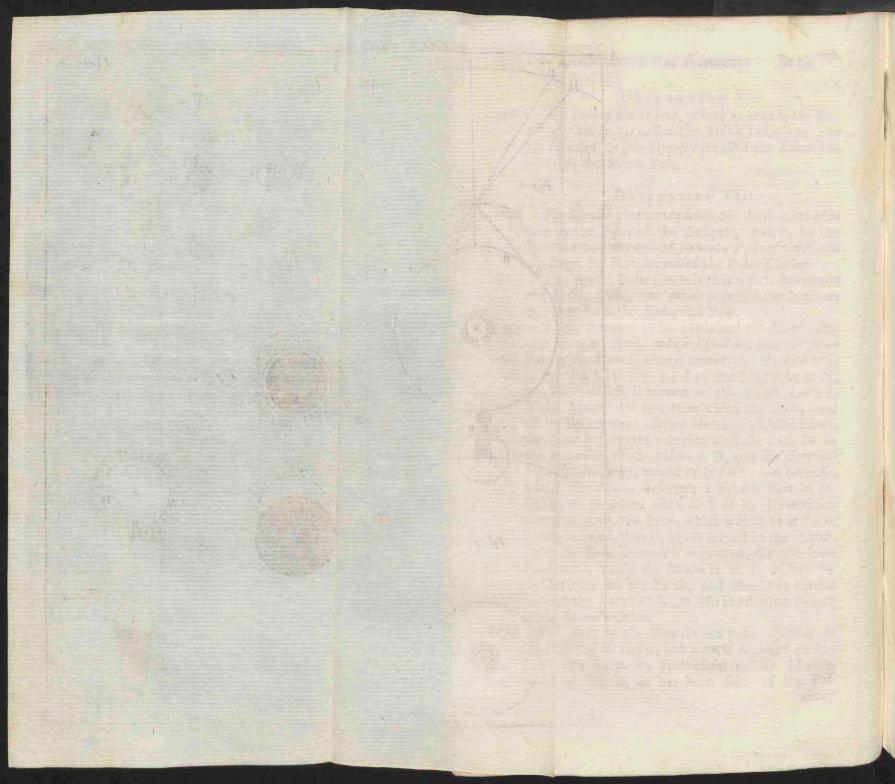
The Arctic Polar Circle is that which furrounds the Arctic Pole; the other opposite one borrows its Name from the Antarctic Pole.

1092 There remains to be explained the Moon's Motion about its Axis, whose Effect is, that the same Face of the Moon is always turned towards the Earth.

Plate XXII. Fig. 3.] Let the Moon be at N. the Face which is turned to the Earth is mni; if the Moon did not turn about its Axis, and all its Points were carried through parallel Lines, the Line mi would coincide with the Line in in the Situation of the Moon at B, and the aforefaid Hemisphere mni, would be at Imn; but because, whilft the Moon defcribes a fourth Part of its Orbit, it performs likewife ‡ of its Revolution round its Axis, the Face, which would be at 1mn, is now at mni, that is, again turned to the Earth. After the fame Mannerit is proved, that this fame Face mni, when the Moon is at P, is feen by an Obferver on the Earth, and that it is turned towards the Earth at E ; as also in all other Points of the Moon's Orbit.

1093 The Axis of the Moon is not perpendicular to the Plane of its Orbit, but a little inclined to it: The Axis keeps its Parallelifm in its Motion round the Earth, as has been faid of the Primary





mary Planets; * therefore it changes its Situation, *952 in refpect of an Obferver upon the Earth, to whom fometimes one, fometimes the other Pole of the Moon is visible, whence it feems to be agitated by a Sort of libratory Motion. There is an-1094 other libratory Motion observed in the Moon ; the Motion about the Axis is equable, and it is carried in its Orbit with an unequal Celerity ; * therefore *966 when the Moon is at its Periganm, that is, at its least Distance from the Earth where it is moved the fwifteft in its Orbit,* that Part of its Surface, *966 which, on Account of its Motion in the Orbit, would be turned towards the Earth, is not wholly turned from it on Account of its Motion round its Axis: Therefore fome of that Part of the Surface of the Moon, which before was not visible, is feen at the Side; which, when the Moon is at its Apogæum, is again vifible.

CHAP. V.

Of the Phanomena which relate to the Surface of the Earth, and its particular Parts.

W E have explained the Coeleftial Phænomena hitherto examined, by confidering the Spectator as carried about by those Motions wherewith the Earth is really moved. Now we shall confider him as placed upon the Surface of the Earth, and carried from one Place to another upon it.

The first Phænomena to be here observed is, 1095 that, by reason of the Interposition of the Earth, one Half of the Heavens is invisible to the Observer who is placed upon the Surface of the Earth.

DEFINITION I.

That Circle in the Heavens, which separates the vi-1096 fible from the invisible Part, when the Rays are

not

not intercepted by the Inequalities of the Earth's Surface, is called the Horizon.

When the Height, to which the Spectator can be raifed above the Surface of the Earth, is very fmall, compared with the Semidiameter of the Earth ; the Eye of the Spectator may be looked upon as placed in the Surface itfelf.

Plate XXIII. Fig. 4.] Let the Earth be at T, and the Obferver at S, and PE p e the Sphere of the fixed Stars; if you conceive a Plane at H H to touch the Earth and go through S, it will be the Plane of the Horizon, whole Section with the Sphere of the fixed Stars is the Horizon. A Plane, as b b, is conceived to gothrough the Center of the Earth, parallel to H H; the Diffance b H is infenfible, by reafon of the immenfe Diffance of the fixed Stars; therefore the Section of that Plane, with the Sphere above-mentioned, may be taken *994 for the Horizon.*

DEFINITION XII.

8097 The Ascent of the Stars, above the Horizon, is called their Rife.

DEFINITION III.

1098 The Defcent, below the Horizon, is called the Setting of the Stars.

DEFINITION IV.

1099 If we conceive a Line drawn through the Observer and the Center of the Earth, which must neceffarily be perpendicular to the Horizon, it will reach the Point Z among the fixed Stars, which is called the Zenith.

DEFINITION V.

1100 The Point N, opposite to it, is called the Nadir. DEFINITION VI.

1101 The Section which a Plane of the Meridian, that goes through the Observer, makes with the Horizon, is called the Meridian Line; and is directed from North to South. DEFI-

DEFINITION VII.

The Eaftern Part of the Heavens is that, from 1102 which we see the Bodies rise above the Horizon; and the Eaft Point is that, in which a Line directed Eastwards, perpendicular to the Meridian Line, and going thro' the Observer, cuts the Sphere of the fixed Stars.

DEFINITION VIII.

The Point, opposite to this, is called the Weft 1103 Point; and the Western Point of the Fleavens is opposite to the Eastern Part.

DEFINITION IX.

The Amplitude is an Arc of the Horizon, which 1104 is contained between the East or West Point, and the Point in which the Star rifes or fets. The first is called the Rifing, and the other the fetting Amplitude : And each is either Northern or Southern Amplitude.

DEFINITION X.

The Height or Altitude of a Star, above the 1105 Horizon, is the Arc of a Circle perpendicular to the Horizon, in whose Center the Spectator is, terminated by the Horizon and the Star.

DEFINITION XI.

The Difference of the Height of a Star, according 1106 to the different Position of the Observer, as he is supposed in the Center, or on the Surface of the Earth, is called the Parallax of the Star.

There is only the Parallax of the Moon, which can 1107 be determined by Obfervations: The Diffance of the reft of the Bodies in the Planetary System is too great to be compared with the Semidiameter of the Earth; and the Parallax depends upon the Ratio which the Semidiameter of the Earth has

to the Distance of a Planet; therefore even the 1108 Parallax of Mars, in Opposition with the Sun, is too fmall for the nicest Observations.

1109 Where there is a Parallax, it diminishes as a Body ascends above the Horizon, and vanishes in the Zenith.

The apparent Height of the Stars is also changed upon another Account, which equally affects all the 1110 heavenly Bodies. The Rays are inflected by the 1111 Refraction of the Atmosphere,* and the Stars ap-*1056 pear bigher than they are; * yet the bigher they *624 are, the less is that Inflexion; * because the Rays *639 fall less obliquely on the Surface of the Atmosphere. 1112 In the Zenith there is no Refraction*; even at the

*629 Diftance of twenty or thirty Degrees from the Zenith, it is not fenfible.

1113 Since the Stars are raised by this Refraction, they are visible before they come to the Horizon.

Plate XXIV. Fig. 2.] All these Things relate to the Surface of the Earth in general, now we must examine the feveral Parts of it; these are determined, by referring to the Earth the several Circles which we have before confidered in the

1114 Heavens; fo on the Earth we confider the Æquaton, the Meridians, the Tropics, and Polar Circles; and thefe Circles divide the Surface of the Earth in the fame Manner as the Sphere of the fixed Stars is divided by the Circles in the Heavens: And therefore the Circles in the Heavens, and those upon the Earth, do fo mutually correspond with each other, that a Line being drawn from the Center of the Earth to a Circle in the Heavens, it will go through the fame Circle in the Earth. If the Poles are P, p, the Æquator will be E e, the Tropics T T, tt, and the Polar Circles A A, aa.

DEFINITION XII.

1115 The Meridian, which goes through a Place, is called the Meridian of the Place. The

The Plane of it is perpendicular to the Horizon; 1116 becaufe it goes through the Center of the Earth, and the Obferver.

A Meridian Line, drawn in any Place, is Part 1117 of the Meridian of the Place.* *1101

DEFINITION XIII.

The Latitude of a Place is its Diftance from 1118 the Æquator; that is, the Arc intercepted between that Place and the Æquator.

DEFINITION XIV.

Circles, parallel to the Æquator, are called Cir-1119 cles of Latitude.

By determining the Latitude of the Place, we determine the Circle of Latitude, which goes through the Place; now, to determine the Situation of feveral Places, in refpect of each other, we must determine Places upon the feveral Circles; which is done, by fuppofing a Meridian to pass through any remarkable Place, which, by its Section, determines a Point upon each Circle of Latitude, from which the Distances of Places are measured.

DEFINITION XV.

The Meridian above-mentioned, taken at Plea-1120 fure, is called the First Meridian.

DEFINITION VI.

The Diftance of a Place from the first Meridian, 1121 measured on a Circle of Latitude that goes through a Place, is called the Longitude of the Place.

Astronomers refer every thing to the Meridian of 1122 the Place in which they make their Observations.

In explaining the Phænomena which relate to the feveral Parts of the Surface of the Earth, we fhall confider the Obferver going from the Pole

to

to the Æquator; and first only take Notice of the diurnal Motion.

1123 Plate XXIII. Fig. 3.] When the Spectator is at S, in the very Pole of the Earth T, the coeleftial Æquator E e coincides with the Horizon, and the Pole of the World P is in the Zenith; in that Cafe, becaufe the Circles which are parallel to the Horizon, are alfo parallel to the Æquator; all the heavenly Bodies appear to be carried by

*1083 a Motion parallel to the Horizon, * in Circles which are reprefented by the Lines A a, B b. The beavenly Bodies in the Hemisphere E P e never set, and the others are never visible. The Ho-

1124 rizon in this Situation is faid to be parallel, or this Situation is called a parallel Sphere.

1125 Plate XXIII. Fig. 4.] If an Observer upon the Earth T recedes from the Pole, and is at S, the Horizon is faid to be oblique, or the Sphere is oblique; then the Axis P p is inclined to the Horizon bb, fo much the more as the Observer is farther from the Pole.

DEFINITION XVII.

1126 The Angle, which the Axis of the Earth makes *1105 with the Horizon, is called the Height of the Pole.* 1127 This Height of the Pole is equal to the Latitude.

- The Height of the Pole is the Angle P T b, whofe Meafure is the Arc P b; the Latitude is meafured by an Arc, which upon the Earth cor-
- *1118 refponds to the Arc Z E in the Heavens: * But it is equal to the Arc P b; for the Complement of each of them, to a Quarter of a Circle, is the Arc Z P.

1128 In this Position of the Observer, because the Æquator is inclined to the Horizon, all the beavenly Bodies are carried by the diurnal Motion in Circles inclined to the Horizon, represented by the Lines A a, B b.

Some

Some of the heavenly Bodies rife and let at every 1129 Revolution of the Earth; namely, those which are between the Parallels to the Æquator Bh and hi; because all the Parallels between those Two are cut by the Horizon.

The Planes of the Æquator and the Horizon go through the Center of the Earth; therefore thefe Circles cut one another mutually into two equal Parts, and Half of the Æquator is above the Horizon; therefore the heavenly Bodies, which are in 1130 the Æquator, are above the Horizon during half a Revolution of the Earth about its Axis; * and, on *1082 Account of the Equability of the Motion about the Axis, are invifible during an equal Time.

These also rise due East, and set due West (that II is, in the very Points of the East and West;) for the Section of the Planes of the Æquator and the Horizon, is perpendicular to a Plane perpendicular to both these Planes; and this last Plane is the Plane of the Meridan of the Place.* Where-*1078 fore the above-mentioned Section is perpendicular to the Meridian Line, * and confequently goes *1117 through the East and West Points.*

Bodies between the Equator and a Parallel B h, 1103 which touches the Horizon, as in the Circle A a, con-1132 tinue longer above than below the Horizon; and this Difference is fo much the greater, the more the 1133 Circle A a approaches that Pole, which is above the Horizon. On the contrary, as the Body goes towards the opposite Pole, its Time of Continuance above the Horizon is the longer.

This Inequality of the Timetbat a Body is above 1134 and below the Horizon encreases as the Height of the Pole does, because of the Diminution of the Angle made with the Horizon by the Æquator and its Parallels.

Bodies, whose Distance from the Pole is equal to 1135 the Height of it, never set; for fuch is the Distance

of

of the Circle B b, which touches the Horizon, but has no Part of it below the Horizon.

Bodies, lefs diftant from the Pole, do not fo much as come down to the Horizon.

1136 It appears by the fame Reafoning, that Bodies, whose Distance from the opposite Pole does not exceed the Height of the Pole, never rife above the Horizon, and are always invisible.

1137 Bodies, whole Diftance E Z from the Æquator is equal to the Height of the Pole, go through the Zenith Z; for E Z is equal to the Latitude of the Place to which the Height of the Pole is *1126 equal. *

1138 Plate XXIII. Fig. 5.] When a Spectator S has receded as far as he can from the Pole, he comes to the Æquator, whofe Points are equally diftant *1075 from each Pole; * then the Axis P p is in the Holisitrizon, with which the Æquator makes a right *1075 Angle,* for which Reafon the Horizon is faid to 1114 be Right; or this is called a Right Sphere.

The Horizon cuts into two equal Parts all the Circles, that are parallel to the Æquator, which are reprefented by the Lines A a, B b; therefore 1139 all the heavenly Bodies, at every Revolution of the

Earth, rife, and fet, and are visible and invisible during equal Times.

1140 The Âquator itself goestbrough the Zenith, and therefore all the Bodies that are in it pass through it also.

If what we have explained concerning the Diurnal Motion be applied to the Bodies of whofe other apparent Motions we have fpoken before, the Phænomena will be eafily determined from the Motions joined together : Thofe that relate to the Sun are more remarkable than the reft, and therefore more particularly to be explained.

DE-

DEFINITION XVIII.

We call a Natural Day the Time elapsed between 1141 the Recess of the Sun from the Meridian of a Place, and its next Return to the same Meridian.

This Day differs from the Time of the Revolution 1142 of the Earth about its Axis, which Times would be equal, if the Sun appeared immoveable amongft the fixed Stars; but whilft by the Diurnal Motion, in the Time of one Revolution of the Earth about its Axis, the Sun is carried round from Eaft to Weft, that is, in Antecedentia,* it is car-*1080 ried by a contrary Motion in the Ecliptic,* where- *996 by it comes later to the Meridian.

But as the Sun does not every Day go through an equal Space in the Ecliptic,* all the Natural *997 Days do not equally exceed the Revolution of the 1143 Earth about its Axis; therefore these Days are unequal to one another.

Natural Days are unequal alfo upon another Account, namely, by reafon of the Inclination of the Ecliptic in respect to the Æquator; whence it follows, that the Annual Courfe of the Sun is unequally inclined to the Æquator in different Points ; and, though the Sun fhould equally go forward every Day in the Ecliptic, the natural Days would not equally exceed the Time of the Revolution about the Axis; for if the Motion of the Sun be refolved into two Motions, * of which 192 the one is parallel to the Æquator, and the other perpendicular to it, the first is only to be confidered in determining the Excess above-mentioned, and that it is unequal, is plain from the different Inclination above-mentioned.

These two Causes of Inequality often concur, and often act contrarywife.

Every Natural Day is divided into twenty-four 1144. equal Parts which are called Hours. Each Hour is Vol. II. O divided

divided into fixty Minutes; and each Minute into fixty fecond Minutes, or Seconds; and fo on. That thefe Parts of Time wary in different Days, appears *1143 plainly from what has been * faid; they are by Aftronomers reduced to Equality, by confidering the Number of Hours in the whole Revolution of the Sun in the Ecliptic, and dividing the whole Time into as many equal Parts as there are Hours, twenty-four of which are taken for one Day.

1145 The Time, whose Parts are by this Method reduced to Equality, is called Mean Time, and that Reduction is called the Æquation of Time.

1146 We always make use of the Days and Hoars of the Mean Time in determining the Periods of the heavenly Motions.

DEFINITION XIX.

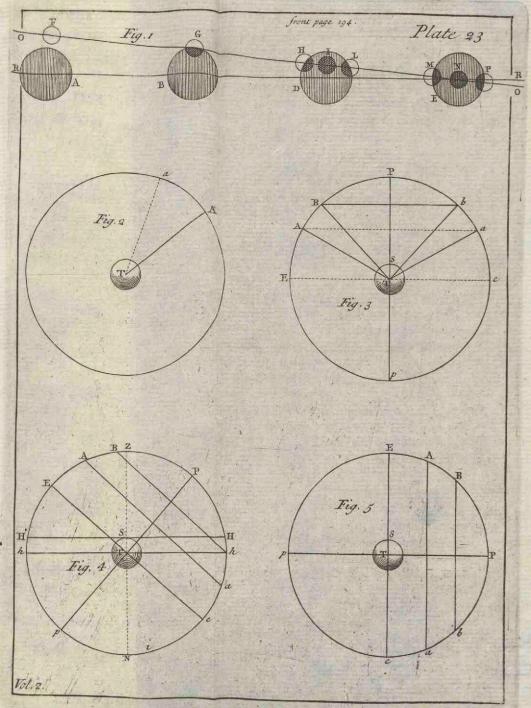
- 1147 The Artificial Day is the Time that the Sunstays above the Horizon.
- We always fpeak of it, when we mention Day 1148 in Opposition to Night. In determining the Length of Artificial Days, we shall not attend to the Æquation of Time.

1149 The Crepusculum always comes before the Sun's

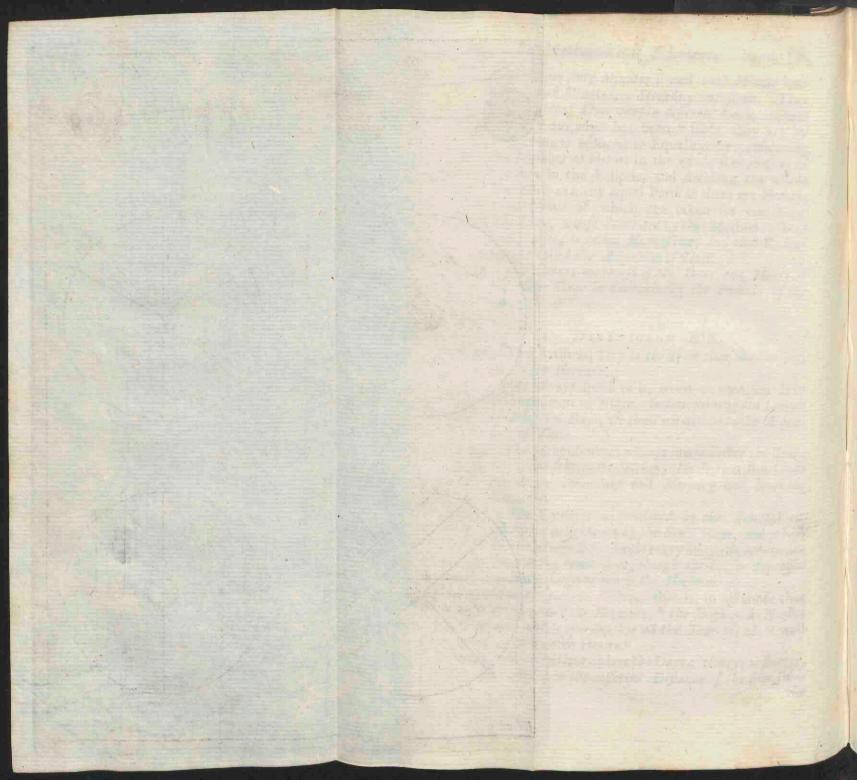
- 1150 Rife, and follows its Setting; this is that dim Light which we commonly call Morning and Evening Twilight.
- 1151 The Twilight is produced by the Atmosphere, which is enlightened by the Sun's Rays, and whose Particles reflect the Light every Way; from whence fome Rays come to us, though the Sun be depressed eighteen Degrees below the Horizon.

1152 In the parallel Sphere, that is, to all those that *1138 dwell under the Æquator, * the Days and Nights *1139 are equal to one another all the Year round, * and *1144 are of twelve Hours.*

1153 In the oblique Sphere the Days are longer or forter, according to the different Diftance of the Sun from the







the Æquator, towards either Pole;* for the Sun *131 recedes from the Æquator towards the Poles 23 *1087 Degr. 29 Min.*

The Sun is in the Æquator about the 11th of March, and the 12th of September, and then the 1154 Day is equal to the Night,* which happens all *1194 over the Earth, except just at the Poles.

DEFINITION XX.

Those Points of the Ecliptic, in which it is cut 1155 by the Æquator, * are called the Æquinoctial *1085 Points; because the Sun is in those Points, when the above-mentioned Equality of Day and Night happens.

DEFINITION XXI.

These Points of the Ecliptic, in which the Tropics touch that Circle,* are called the Solfticial *1089 Points; because for a few Days, when the Sun comes to those Points, and goes beyond them, it does not fensibly change its Declination, and the Length of the Days not fensibly vary.

Under the Poles, if there be any Inhabitants 1156 there, they can only once in a Year fee the rifing and the fetting Sun, and only one Day with one Night make up their whole Tear. The Sun continues above the Horizon all the while it goes through one Half of the Ecliptic;* the reft of the Time it is hid under the Horizon. But yet their 1157 Day is lengthened upon account of the Refraction;* *1113 and the Twilights laft very long, for they laft as long as the Declination of the Sun towards the hidden Pole does not exceed 18 Degr.* *1151

At the Arctic Pole, in the first fix Signs, from 1158 Aries to Libra, the Sun is above the Horizon; therefore at that Pole the Day exceeds the Night nine Natural Days,* befides the Diminution of the *1000 Night on Account of the Refraction.* *1157 Vol. II. O2 Thefe

195

These general Things which relate to the different Politions of the Horizon being explained, fome more particular Things are to be examined. 1159 The whole Surface of the Earth is divided into five Zones. The first is contained between the two Tro-

1160 pics TT, tt, (Plate XXIV. Fig. 2.) and called the Torrid Zone ; there are two Temperate Zones, and two Frigid Zones. The Northern Temperate Zone is terminated by the Tropic of Cancer TT, and the Arctic Polar Circle A A. The

1161 Southern Temperate Zone is contained between tt, the Tropic of Capricorn, and the Polar Circle

1162 a a. The Frigid Zones are circum[cribed by the Polar Circles, and the Poles are in the Centers of them.

In the Torrid Zone, twice a Year, the Sun goes 1163 *1137 through the Zenith at Noon.* For the Elevation *1160 of the Pole is lefs than 23 Degr. 29 Min.* and the 1127 Diftance of the Sun from the Æquator towards the Pole, which is above the Horizon, is twice in *1087 a Year equal to the Height of the Pole.* For 1088 which Reafon alfo in the Limits of that Zone, 1164 namely under the Tropics, the Sun comes to the * 1087 Zenith only once in a whole Tear.*

In the Temperate and Frigid Zones, the leaft 1089 1165 Height of the Pole exceeds the greatest Distance *1087 of the Sun from the Æquator;* and therefore to 1162 their Inhabitants the Sun never goes through the 1162 Zenith.* Yet the fame Day the Sun rifes to a greater 1166 Height, the less the Height of the Pole is; because *1137 thereby the Inclination of the Circles of the Diurnal Motion with the Horizon is lefs.

1167 In the Torrid Zone, and in the Temperate Zones, *1129 every Natural Day the Sun rifes and fets; for the 1137 Diftance of the Sun from the Pole always exceeds 1168 the Height of the Pole.* Yet every where but under * 1087 the Æquator* the Artificial Days are unequal to 1160 one another; * which Inequality is fo much the greater. 1161 1152 *1132

greater, the lefs the Place is diftant from a Frigid Zone.

But in the Polar Circles, just where the Tempe-1169 rate Zones are feparated from the Frigid ones, the Height of the Pole is equal to the Distance of the Sun from the Pole, when it is in the neighbouring Tropic; * and therefore in that Cafe, *1089 that is, once a Tear, the Sun in its diurnal Motion 1091 performs one entire Revolution without going down under the Horizon.

But every where in a Frozen Zope the Height of 1170 the Pole is greater than the leaft Diftance of the Sun from the Pole;* therefore, during fome Revo-*1089 lutions of the Earth, the Sun is at a Diftance from the Pole which is lefs than the Pole's Height, and during all that Time, it does not fet, nor fo much as touch the Horizon.* But where the Diftance *1135 from the Pole, as the Sun recedes from it, does exceed the Height of the Pole or Latitude of the Place, * the Sun rifes or fets every Natural *1127 Day;* then in its Motion towards the oppofite *1129 Pole, it ftays in the fame Manner below the Horizon, 1171 as was faid of the Motion above the Horizon.* *1136

Thefe Times, in which the Sun makes entire Revolutions above the Horizon, and below it, in its diurnal Motion, are fo much the greater, that is, the longest Day and Night last the longest, the less the Place in the Frigid Zone is distant 1172 from the Pole, till at last, at the Pole itself, they take up the Time of the whole Year.

Form the fame Caufes, namely, the Obliquity of the Ecliptic in refpect of the Æquator, by which are occasioned all the Things which relate to the Inequality of Days, which is different in different Places; we also deduce the Difference of Seasons, which fucceed one another every Year; I shall speak of them first in O 3 respect

respect to the Frigid and Temperate Zones, and then in respect to the Torrid Zone.

The Rays of the Sun communicate Heat to the Air, not only when they come directly from the Sun, but when they are reflected irregularly *582 from Bodies or the Surface of the Earth.*

This Effect is fo much greater as the Rays ftrike the lefs obliquely against the Surface of the Earth; and that upon a double Account. I. If you refolve the Motion of the Light into two "192 Motions, * one of which is parallel, and the other perpendicular to the Surface of the Earth, the Light acts upon Bodies only by this last Motion, which diminishes as the Obliquity encreases. 2. There are more Rays acting at one Time upon the fame Part of the Surface of the Earth, the more directly they come upon it.

1173 Hence we deduce, that the Caufes of Heat enereafe when the Days encreafe, by the Sun coming towards the Pole, which is above the Horizon; becaufe the Sun does daily afcend to a greater Height; fo that to the diminished Obliquity is added the longer Continuance of the Sun above the Horizon, both which concur to the encreafing of the Heat; the Nights alfo are diminished as the Days encreafe, and the Height that is produced by Day has less Time to decreafe in.

In the Northern Zones, as follows from this, the Caufe of the Heat is the greateft of all when *1090 the Sun comes to the Tropic of Cancer.* Tet the 1174 Heat is not always the greateft where the Caufe of Heat is the greateft; for the Heat encreafes as long as that which is acquired by Day is not wholly deftroyed by Night; for though the daily Agumentations be diminished, as long as there is an Augmentation the Heat encreafes. 1175 The most intense Cold is not upon the shortest Day,

Day, in which the Obliquity of the Sun's Rays is the greateft, and the Absence of the Sun the longest; but the Cold encreases, as long as the Diminution of Heat does last; concerning which one may reason in the same manner as concerning the Encrease of Heat.

The Year is divided into four Seafons; the botteft 1176 is called the Summer; the coldeft the Winter; the temperate Seafon that follows the Winter, Spring; and the Autumn comes in between Summer and Winter.

In the Northern Regions, in the Beginning of 1177 Spring, the Sun appears to be in the Beginning of Aries. In the Beginning of Summer the Sun comes to the Tropic of Cancer. When the Sun enters Libra, the Autumn begins. In the Beginning of Winter the Sun performs its diurnal Motion in the Tropic of Capricorn; all which may be eafily deduced from 1174 what has been explained.*

In the Southern Regions, the Summer happens 1178 in the Time of the Winter above-mentioned, and they have their Spring whilf the former have their Autumn; and so of the other Seasons.

The general Caufes, upon which the Division above-mentioned depends, are often diffurbed by Caufes relating to particular Places; effectially in 1179 the Torrid Zone, of which we faid we must treat of feparately. In most Places of this Zone there are only two Seasons observed, (viz.) Summer and Winter, which are chiefly distinguished by dry and wet Weather.

When the Sun comes to the Zenith of any Place, 1180 there are almost continual Rains; upon which Account the Heat is diminished, which Time is referred to, or called Winter. As the Sun recedes, the Rains diminish, the Heat is encreased, and that 1181 Time is referred to Summer.

In

1182 In the Middle of the Torrid Zone there are two Summers, and as many Winters; because the Sun *1163 comes up twice to the Zenith. *

Towards the Sides of the Zone, tho' the Sun comes twice to the Zenith, yet fince there is but a finall Time between its coming to it the first and fecond Time, both the Winters are confounded into one; wherefore only two Seafons in a Year are obferved there.

CHAP. IX.

Concerning the Phænomena arifing from the Motion of the Axis of the Earth.

*952 WE have faid that the Axis of the Earth is carried by a parallel Motion *; we have not confidered a fmall Motion, whereby it is really moved, of which we fhall now fpeak.

The Axis of the Earth, keeping the Inclination 1183 of 66 Degr. 31 Min. to the Plane of the Ecliptic, revolves in Antecedentia, that is, is fucceffively 1184 carried towards all Parts; and its Extremities, (viz.) the Poles of the World describe Circles round the Poles of the Ecliptic, from Eaft to Weft. And this Revolution is performed in the Time of about 25000 Tears; which Period is called the Great Tear.

Becaufe the Earth is looked upon as immoveable by its Inhabitants, this Motion is referred to the heavenly Bodies, as has been faid of the other Motions. 'Therefore whilft the Poles of the World are moved about the Poles of the Ecliptic in Antecedentia, and pafs fucceffively thro' all the Points that are 23 Degr. 29 Min. diftant from thefe Poles, thefe Points themfelves, or rather the

the fixed Stars that are in them, come towards the Poles of the World fucceffively, and feem to be carried in *Confequentia*, and to defcribe Circles which are really defcribed by the Poles of the World about the Poles of the Ecliptic, which, being placed in Centers, alone are at reft. For together with the Stars above-mentioned the reft of the Stars (becaufe they keep the fame Situation in refpect to one another) * do alfo ap- $*_{923}$ pear to be moved.

Therefore the whole Sphere of the fixed Stars 1185 feems to move, in Confequentia, about an Axis paffing through the Poles of the Ecliptic; and each Star apparently defcribes a Circle parallel to the Ecliptic; by which Motion the Latitude of the Stars is not changed.

The Plane of the Æquator makes a right Angle with the Axis of the Earth; therefore, by the aforefaid Motion of its Axis, the Section of the Plane of the Æquator, with the Plane of the 1186 Ecliptic, is moved round; wherefore the first Points of Aries and Libra*, which are always op-*1085 posite, move through the whole Ecliptic Line in the Space of about 25000 in Antecedentia: Yet they they are looked upon as immoveable by the Inhabitants of the Earth, who imagine that the fixed Stars themselves are moved in Confequen-*1185

CHAP. X.

Concerning the fixed Stars.

WE have faid that the fixed Stars are lucid Bodies, removed fo far off, that their Diftances can be compared with no Diftances in the

1187 the Planetary System. For Astronomers bave not been able, by their nicest Observations, to observe the Poles of the World carried out of their Place in the annual Motion of the Earth, although they defcribe Circles in the Heavens which are equal to *952 the Earth's Orbit.*

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DEFINITION I.

¹¹⁸⁸ This Motion of the Pole is called the Annual Parallax.

That the fixed Stars are at an immenfe Diftance, is also proved by Observations with the Help of 1189 Telescopes. If any fixed Star, even the lucid and confpicuous, be beheld with a Telescope, through which the Diameter of the Sun would appear equal to the Diameter of the Earth's annual Orbit, it will appear to be a lucid Point, without any sensible Magnitude; for all the fixed Stars appear lefs when they are seen through Telescopes than they do to the naked Eye; for it is only their Twinkling which makes them appear to have any fensible Magnitude.

- 1190 That the Stars may be diftinguished, they are referred to various Figures, which are imagined in the Heavens, and are called Constellations.
- 1191 Twelve Conftellations are imagined to be in the Zodiac, which are called the Sines of the Zodiac; they receive their Names from the Animals or
- 1192 Things which they reprefent : Aries, Taurus, Gemini, Cancer, Leo, Virgo, Libra, Scorpius, Sagittarius, Capricornus, Aquarius, Pisces. These Signs have given their Names to twelve Parts of the E-*999 cliptic, of which we have spoken before.*

In the Time of Hipparchus, the Sections of the Ecliptic and Æquator were between the Conftellations of Pisces and Aries, and Virgo and Libra, and the Conftellations gave their Names to those Parts

The Zodiac feparates the North Part of the Heavens from the South Part.

In the Northern Region are the following Conftel-1194 lations; the leffer Bear, the greater Bear, the Dragon, Cepheus, the Hounds, Bootes, the Northern Crown, Hercules, the Harp, the Swan, the Lizard, Caffiopeia, Camelopardus, Perfeus, Andromeda, the Triangle, the leffer Triangle, the Fly, Auriga, Pegafus, or the Flying-Horfe, Equuleus, the Dolphin, the Fox, the Goofe, the Arrow, the Eagle, Antinous, Sobiesky's Shield, Serpentarius, the Serpent, Mount Mænalus, Berenice's Hair, the leffer Lion, the Linx.

In the Southern Region of the Heavens, are the following Constellations, many of which are invisible 1195 to us: * namely the Whale, the River Eridanus, the *1131 Hare, Orion, the great Dog, Rhinoceros, the leffer Dog, the Ship Argo, Hydra, the Sextant of Urania, the Cup, the Crow, the Centaur, the Wolf, the Altar, the Southern Crown, the Southern Fish, the Phænix, the Crane, the Indian, the Peacock, the Bird of Paradife, the Southern Triangle, the Crofs, the Fly, the Chamæleon, King Charle's Oak, the flying Fish, the Toucan, or American Goose, Hydrus, or the Water-Serpent, Xiphias, or the Sword-Fish,

DEFINITION II.

The Stars, which are between the Constellations, 1196 are called unformed Stars.

The Stars are not equally lucid, and they are 1197 referred by Aftronomers to fix Class, the most lu-

cid

cid are called Stars of the first Magnitude; others are faid to be Stars of the fecond Magnitude, others of the third, &c. to the fixth Magnitude.

1198 Some are not referred even to this last Class, and are called Nebulous Stars.

1199 There is also a certain Zone or Belt observed in the Heavens, which is not every where of the fame Breadth, and goes round the whole Heavens, and in fome Places is feparated, fo as to become dou-

1200 ble. From its Colour it is called the milky Way. It is plain from Observations, by the Help of the Telescope, that this Way is an Assemblage of innumerable fixed Stars, which cannot be seen by the naked Eye, either because they are less than the other Stars, or more distant.

1201 Towards the Antarctic Polethere are two Nubeculæ, of the fame Colour as the Milky Way, which are alfo Heaps of fmall Stars, and cannot be feen without a Telefcope. Befides the Stars, which are obferved in thefe Nubeculæ, and in the Milky

1202 Way, to what soever Part of the Heavens you point the Telescope, you may discover small Stars in a great Number, which are not visible to the naked Eye. Very often an Heap of Stars appears to the naked Eye to be but one Star.

12°3 Amongh the Stars, fome are visible and invisible by Fits, and observe regular Periods; others are fucceffively fometimes more lucid, fometimes of a duller Light, and to be seen only by the Help of a Telescope, and that at certain Times.

1204 Yet they are not equally bright at every Period. Sometimes Stars have appeared fuddenly, exceeding the brightest in Light, which afterwards, fuccefsively decreasing, have vanished in a short Time, and still remain invisible.

1205 Belides the Stars, we observe in the Heavens several whitish Spots, which are in some Measure lucid, and

and invifible to the naked Eye; their Light is referred to the Stars which are in them, or they are looked upon as nebulous Stars.

What thefe Spots are, cannot be determined; perhaps they are a Congeries of Stars, which have the fame Relation to the Telescopic Stars, as those which form the Milky Way, have to those which are seen to the naked Eye.



BOOK IV.



BOOK IV.

PART II.

The Phyfical Caufes of the Celeftial Motions.

CHAP. XI.

Concerning universal Gravity.



Aving explained the Motions of the heavenly Bodies, and the Phænomena arifing from them, we must now examine by what Laws these Motions are performed.

We have before laid down the Laws according *124 to which the Motions of Bodies are directed. * 125 If we add one to thefe, we fhall fee the whole 126 Contrivance by which that vaft Machine, the Planetary Syftem, is governed.

1206 The Law, to be added to the reft, is this. All Bodies are mutually beavy (or gravitate) towards
1207 each other. This Gravity is proportional to the Quan1208 tity of Matter: At unequal Diftances it is inversity as the

the Square of the Diftance: That is, all Bodies mutually attract or tend towards each other, with the Force which belongs to each Particle of Matter acting upon each Particle; and the Force, with which a Body acts upon others, is compounded of the joint Forces of all the Particles of which the Body confifts; fo this Force encreafes in the fame Proportion as the Quantity of Matter, and is unchangeable in every Particle; it is always the fame at the fame Diftance; but the Diftance encreafing, the Force decreafes as the Square of the Diftance encreafes.

We call this Force Gravity, when we confider a 1209 Body which of itself tends towards another; because this Force is called by this Name near the Earth's Surface.*

But when we confider a Body, towards which an-1210 other tends, we call this Force Attraction. We mean the fame Effect by these Names, and nothing but the Effect; for fince all Gravity is reciprocal, * it *126 is the fame to fay, all Bodies gravitate mutually towards one another, as that Bodies mutually attract one another, or mutually tend towards each other.

We look upon this Effect as a Law of Nature^{*}, ^{*} 4 becaufe it is conftant, and its Caufe is unknown to us, and cannot be deduced from Laws that are known, as we fhall fhew by and by. Now, that there is fuch a Gravity, is to be proved from Phænomena.

All the Primary Planets are kept in their Orbits by Forces, which tend towards the Center of the Sun^{*}; therefore there is a Force by which ^{*}944 the Planets are carried towards the Sun, and ²²⁶ whereby the Sun tends reciprocally towards each of them^{*}: That is, the Sun and Planets gravi-1211 tate mutually towards each other. ^{*126}

After

- 1212 After the fame Manner it is plain, that the Satellites of Jupiter and Saturn gravitate towards each
 - *966 other; as also that the Satellites of Saturn gravi-226 tate towards their Primary Planet, and that towards. 126 them.*
- 1213 The Moon and the Earth alfo gravitate towards *966 each other.*
- 226 All the Secondary Planets gravitate towards
 126 the Sun. For they are all carried by a regular
 1214 Motion about their Primary Planets, as if the Primary Planets were at reft; whence it is plain, that they are carried about by the common Motion with the Primary Planets; that is, that the fame Force, by which the Primary Planets are every Moment carried towards the Sun, acts upon the Secondary ones, and that they are carried towards
 1215 the Sun with the fame Celerity as the Primary Planet
- nets. Even the Irregularities of the Secondary Planets, which are fo fmall, as only to be fenfible with refpect to the Moon, confirm this Gravity of the Secondary Planets towards the Sun; for we fhall fhew hereafter, that all the Irregularities are caufed by the Change of the Moon's Gravity towards the Sun, at a different Diffance; and becaufe the Lines, in which the Earth and Moon tend towards the Sun, are not altogether parallel.
- 1216 From the Gravity of the Secondary Planets towards the Sun, it follows, that the Sun gravitates *126 towards them.*
- 1217 In refpect of the Gravitation of the Primary Planets towards one another, Aftronomers have obferved, that Saturn changes its Way when it is nearest to Jupiter, which is far the greatest of all the Planets; fo that it is plain from immediate Observations, that Jupiter and Saturn gravitate towards each other.

Jupiter

Jupiter also in this Cafe, as Flamsteed has ob-1218 ferved, diffurbs the Motion of the Satellites of Saturn, attracting them a little to itfelf ; which proves that these Satellites gravitate towards Inpiter, and Jupiter towards them. From all which Confiderations * compared together, it follows, *1211 that the feventeen Bodies, of which the Planetary 1212, Syftem is made up, mutually gravitate towards 1214. each other, although no immediate Obfervations 1216, can be made concerning the Gravitation of each 1217. 1218. particular one towards the reft.* *7

The fecond Part of the Law is,* that Gravity *1207 is proportional to the Quantity of Matter; that is, that all the Particles of Matter gravitate towards each other; and therefore that the Law of Gravity is universal, and that every Body acts upon all other Bodies; which is deduced from Phænomena.

The Forces of Gravity are as the Qualities of 1210 Motion which they generate, * and thefe Quan-* 58 tities, in unequal Bodies that are equally fwift, are to one another as the Quantities of Matter: therefore fince unequal Bodies, at the fame Di-*62 ftance from the attracting Body, move equally fwift by Gravity, * it is evident, that the *1215 Forces of Gravity are proportional to the Quantity of Matter. We find the fame in all Bodies 1220 near the Earth's Surface, which have a Gravitation towards the Earth, proportional to their Quantity of Matter.* But the mutual Gravity *79 of all these Bodies towards one another is not senfible ; becaufe it is exceeding fmall in refpect of their Gravity towards the Earth; and therefore cannot difturb their Motion arising from their Gravity towards the Earth ; * at leaft, fo as to *190 make any fenfible Change in the Direction of their Motions. P

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We fhall prefently fhew, by another Method, that this universal Gravity of all the Particles of Matter, whereby they act upon one another, *1222 may be proved from Phænomena.*

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The third Part of the Law, which we examine, is, that Gravity decreafes when the Diftance encreafes, and is inverfly as the Square of the Distances, which also follows from Phænomena.

1221 Bodies, upon which Gravity acts according to their Quantity of Matter, as in our Syftem, are moved with an equal Celerity in the fame Circumftances, as we faid before; fo that it is no Matter, whether the Bodies are greater or lefs; and they are moved as if they were equal. But in this Cafe, if the Force towards a Point decreafes in an inverse Ratio of the Square of the Diftance from that Point, and the Bodies move at various Diftances from it, and are kept in Circles by that Force ; the Squares of the Periodical Times will be to one another, as the Cubes of the Di-*239 ftances.* Which is demonstrated by Geometers to obtain (in respect of the mean Diftances) in Elliptic Lines, whole Forces are directed to their Foci. But this is the Cafe in Bodies which revolve about the Sun, Saturn and '7u-*974 piter; * whence it follows, that the Force of Gravity, receding from the Centers of thefe Bodies, decreafes in an inverse Ratio of the Squares of the Diftances.

1222 By this Reafoning, fuppofing Gravity proportional to the Quantity of Matter, we demonstrate, that it decreases in an inverse Ratio of the Square of the Diftance. And by the fame Reafoning, fuppofing the Diminution of Gravity to be in this Proportion, it follows, that Gravity is proportional to the Quantity of Matter, as is very evident. But

But we prove by another Argument, that the Diminution of Gravity, fo often mentioned, is in an inverse Ratio of the Square of the Diftance : fo that there can remain no Doubt concerning the two Laws of Gravity, which we now treat of.

The Planets are moved in Orbits at reft*; ×1223 and are kept in them by Forces, which are di-*929 rected to an Excentric Point ; * but it is plain, *931 that this would not obtain, if the central Force did not encrease in an Inverse Ratio of the Square *141 of the Diftance.*

It follows from the fame Reafoning, that re- 1224 ceding from the Center of the Earth, Gravity decreafes according to the fame Law. For the Moon is retained in its Orbit by a Force which tends towards the Center of the Earth, that is, to an Excentric Point * And though the Line of *967 the Apfides is not carried by a parallel Motion, 966 its Agitation is fo fmall, if we confider every 226 Revolution, that it may be looked upon here as quiescent : For if we compute the Force which keeps the Moon in its Orbit fo agitated, we shall find the Diminution of the Force of Gravity, in respect of the Moon, to differ very little from an Inverse Ratio of the Square of the Distance; and we shall shew hereafter, that this Difference depends upon the Action of the Sun.

And no Doubt will remain concerning this 1225 Diminution, if we confider, that the Moon is kept in its Orbit by that very Force, wherewith Bodies are carried towards the Earth, near the Earth's Surface ; which is diminished, according to the Law of Diminution to often mentioned. The Mean Distance of the Moon is 601 Semidiameters of the Earth : We have before fhewn, that a Diameter of the Earth contains 3400669 Rhynland Perches; * whence knowing the Periodical Time, *976 P 2 VOL. II. we

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we eafily difcover that the Moon, in one Minute of Time, goes through 16425 1 Rbynland Perches of her Orbit. This Arc is not the hundredth Part of one Degree, and may be looked upon as its Subtenfe ; therefore the Diameter of its Orbit is to this Arc, as the Arc itfelf is to its Verfed Sine; which is difcovered to be of 15. 736 Rhynland Feet; and it is the Space which the Moon and Earth would go through in one Minute, coming to one another by their mutual Attraction. The Celerity with which a Body comes to another by the Force of Gravity, depends upon the Force with which it is attracted by that other, all whofe Particles of Matter attract it; therefore the Celerities of the Moon and Earth, as they come towards each other, are inverfly as the Quantities of Matter in them : which is also deduced from the equal Quantity of *126 Motion that is in each Body.* Therefore by 65 this Proportion we discover how much of the aforefaid Space (15, 736 Feet) is gone through by the Moon : As the Quantity of Matter, in both Bodies, is to the Quantity of Matter in the Earth; fo is the Space gone through by both Bodies, in their mutual Accefs towards each other, to the Way gone through by the Moon only. The Quantities of Matter in the Moon and in the Earth, as we shall shew hereafter, are to one another as 1, to 39, 37; and 40, 37, is to 39, 37, as 15, 736 to 15, 344, the Space gone thro' by the Moon; which therefore would be gone through in one Minute by any Body, which at the Moon's Diftance should be impelled by Gravity towards the Earth. This Force increasing in an Inverse Ratio of the Square of the Diftance from the Center, the Space gone through in the fame Time at the Diftance of a Semidiameter of the Earth, that is, on its Surface, will be 60¹/₂

 $60^{\frac{1}{2}\times60^{\frac{1}{2}\times15},344}$, (viz.) 56158 Feet; but becaufe in every Motion equably accelerated, as here (for we confider the Force as removed from the Earth's Center to the Diffance of its Surface) the Squares of the Times are as the Spaces gone through in the Fall;* by dividing the Number by *131 60×60 , that is 3600; we have the Space gone through by a Body in one Second, near the Earth's Surface, by the Force with which the Moon is kept in its Orbit, which is diffeovered to be 15, 6. R bynland Feet.

Now if we examine the Gravity which we daily find in all Bodies near the Earth's Surface, * *72 it is plain, from what has been faid concerning the Motion of Pendulums,* and from Experi-*157 ments made upon Pendulums, that a Body in ¹⁵⁸ falling goes through 15, 6. *Rbynland* Feet in one Second of 'Time, and therefore falls with the Force by which the Moon is kept in its Orbit.

In this Computation we have neglected to confider the Action of the Sun, becaufe it is fmall, and fometimes encreafes, fometimes diminifhes the Gravity of the Moon towards the Earth.

We have confidered the Centers of the Bodies in examining the Law of the Diminution of Gravity, although Gravity belongs to all the Particles of Bodies; becaufe it is plain by Mathematical Demonfitration, that the Aftion of a [pherical Body (in 1226 which, in every Part, the Particles, that are equally diftant from the Center, are homogeneous, and which is made up of Particles, towards which there is a Gravity that decreases, receding from each of them in an Inverse Ratio of the Square of the Distance) is directed towards the Center of the Body, and receding from it, is diminished in the same Inverse Ratio of the Square of the Distance : So that such a Body P 3 acts,

acts, as if all the Matter, of which it confifts, was collected in its Center. Whence we deduce the following Conclusions.

- *1207 Matter, and inverfly as the Square of the Diameter;*
- 1208 for in these Bodies the Distances from the Center are as the Diameters.

That on the Surfaces of Bodies that are fpherical, homogeneous, and equal, the Gravities are as the Densities of the Bodies; for the Diftances from the Center are equal; in which Cafe the Forces of
1207 Gravity are as the Quantities of Matter ; which,
228, in equal Bodies, are as their Densities:

79 That on the Surfaces of the Bodies, that are sphe-1229 rical, unequal, homogeneous, and equally dense, the

Gravities are inverfly as the Squares of the Diame-

- *1208 ters ;* because the Distances from the Center are *1207 in the Ratio of those Diameters, * the Gravities also are directly as the Cubes of their Diameters ; for the Quantities of Matter in Spheres are in that Ratio : And the Ratio compounded of that direct Ratio of the Cubes of the Diameters, and the inverse Ratio of their Squares, in the direct Ratio of the Diameters themselves.
- 1230 Therefore if both the Densities and the Diameters differ, the Gravities on the Surface will be in a *1228 Ratio compounded of the Densities* and the Diame-
- ¹228 Icario composition of the Dengine Gravity on the Surface by the Diameter, you will have the Den-
- 1231 fity; which confequently is in a direct Ratio of the Gravity on the Surface, and an Inverse Ratio of the Diameter.
- 1232 If a Body be placed in a Sphere that is bomogeneous, hollow, and every where of the fame Thickne/s, wherefoever it be placed, it has no Gravity, the opposite Gravities mutually deftroying one another

¹²²⁷ That on the Surfaces of Bodies, in which the Matter is homogeneous at equal Distances from the Center, the Gravity is directly as the Quantity of

another precifely: Whence it follows, that in an 1233 homogeneous Sphere, a Body, coming towards the Center, gravitates towards the Center only from the Action of the Sphere, whofe Semidiameter is the Diftance of the Body from the Center, which Gravity decreases in coming towards the Center, in the Ratio of the Distance from the Cen-*1229 ter; for all Matter, which is at a greater Diftance from the Center, forms an hollow Sphere, in which the feveral Actions on a Body deftroy each other.* * 1232

We have faid that the Gravity, which we have hitherto explained, is to be taken for a Law of Nature, becaufe we don't know the Caufe of it; and becaufe it depends upon no Caufe that is known to us, which will evidently appear, if we attend to what follows.

(Viz.) That Gravity requires the Prefence of 1234 the attracting Body; fo the Satellites of Fupiter. ex. gr. gravitate towards Jupiter, wherefoever it be.* *1212

That the Distance remaining the fame, the Cele- 1235 rity with which Bodies are carried by Gravity, depends upon the Quantity of Matter in the attract- *1207 ing Body:* And that the Celerity is not changed, let 1236 the Mass of the gravitating Body be what it will * *1207

Befides that, if Gravity depend upon any known 1237 Law of Motion, it ought to be referred to a Stroke from an extraneous Body; and becaufe Gravity is continual, a continual Stroke would be required.

If there be fuch a Sort of Matter continually ftriking against Bodies, it must of Necessity be fluid, and very fubtile, fo as to penetrate all Bodies; for Bodies, that are any how thut up in others, are heavy.

Now let a Mathematician confider, whether a Fluid fo fubtile as freely to penetrate the Pores of all Bodies, and fo rare, as not fenfibly to hinder

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der the Motion of Bodies (for in a Place void of Air, the Motion of a Pendulum will be continued very long) can impel vaft Bodies towards one another with fo much Force? Let him explain how this Force increases in a Ratio of the Mass of the Body towards which another is car-

Lafly, Let him flew, what feems most difficult to me, how all Bodies, in any Situation whatfoever (if the Diftance, and the Body to-*1236 wards which the Gravitation is, remain the fame) are carried with the fame Velocity ;* that is, how a Liquid which can only act on the Surfaces. either of the Bodies themfelves, or their internal Particles, to which it is not hindered from coming by the Interpolition of other Particles, can communicate fuch a Quantity of Motion to Bodies, which in all Bodies exactly follows the Proportion of the Quantity of Matter in them ; and which in this Chapter we have proved to obtain every where in Gravity, and which we have *77 demonstrated by a direct Experiment, in refpect 1238 of the Gravity near the Earth's Surface.*

Yet we don't fay, that Gravity does not depend upon any Stroke, but that it does not follow from that Stroke, according to any Laws known to us, and we confers that we are entirely ignorant of the Caufe of Gravity.

CHAP. XII.

Of the Celestial Matter; where a Vacuum is proved.

TI AVING explained the Laws whereby the whole Planetary System is governed, feveral Things must be first laid down, before we proceed to the Physical Explication of the System.

System. We must begin, by faying fomething of the Celestial Matter, that is, of the Medium in which the Bodies that make up the System are moved, which would be done in few Words, if all Philosophers agreed that there is a Vacuum.

We have before, proved that a Vacuum is poffible; * now we are to demonstrate that there is $*_{13}$ really one. From only confidering Motion we can 1239 deduce a Vacuum; and this is a very common and ufual Way of proving it: To fee the Force of which Argument, we must confider, that indeed all Motions are not impossible without a Vacuum, but most of those which are daily observed; which might be fully evinced by a longer Discuffion: But it feems to me to be fo evident from the following Confideration, that it would be uselefs to add much more.

The Figure of the leaft Particles is unchangeable; for the Particle, whofe Figure may be changed, confifts of fmaller Particles, which are moved in refpect to each other; and therefore if it has a changeable Figure, it is not one of the leaft Parts.

But if the Figure of thefe Particles be unchangeable, and a Body can move between them, without fuch a Separation of the Particles as to leave a void Space, this will depend on the Figure of the Particles, and the Relation which they have to one another, which a Mathematician will not deny: Therefore, if keeping Things in this State (as to their Figure and Relation) the Particles are encreafed, even in that Cafe Bodies may be moved without a Vacuum.

Now fuppofing the fmalleft Particles encreafed to the Bignefs of a Cubic Foot, whatever be their Figure and Relation with the other Parts, which we fuppofe encreafed in the fame Proportion as the firft; let any one confider, whether Bodies

Bodies of any Bignels can be carried between those Parts in Right Lines, and in all forts of Curves, and yet never feparate the Particles fo as to leave Vacuities between them.

We cannot conceive how the moft fubtile Parts are made, and therefore often attribute to them fuch Properties as do not follow from their Figure ; but thefe Errors will be corrected, by imagining the Particles encreafed.

1240 We also prove a Vacuum by an Argument taken from Relifiance.

We have faid that Matter is inactive ;* fome *12 difpute about the Word, but no Man denies the Thing; whence it follows, that a Body cannot move through a Fluid, without undergoing a Re-*319 fiftance,* and confequently a Retardation.* The *330 Refiftance arifing from the Inertia of the Matter (which Refiftance alone is here confidered) depends upon the Quantity of Matter to be removed out of its Place, which is the fame, whether the Parts of a Fluid be greater or lefs, if the Celerity of the Body remains the fame: Whence it follows, that in determining what relates to the Refiftance, we must have no Regard to the Subtility of the Fluid, as long as it cannot go through the Pores of Bodies; for if we come to fuch a Fineness of Parts, that a Fluid shall partly penetrate a Body, it will lefs refift the Body.

Now let us fuppofe any Ball or fpheric Body to be moved along in a Medium of the fame Denfity as itfelf, and fo clofe, that the Parts of the Medium cannot pafs through the Pores of the Body; this Body will be retarded every Moment, fo that its Velocity at last will be reduced to half (as may be proved by a Mathematical Demonstration) before the Body has gone through twice the Length of its Diameter.

In order to apply this Proposition to a Motion in a very fubtile Fluid, which freely penetrates the Pores of all Bodies, and fills all Places, we must conceive a fpherical Body without any Pores at all; and that fuch a Body may be fupposed by imagining all the Particles of Matter closely joined, no body will deny.

That the Refiftance of fuch a Body in any Fluid does not depend upon the Bignefs of the Parts of the Fluid, and is the fame, whether the Parts of the Fluid be equal, or any how unequal, is evident.

If every thing be full of Matter, this Body can only move through a Fluid of the fame Denfity as itfelf; for it muft run against all the Matter which is in those Places through which it paffes, and in them the Matter is without Interstices as it is in the Body; therefore it will lose half its Velocity, before it has run through the Length of twice its Diameter.

Now let us fuppofe the Body to be encreafed, the Quantity of Matter remaining the fame, and the Body continuing homogeneous; that is, let there be Pores in the Body, through which the most fubtile Particles of Matter may pass very freely; and let thefe Pores be equally difperfed all over the Body. If the Body, thus changed, be moved, the very fubtile Fluid, of which we fpeak, will not run against the whole Surface, but only those Parts of the Surface which are between the Pores, which Parts being taken together, becaufe we fuppofe the Body homogeneous, are equal to the Surface of the Body in the first Supposition, when we conceived it to be without Pores; for the Body being encreafed. the Surface has not been changed, but only dilated by the Interpolition of Pores : So that in both Cafes the Body will undergo the fame Refiftance

fiftance from the Impulfe upon the Surface; and the Refiftance on the dilated Body is greater from the Fluid running against the internal Parts of the Body. Wherefore this Body will fooner lose half its Motion in the fecond, than in the first Cafe; that is, before it has run through the Length of two Diameters of the first fupposed Bigness; and therefore it loses a greater Quantity of Motion, in going through two Diameters of the Bigness fupposed in the fecond Cafe.

But this is contrary to Experience; for a homogeneous Ball of Gold, or Lead, &c. lofes a much lefs Quantity of Motion than what we have mentioned, in Water or Air; whence it follows, that the Supposition that all Things are full of Matter, is falle. Therefore there is a Vacuum.

1241 That there is a Vacuum does also agree with the Phænomena relating to Gravity; by which it follows that it is proportional to the Quantity of Matter. If all was full of Matter, Gravity would act equally every Way, and the Forces which are directed towards opposite Parts would deftroy one another; and therefore no fensible Gravity would be observed, which is contrary to Experience.

Thefe being premised, we must return to the Celestial Matter.

1142 The Motions of the heavenly Bodies do not depend upon the Motion of the Celeftial Mat-

1238 ter, if there be fuch a Matter ;* whereby is overthrown the Opinion of those which say that the beawenly Bodies are carried along by the common Motion of the Matter which fills our System. This Opinion is also overthrown by the Motion of the Comets: If there was a Medium in the System which carried about the Planets in its Motion, and also the Comets, it would at least fensibly difturb these last in their Motions, whilf they come

come almost directly towards the Sun, or directly recede from it, or are carried in Antecedentia : that is, in a Motion contrary to the Motion of that Matter. Now as this Motion is not diffurbed, but follows the Way which depends upon Gravity, as it is obferved, it is plain, that if there be any Celeftial Matter, and that it is in Motion, it does not exert a fenfible Action on the Bodies of the Planetary Syftem, which is alfo deduced from the fmall Refiftance of fuch a Medium ; for by comparing the most ancient Obfervations with the modern, it does not appear that the Planets are fenfibly retarded in their Motions. Yet in Air the Refiftance is fenfible; wherefore the Denfity of the Medium, in which the Planets fhould move, must be almost immenfly lefs ; therefore the Planetary System is not 1243 filled, unless it be by fuch a subtile Medium.

But we may from the Divifibility of Matter deduce, that a Quantity of Matter, how fmall foever it be, may be difperfed all over the Planetary Syftem, leaving but very fmall Interftices. *

CHAP. XIII.

Concerning the Motion of the Earth.

BESIDES the Queffion that has been handled in the foregoing Chapter, there is also another to be examined, before we proceed to the Explication of the whole System.

That no Doubt may be made concerning the Syftem, which has been explained in the first Chapter of this Book, we must here prove the Motion of the Earth, concerning which it is no Wonder that many have doubted; for the Celeftial * 20

leftial Motions cannot be determined by us, but by Obfervations made by Obfervers on the Earth; and the fame Phænomena appear, whether the Bodies themfelves be moved, or the Spectator

*993 be moved; * fo that it is not to be proved by immediate Obfervations, whether the Motion of the Earth is to be referred to the heavenly Bodies or not.

44 That the Earth is carried about the Sun, is deduced from the Analogy of the Motions, and from an Examen of the Laws of Nature.

As to what relates to the Analogy of the Motions, it is to be obferved, that Satellites revolve about Jupiter and Saturn, which are lefs than the central Body; that the Moon revolves about the Earth, than which it is lefs. Laftly, That the Sun has revolving about it lefs Bodies than itfelf, as Mercury, Venus, Mars, Jupiter, and Saturn. Now if the Earth revolves with the 1245 reft, then every where in our System the leffer Bodies move about the greater : Now there would be an Exception in this Rule, in respect of the Sun, if that vaft Body was to goround so fmall a Body as the *975 the Earth. *

About the Sun, Jupiter and Saturn, about which feveral Bodies revolve, those move the flowest which 1246 are most distant from the central Body, and according to this Rule, that the Squares of the Periodical Times follow the Ratio of the Cubes of the *974 Distances; * which Rule may be applied to the Earth, if it be carried about the Sun with the rest of the Planets, as appears, if its Periodical Time (namely the Time in which the Sun appears to perform an entire Revolution) and its Distance from the Sun be compared with the Distances and Periodical Times of the rest of the Planets.

Now this Rule bas only one Exception, if the Sun be moved about, the Earth is at rest. In this Cafe

Cafe Mercury, Venus, Mars, Jupiter, and Saturn, are fubject to this Rule in their Motions, as alfo the five Satellites of Saturn, and the four Planets that accompany Jupiter; only the Moon and the Sun would move above the Earth in a Proportion quite different; and then the Celerity of the Sun would not only be greater than is required by this Law, but its Velocity would at leaft be fix and twenty Times greater than that of the Moon, though it be removed to a vaft Diffance from the Earth, in refpect to the Moon's Diffance : And therefore in this refpect, the Analogy of the Celeftial Motions would be diffurbed.

To thefe Arguments I shall add others, whereby it will clearly appear, that the Motion of the Earth is a necessary Confequence of the Laws of Nature, which are deduced from Phænomena.

All Bodies gravitate towards one another, * 1247 therefore the Sun and Earth do; but the Mo- *1206 tion, whereby these Bodies tend towards one another, is deduced from direct Obfervations. Whichfoever of thefe Bodies moves about the other, defcribes Area's by Lines drawn to the Center of it, proportional to the Times, which is evident from Aftronomical Obfervations; therefore the Body moved is retained in a Curve by a Force, which is directed towards the Center of the other. * Now as Re-action is always *226 equal to Action,* unlefs the Laws of Nature, *126 which obtain conftantly every where, be wholly overturned, these two Bodies tend towards one another with equal Motions ; that is, with Celerities that are inverfly as their Maffes; * which *65 is alfo immediately deduced from the Law of Gravity. * *1235

The Quantity of Matter in the Earth is next to nothing, in comparison to the Quantity of Matter Matter in the Sun, as we shall shew in the following Chapter; wherefore the Sun must move very flowly whils the Earth comes towards it very swiftly.

Whence it follows, that the Earth is carried round the Sun, left it should fall upon the Sun by that very violent Motion whereby it is retained in its Orbit.

This Motion of the Earth is alfo deduced from the fame Principles another Way.

Two Bodies, that are carried towards one another by any Force, will at last concur, or continually recede from one another, unlefs each of them be fo moved, as to have a centrifugal Force equal to the Force whereby it is carried towards the other Body; but as the Bodies, which gravitate towards one another, tend towards *126 each other with equal Forces,* or what is the * *os fame,* with Celerities that are inverfly as the *123: Quantities of Matter;* thefe Bodies cannot per-1248 fevere in their Motions about one another, unlefs both of them be fo moved, as to have equal centrifugal Forces, which does not happen, unlefs they both revolve in equal Times about their common Center of Gravity: That is, if this Proposition be applied to the Sun and Earth, unlefs they both move about a Point, whole Diftance from the Center of the Sun is to its Diftance from the Center of the Earth, as the Quantity of Matter in the Earth is to the Quantity of Mat-*234 ter in the Sun, they cannot perfevere in their ²³⁵ Motions about one another :* This Point or Center of Gravity must of Confequence be very near to the Sun's Center. Now fince whichfoever of thefe Bodies moves, it perfeveres in its Motion about the other, it follows, that both of them are affected by the Motions above-mentioned, and that the Sun is moved but

but a little, whilft the Earth defcribes a very great Orbit. Whence it follows, that the Motion of the Earth cannot be denied by any one, who reafons from the Laws of Motion that are deduced from Phænomena.

Having proved the annual Motion of the Earth 1249 and brought back the Earth amongft the Planets, there remains but little Difficulty in relation to the Motion of it about its Axis; for no body, that believes the annual Motion, doubts of this; a great many, which allow of the Motion about the Axis, deny the annual Motion; therefore it will be enough to obferve by the By, that all the Planets, concerning which any Obfervation could be made in refpect of this Motion, do move about their Axes; and that the Earth has fuch a Motion, the uniform diurnal Motion in Bodies, at any Diftances, does plainly enough thew. To which we must add, that the Celerity of the fixed Stars, going through one whole Revolution lefs than 24 Hours, can hardly be more probable than it is conceivable.

This Motion alfo is difagreeable to the Nature of all the heavenly Bodies; for, if they are car- 1250 ried round, they must every Day, with an equable Motion, defcribe Circles that have the Earth for their Center; that is, they muft, by Lines drawn to the Center of the Earth, fweep through Area's proportionable to the Times, and be retained in their Orbits by Forces which are directed towards the Center of the Earth,* and by *226 which (by reason that Action and Re-action * *126 are equal) the Earth must also be continually attracted towards those Bodies; fo that it must neceffarily be agitated by a very violent Motion ; whence it appears that the diurnal Motion muft not be referred to the heavenly Bodies, but to the Rotation of the Earth about its Axis.

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1251 Those that obstinately affirm that the Earth is at Rest, object that Bodies, upon the Surface of the Earth, must (on account of their centrifugal Force) recede from the Earth, along a
217 Tangent to a Circle parallel to the Æquator. We answer, that the Bodies in the Places where they are, are carried round with the same Motion as the Surface of the Earth; and therefore that, in respect of the Points of the Surface, they en*223 Axis;* but also that Bodies by Gravity tend to the *1226 Center of the Earth;* and therefore by a Motion compounded of both these, the Body is continu-

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190 ally moved, or endeavours to move ; but becaufe 198 the first Motion is extremely fmall in respect of the other, a heavy Body is turned but very little out of its Direction towards the Center, and the Gravity is a little diminished, fo much the more as the Place is more diftant from the Pole; which agrees with Experience. We fhall hereafter fhew, when we come to fpeak of the Figure of the Earth, that the above-mentioned Direction of heavy Bodies is every where directed perpendicularly to the Surface of the Earth. A Body, which is thrown upwards, is acted upon, not only by the Motion wherewith it is thrown up; but it is also carried by the Motion that is imprefied to the Perfon or Machine that impels the Body ; that is, it is carried by the Motion which is common with the Surface of the Earth ; and therefore the Body moves in the fame Line (the Line being carried on with the Surface of the Earth) as it would do if the Earth was at Reft.

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CHAP. XIV.

Concerning the Density of the Planets.

BEFORE we proceed to the Phyfical Ex-planation of the System, we must determine the Quantities of Matter in fome Bodies, and their Diverfities ; which being known, the Effects of the Laws, by which these Bodies are governed, will more eafily appear.

The Quantities of Matter, in different Bodies, are to one another, as the Gravities at the fame Diftance from thefe Bodies ;* which Gravities *1207 are to one another inverfly as the Squares of the Periodical Times of the Bodies revolving about those different Bodies at the fame Diftance.* *236 By multiplying the Quantities which are in this Ratio by the fame Quantity, (viz.) by the Cube of this Diftance, the Ratio of these Quantities will not be changed ; which are therefore to one another as the Quotients of the Divisions of the above-mentioned Cube, by the Squares of the Periodical Times aforefaid : But the Quotient of fuch a Division is found for any Body, by dividing the Cube of the other Diftance; let it be what it will, by the Square of the Periodical Time of the Body revolving at that Diftance; for fuch Quotients are equal to one another, for all Bodies that revolve about the fame Body at any Diftances; as follows from the Equality of the Ratio between the Cubes of the Diftances, and the Squares of the Periodical Times at those Diftances.* From which we deduce, that the *974 Quantities of Matter, in any Bodies in our System, 1252 are to one another directly as the Cubes of the Diftances at which other Bodies revolve about thefe. and

and inversity as the Squares of the Periodical Times of these revolving Bodies.

These Things are demonstrated by setting afide the Agitation of the Central Body, whose Quantity of Matter is enquired after.

By reafon of the Sun's Magnitude, in refpect of Venus, ex. gr. which alone we confider of the Planets, the Sun is fcarce moved by the Action *2135 of that Planet.* And Venus may be confidered as moving about a quiefcent Body.

The Satellites of Jupiter and Saturn are indeed carried by the common Motion along with the Primary Planets, but by reafon of the Magnitude of the Primary Planets, they are carried about them as about Bodies that are at reft.

But the Moon acts fenfibly enough upon the Earth and moves it; wherefore before we can compute the Motion of the Moon by the Help of the *1252 aforefaid Rule,* in order to compare the Quanti-1253 ty of Matter in the Earth with the Quantities of Matter in the Sun, Jupiter, and Saturn, we must determine the Diffance at which the Moon would move about the Earth, if it was at Reft (that is, not carried about by the Action of the Moon) in the fame Periodical Time in which it now performs its Revolution. Here also we don't take Notice of the Motion that is common to the Earth and Moon, by which they are both carried about the Sun.

The Moon perfeveres in its Motion about the Earth; therefore the Earth and Moon are moved about a common Center of Gravity; as follows from what has been demonstrated concerning the *1248 Earth and the Sun,* and the Moon (with that Force with which it tends towards the Earth) revolves in an Orbit whose Semi-Diameter is the Distance of the Moon from the aforefaid common

mon Center of Gravity of the Moon and the Earth.

Let L be this Diftance of the Moon from the common Center of Gravity; T the Diftance of the Earth from the fame Center; L + T therefore is the Diftance of the Moon from the Earth. and is 60 ; Semi-Diameters of the Earth ; for here we confider the mean Diftance. Let D be the Diftance which we would have, at which the Moon, by its Gravity towards the Earth, would move about the Earth, if it was at Reft, in the fame Time in which it is now moved about the common Center of Gravity at the Diftance L.

By reafon of this Equality of the Periodical Times, the Force, whereby the Moon would be kept in its Orbit at the Diftance D, is to the Force whereby it is kept in its Orbit at the Diftance L, as D to L.* \$232

But the Force whereby the Moon would tend to the Earth, and be kept in its Orbit at the Diftance D, is to the Force whereby it is now kept in its Orbit at the Diftance L + T, as $L + T^{q}$ to D^{q*}. Therefore *1208

$D, L :: L + T^q, D^q$

Confequently $D^c = L \times L + T^q$, and $D^c \times L + T$ $= L \times L + T^{\circ}$: Whence we deduce the following **Proportion**:

L+T, D^c :: L+T, L. Therefore L + T, D :: L + T, is to the first of two mean Proportionals between L + T and L.

L+T is to L, as the Quantity of the Matter in the Earth and Moon taken together, to the Ouantities of Matter in the Earth alone ;**234 which Quantities of Matter, as we shall shew 235 hereafter, are to one another, as 40, 37. to 39, 37.

39, 37. and the first of two mean Proportionals to these Numbers is 40,035; therefore 40, 37. is to 40,035. as $60\frac{1}{2}$ to the Distance required, which is found to be 60 Semi-Diameters of the Earth.

Concerning this Operation it is to be noted, that the Diftance D cannot be difcovered, unlefs the Ratio between the Mass of the Moon and the Earth be known, which cannot be determined unlefs the Ratio between the Denfity of the Sun and the Earth be found : To difcover which, it is neceffary that the Diftance D be known. Wherefore D is difcovered at first by Trials, and is exactly determined by Approximation. But it is certain, that this is 60 Semi-Diameters of the Earth ; becaufe, this being fuppofed, it is found that the Ratio between the Quantities of Matter of the Moon and the Earth is as I to 39, 37. as we shall fee hereafter; by making use of which Proportion, this Diftance is discovered to be 60 Semi-Diameters, as we have thewn.

These Things being premised, we proceed to the Computation.

The Diftance of Venus from the Center of the Sun is 723, and its Periodical Time is 5393 *959 Hours.

The Fourth Satellite of Jupiter is diftant from the Center of Jupiter 12, 507. fuch Parts of which Venus is diftant from the Sun 723. The Periodical Time of this Satellite is 402 Hours, *971 5 Minutes.*

The Fourth Satellite of Saturn is diftant from the Center of Saturn 9, 292. of the fame Parts; *97² and its Periodical Time is 382 Hours 41 Min.*

Laftly, the Diftance of the Moon is 60 Semi-Diameters of the Earth from its Center, and 2,909.

2,909. of the aforefaid Parts. Its mean Periodical Time is 655 Hours, 43 Min.

If you divide the Cubes of thefe Diftances re-1254 fpectively by the Squares of the Periodical Times, you will have, in the Quotients, Numbers which are to one another as the Quantities of Matter in the aforefaid Central Bodies; **1252 which Quotients are to one another as the following Numbers:

Quanti-? In the Sun. In Jupiter. In Saturn. In the Earth ties of Matter. \$ 10000. 9, 248. 4, 223. 0,0044.

We have also the Proportion of the Diameters 1256 of these Bodies from Astronomical Observations, as follows.

Diame- ¿Of the Sun.Of Jupiter.Of Saturn.Of the Earth ters 5 10000. 1077. 889. 104.

If the Quantities of Matter above faid be di-1257 vided by the Squares of the Diameters, the Quotients will be to one another as the Weights on the Surfaces of the afore faid Bodies; * and the fe *1227 Quotients are as the following Numbers.

Gravities Of the Sun.Of Jupiter.Of Saturn.Of the Earth 1258 on the Surfaces \$ 10000. 797, 15. 534, 337. 407. 832.

If you divide these Numbers by the Diame-1259 ters, you will have the Proportion of the Denfities of those Bodies.*

The Quotients, found by these Divisions, are as the following Numbers.

Den- 2 Of the Sun. Of Jupiter. Of Saturn. Of the Earth 1260 fities 5 10000. 7404. 6011. 39214.

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We

We shall determine the Density of the Moon in the last Chapter.

It is not probable that the aforefaid Bodies are homogeneous. We fhall fhew in relation to the Earth, in the 17th Chapter, that it is denfer towards the Center than towards the Surface; from whence it follows, that the Denfities cannot be exactly determined: Wherefore we only determine the mean Denfities, that is, which the Bodies would 1261 have, if the fame Bodies, keeping the fame Quan-

tity of Matter and Bulk which they now have, Jhould become homogeneous.

1262 The above-mentioned Proportion * between the * 1260 Densities, in respect of all the Bodies, and the rest of the Computations in respect of the Sun, Jupiter, and Saturn, are free from any lensible Error : When they are compared with the Earth there may be some Error, which must be corrected by Observations to be made hereafter ; for we suppose the Distance of the Moon (which is 60 Semi-Diameters of the Earth) to be 2, 909 fuch Parts, of which Venus is diftant from the Sun 723, that is of which the ⁹⁵⁹, Earth is diftant from the Sun 1000 ;* which Di-960 ftance of the Moon is difcovered by fuppofing the Horizontal Parallax of the Sun 10", which cannot be looked upon as abfolutely true, although it be deduced from the most exact Observations that have hitherto been made, of the Parallax of Mars, when it is nearest of all to the Earth, which is too fmall to leave us without Sufpicion of *1108 fome Miftake.*

But the Error, in not determining truly the Proportion between the Semi-Diameter of the Earth and the Diftance from the Sun, does not change the determined Denfity of the Earth, as is deduced from Computations made about it.

For it follows from thefe, that the Denfities of Bodies are to one another in a Ratio compounded

pounded of the direct Ratio of the Cubes of the Diftances of the Bodies carried about, and the inverse Ratio of the Squares of the Periodical Times of these revolving Bodies ;* as also of *1254 the inverse Ratio of the Cubes of the Diameters of the Central Bodies whofe Denfities are required ;* the Ratio, compounded of thefe, is com- *1257 pounded of the direct Ratio of a Fraction whole 1259 Numerator is the Cube of the Diftance of the revolving Body, and whofe Denominator is the Cube of the Diameter of the Central Body, and the inverse Ratio of the Square of the Periodical Time of the Body carried about. But you have fuch a Fraction, if you know the Ratio between the Diameter of the Central Body and the Diftance of the revolving Body from that Center. although this Diftance can be compared with no other ; but this Ratio is given in refpect of the Earth and Moon, as well as in respect of the other Bodies; wherefore alfo the Ratio of the Denfity of the Earth to the Denfities of the other Bodies is exactly difcovered.

CHAP.XV

The Phylical Explanation of the whole Planetary System.

T N the first Part of this Book we have shewn. what are the Motions of the Bodies in the Planetary Syftem; now we must explain how these Motions follow from the Laws of Nature :* that is, how thefe Bodies, being once put in Motion, perfevere in those Motions which we observe. 1208

Let us conceive the Sun and Mercury to be left 1263. to themfelves, and they will come together ;* *1206 but if they be projected, they may revolve about

124. 1250 126. 1206, 1207.

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a common Center of Gravity in equal Times, *1248 and defcribe immoveable Elliptic Lines, * and 1208 perfevere in that Motion; for it is plain, by Ma-241 thematical Demonstration in that Cafe, that the Bodies will defcribe Ellipses about the common Center of Gravity, fimilar to that which the one of them could defcribe with the fame Forces about the other, if it was at Reft : This Center, *235 on account of the Sun's Magnitude, * is very little diftant from the Center of the Sun itfelf.

Let us conceive befides, Venus to be projected at a greater Diftance from the Sun, it will a little difturb the Motion of Mercury, which alfo, by its Action upon Venus, will turn it a little out of the Way, and both will draw the Sun, fometimes the fame Way and fometimes different Ways; but we find all thefe Irregularities are infenfible, if we confider the Magnitude of the Sun; and therefore that thefe three Bodies tend towards a Point that is between them near the Sun, which therefore is very little diffant from the common Center of Gravity of them all.

If the Earth, Mars, and the other Planets be fucceffively projected at different Diffances from 1264 the Sun, the fame Reafoning will hold good. Whence it follows, that all the Planets are revolved about the common Center of Gravity of all the Bodies which compose the System, which is but little diftant from the Sun, and that the Planets do not 1265 fenfibly difturb one another in their Motions; wherefore they all describe the same Line fingly, which they would describe about the Sun, if every one of them was alone with the Sun in the Planetary Syfrem, that is, immoveable Ellipfes : For it is plain, that thefe will be defcribed by the Force *1208 of Gravity; * and it is proved by Mathematical 2+1 Demonstration, that no other immoveable excentric

centric Lines can be described by a Central Force acting equally at equal Distances.

It will also more plainly appear, that all the Planets tend to a Point near the Sun, if we confider that the Quantity of Matter in the Sun is a thousand times, and more, greater than the Quantity of Matter in *Jupiter*, which is far the greatest of all the Planets.*

When all the Planets move, though they move 1226 the Sun but little, yet they do move it, and draw it differently according to their different Situation in refpect of one another; whence there arifes a fmall Motion in the Sun, which always depends upon the Motion already acquired, and the Change which happens in it from the Action above-mentioned, which varies every Moment.

It is owing to this Agitation of the Sun, that the 1267 Planets disturb one another less in their Elliptic Motions round the Sun, than if the Sun was at Reft in the Middle of the System. If Jupiter, ex. gr. was equally diftant from Mercury and the Sun, it would attract both those Bodies to itself with an equal Celerity ;* whence the Situation, in refpect *1235 of the Sun, is lefs changed than if the Sun was not agitated by this Motion, and Mercury only was attracted by Jupiter : According to the various Diftances of Mercury and the Sun from Jupiter, the one or the other is more attracted. and there is always a lefs Change in their refpective Situation when both are carried the fame Way, than if (the Sun being at Reft) Mercury only should be carried towards Japiter.

This Reafoning may be applied to all the Actions of the Planets, that are more diftant from the Sun, upon those that are less diftant. As to what relates to the Action of those that are nearer, upon those that are farther from the Sun, according to the different Situation they draw

draw a Planet to the Sun, or drive it from the Sun, and in confidering one whole refpective Revolution, that is, the Motion from one Conjunction to another, the Difturbance is lefs than if the Sun was immoveable.

1268 The Magnitude of the Sun, compared with the reft of the Bodies of our Syftem, is the Reafon (as appears by what has been already demonftrated) that the Planets diffurb one another but little; but fince the Magnitude is not infinite, thefe mutual Actions must not be wholly overlooked.

We have faid, that it appears by Aftronomical Obfervations, that *Jupiter* alters the Way of *1217 Saturn when it is neareft to it; * why this Difturbance is more fenfible than the reft, is deduced from the Law of Gravity.

- 1269 The Actions of Jupiter upon Saturn when it is nearest to it, and of the Sun upon the fame Planet, by which it is kept in its Orbit, are to one another directly as the Quantities of Matter in *1207 Jupiter and the Sun,* (viz.) as 9, 248. to 10000,*
- *1255 and inverfly as the Squares of the Diftances of Jupiter and the Sun from Suturn; that is, directly as 81 to 16; for the Diftances of Saturn and
- 1270 Jupiter from the Sun are almost as 9 to 5; wherefore, when Jupiter is nearest to Saturn, the Distances of Saturn from Jupiter and the Sun are as 4 to 9. The Ratio compounded of the two aforefaid Ratio's is as 749 to 160000, or as 1 to 214. This Action of Jupiter conspires with the Gravity of Saturn towards the Sun, and therefore encreases the Part: Whence it is no Wonder that the Disturbance is fensible.

We don't here confider the Force by which Jupiter attracts the Sun, for the Orbit of Saturn is not changed by it, and what we had to explain was, why Aftronomers obferve Saturn to be

be turned out of the Way; yet by the Action of Jupiter upon the Sun, the Sun is brought nigher to Saturn, and the refpective Situation of thefe Bodies is more diffurbed, than is difcovered by Aftronomical Obfervations. The Force with which Jupiter in the aforefaid Polition attracts the Sun, and with which therefore the Sun is attracted towards Saturn, is to the Force with which Jupiter attracts Saturn, as 16 to 25;* that is, *1208 as 479 to 749, which Number expresses the Force with which Saturn tends towards Jupiter, when the Gravity of Saturn towards the Sun is expressed by 160000. If we collect into one Sum the Forces of Jupiter, by which it attracts Saturn and the Sun; the Force, by which, from the Interposition of Jupiter, these Bodies tend towards each other, will be to the Gravity of Saturn towards the Sun, as 1228 to 160000; but this Gravity is to the Gravity of the Sun towards Saturn, as 160000 to 67, 5.* wherefore the * 1207 mutual Access, or Approach of the Sun and Saturn 1255 is to the Encrease of this Approach by the Action of 1271 Jupiter interpo/ed, as 160067 to 1228, or, as 130 to I.

ThisDifturbance is remarkable, and far the greateft of any that happens in the Motion of any of the primary Planets: This alfo obtains only in this one Cafe of the Conjunction; for when Jupiter recedes from Saturn, the Difturbance of the Motion of Saturn, in a fhort Time, becomes infenfible.

In the fame Position of Jupiter, when it is nearest to Saturn, the Force of Saturn, although it be the greatest of all in this Case, does not so fensibly alter the Way of Jupiter about the Sun. The Action of Saturn, attracting Jupiter, is to its Action by which it attracts the Sun, as \$1 to 16;* therefore it attracts Jupiter with greater *1208 Celerity; and fince they are both attracted the fame

fame Way, the Difference of these Forces is the Force with which (from the Action of Saturn) *175 Jupiter and the Sun are separated from each other*; which is therefore to the Gravity of the Sun towards Saturn, as 65 to 16; but this Gravity of the Sun towards Saturn is to the Gravity of Jupiter towards the Sun, as 4, 223. to *1207 10000*, and as 25 to 81*; that is, as 106 to 1255 810000, or as 16 to 122756; therefore the di-*1208 sturing Force of Saturn is to the Gravity of Jupiter towards the Sun, as 65 to 122756, or 1272 as 1 to 1888; therefore by the greatest Action of Saturn, the Gravity of Jupiter towards the Sun is

diminished only by TST Part, which Difturbance is infentible. The other mutual Difturbances of the Planets

are much lefs, as will appear by determining that, which is the greateft of them all, (viz.) that of Mars by Jupiter, which is differed by the fame Sort of Computation as the foregoing.

The Distance of Jupiter from Mars and the Sun, when Mars is between the Sun and Jupiter *961 in the fame Line, are as 7 to 10*; wherefore 962 the Forces with which Jupiter attracts thefe *1280 Bodies, are as 100 to 49*, the Difference of which Forces is to the Gravity of the Sun towards Jupiter, as 51 to 49. This Gravity of the Sun towards Fupiter, is to the Gravity of *1207 Mars towards the Sun, as 9, 248 to 1000*, and 1255 as 9 to 100*; that is, as 83 to 1000000; or as *1208 49 to 590443; and the difturbing Force of 7upiter is to the Gravity of Mars towards the Sun. as 51 to 590443; or as I to 11577 : wherefore 1273 the Gravity of Mars towards the Sun, is diminished only 11377 Part by the Action of Tupiter when neareft to it.

Although

Although these Disturbances, arising from the 1274 Actions of the Planets upon each other, be very fmall, and although those which happen in a differrent Position of the Planets do in fome Measure compensate each other, yet the Proportion in which the Force, which keeps the Planets in their Orbits decreases, is a little changed by these Actions, fo that it does not decrease exactly in an inverse Ratio of the Square of the Distance : Therefore although the Orbits are at Rest as to Sense, after a great many Revolutions, a small Change is ob-*243 ferved in their Situation.*

From all this it follows, that if we fuppofe the ¹²⁷⁵ Planets at first once projected at the Distances from the Sun at which they are moved, they will, by the Laws already explained, perfevere in those Motions; and the Excentricity of the Orbits depends upon the Celerity and Direction of the first Projection. But these Motions may be preferved very long, by reason of the finall Resistance of the Celestial Matter.

It is alfo plain, why, by Lines drawn to the Center of the Sun, they defcribe Areas proportionable to the Times ; namely, becaufe all other Gravities, in the Syftem, are very fmall in refpect to the Gravity towards the Sun*; therefore by this *1265 Gravity alone, it is that they are retained in their Orbits ; whence follows this Proportion of the Areas*. And alfo the Motion in Elliptic Lines, *225 which are carried on very flowly, follows from the Law of Gravity; and thefe Lines would also be immoveable, if the Planets gravitated only towards the Sun *; but this flow Motion of the *241 Orbits is deduced from the Action of the Pla- 1208 nets upon one another.* Now in respect to the *1214 Proportion which is observed between the Cubes of the Diftances, and the Squares of the Periodical Times, it is also deduced from the Law of Gravity

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239 Gravity ; fo that if we add to thefe what we 1208 have faid of the Deflection of Saturn,* nothing 1276 will remain to be explained in respect to the primary Planets.

1276 That the Motion of Comets depends upon the Law of Gravity, is alfo deduced from Obfervations; and in refpect of them, as has been faid concerning the Planets, the Sun's Gravity prevails, and by that Gravity they deflect from a rectilinear *980 Courfe;* but that the Curvature of their Way 226 depends upon the fame Gravity, follows from this; That a Body, by that Gravity, will definite an Ellipfe, or a Parabola, or an Hyper-*241 bola;* which Lines it appears that those Comets have been determined.

1277 The Satellites of Jupiter and Saturn are moved *944 by the fame Laws about their Primaries, as the 966 Primaries are moved about the Sun;* wherefore the 974 Explication of those Motions * may be also refer-

*1275 red to them; for in these three Cases, smaller Bodies are revolving at different Distances about a much greater Body: Namely, Satellites about Jupiter and Saturn and Frimary Planets about the Sun.

1278 Whilft fecondary Planets are moved about a Primary one, it is evident that they may all be moved with one common Motion, whereby the refpective Motions with which they are moved, in refpect of each other, will not be diffurbed, becaufe a Body may at the fame Time be moved by different Im-

*125 preffions : * The Motion that a Primary Planet has, in common with its Satellites, is the Motion of a Primary Planet about the Sun.

1279 Yet the Motion of the fecondary Planets is diffurbed by the Action of the Sun, towards which they are carried fometimes faster, and fometimes flower, according to the different Position of the Primary; and they

they also often concur towards the Sun's Center in different Directions. These Irregularities, which are very small, cannot be observed in the Satellites of *Jupiter* and *Saturn*, though they be really like those which are observed in the Motion of the Moon: The least Deviation of this last is very fensible to us. But that the Irregularities of the Moon exactly follow from the Theory of Gravity, will appear in the next Chapter.

CHAP. XVI.

The Physical Explication of the Moon's Motion.

TT is certain that the Moon and Earth having 1280 once a projectile Motion given them, they can perceive in their Motion about their common Center of Gravity ;* if they be carried any Way *1263 by a common Impression directed in parallel Lines, as was faid of the Satellites Jupiter and Saturn,* *1278 this Motion will not diffurb the Motion about the common Center of Gravity, which will follow that Direction only, because in respect of the two Bodies it is at Reft. But the Bodies are carried by a Motion compounded of that Impreffion, and of the Motion about the common Center of Gravity ; * that is, they are whirled *199 about that Center as it is carried along, as before its Motion when it was at Reft. If every Moment new Impreffions common to both the Bodies act upon them, the Way of the Center of Gravity may be changed every Moment, which Change will be like that which the Bodies themfelves would undergo, if they had no refpective Motion.

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Hence we deduce, that if whilft the Moon and Earth are whiled round their common Center of Gravity, they be both projected, the Way of the Center of Gravity, by the Action of the Sun acting upon both Bodies, is the fame as a Body, projected in the fame Manner, would defcribe about the Sun.

1281 Whence it follows, that the Moon diffurbs the Motion of the Earth, and that the common Center of Gravity of those Bodies describes that Orbit about the Sun, which we have hitherto said, that the Earth described; because we looked the Action

- 1282 of the Moon; but the Earth describes an irregular Curve.
- 1283 Fig. 4. Plate XXIV.] Let the Sun beat S; and the common Center of Gravity of the Moon Q, and the Earth at M, at the Time of the full Moon, be at F: After one whole Lunation, that is, the next full Moon, let that Center be at A; and let F D A be the Orbit which we call that of the Earth; but in which it is the Center of Gravity above-mentioned that does really move.

If this Lunation be divided into four equal Parts, after the first, the Center of Gravity will be at E, the Moon at P, and the Earth at L; after the fecond Part of the Time, at New-Moon, the Center of Gravity will be at D, the Moon at R, and the Earth at I; in the following Quadrature, the Center of Gravity will be at B, the Moon at O, and the Earth at H. Laftly, at Full-Moon the Center of Gravity being at A, the Moon will be at N, the Earth at G: All which follows from the Revolution of the Earth and Moon about their common Center of Gravity, whilf it is carried in the Orbit about the Sun.

Therefore we fee the Earth moves in the Curve MLIHG, which is twice inflected in each

each Lunation; which Curve alfo does not return into itfelf, becaufe the Inflections in the feveral Revolutions about the Sun do not coincide; for 12 Lunations, and a third Part of another, are performed every Year.

This Irregularity of the Motion of the Earth, 1284, which is deduced from the Laws of Nature, is too fmall to become fenfible in AftronomicalObfervations; wherefore we may without any Error fay, that the Center of the Earth itfelf defcribes the Orbit F D A; for M F, or D I, the greateft Diftance of the Earth from that Orbit, is about the 40th Part of the Diftance I R, which Diftance itfelf is not the 300th Part of the Diftance F S.

In explaining what relates to the Moon, we alfo 1285 neglect the Confideration of the Motion of the Earth about the common Center of Gravity above-mentioned, but we fuppofe it to revolve at the Diftance of 60 Semidiameters from the Center of the Earth; becaufe, as we have before demonstrated,* fuch is the Di-*1253 ftance at which, in its periodical Time, it could revolve about the Earth at Reft, or be carried along in an Orbit in which it fhould not be difturbed by the Moon's Action. By this Method, the Moon's Irregularities will be much more eafily difcovered; for they are the fame; as is evident, whether the Moon moves about the common Center of Gravity of the Moon and Earth, or about the Center of the Earth itfelf.

Plate XXV. Fig. 1.] Let S be the Sun; T the 1286 Earth; and the Orbit of the Moon ALBI; and laft of all, let the Moon be at A in the Quadrature; it tends towards the Sun in the Direction A S, in the fame Manner, and with the fame Force that the Earth is carried 'towards S along T S, becaufe the Diftances A S and T S are equal: This Force may be reprefented by T S or A S, R 2 whereby

whereby the Moon endeavours to defcend along AS, and is refolved into two Forces, by drawing the Farallelogram ADST; fo that the Moon will endeavour to move in the Directions AD and AT, by Forces reprefented by those *192 Lines.*

By the Force which acts along AD, the Moon is carried with the fame Celerity and the fame Way of the Earth; by reafon of the equal and parallel Lines TS and AD: Wherefore by this Motion, the Relation between the Moon and Earth is not changed; but the Force along AT 1287 confpires with the Gravity of the Moon towards the Earth, and this Gravity is encreased by the Attion of the Sun, when the Moon is in the Quadratures; and the Augmentation or Addition is to the Gravity of the Earth towards the Sun, as A T the Moon's Diftance from the Earth is to TS, the Earth's Difance from the Sun.

1288 TS the Earth's Distance from the Sun remaining the fame, the above-mentioned Addition of Gravity encreases, and diminiscont the Ratio of the Line AT, the Distance of the Moon from the Earth.

But this Diftance of the Moon from the Earth A T, if it remains the fame, and T S be encreafed, then AT will be lefs in refpect of AS: Therefore though there should be no Change in the Force, whereby the Earth and Moon fall towards the Sun, the Addition will be lefs, and fo much lefs as T S is greater; that is, it will be inverfly as TS; but the Force of Gravity does not remain the fame, when TS is encreafed, but is diminished ; wherefore alfo in that refpect, the above-mentioned Addition is diminished, and in the fame Ratio with that Force of Gravity; therefore it is the inverse Ratio of the Square of the *1208 Diftance TS;* if this Diminution be added to that above-mentioned, we fee that the Addition of which

which we fpeak, follows the inverse Ratio of the 1289 Cube of the Distance of the Earth from the Sun.

The Diftance of the Earth from the Sun remain-1290 ing the fame, the Gravity of the Moon towards the Earth decreafes more flowly in the Quadratures, than according to the inverse Ratio of the Square of the Distance from the Center of the Earth; for if the Addition in that Cafe should follow the inverse Ratio of the Square of the Distance, as the Gravity from the Action of the Earth does,* this *1208 Ratio would not be disturbed; but the Addition encreafes when the Gravity itself is diminished; wherefore the Addition, when the Distance is encreafed, is always greater than is required, and confequently the Diminution the lefs.

This Addition is determined in the mean Distan-1201 ces of the Moon from the Earth, and of the Earth from the Sun: Let AT and TS be thefe mean Diftances ; the Addition required is to the Gravity of the Earth towards the Sun, as A T to T S*;*1287 the Gravity of the Earth towards the Sun, is to the Gravity of the Moontowards the Earth (becaufe these Bodies are retained by these Gravities in their Orbits) directly as T S to T A, and inverfly as the Squares of the Periodical Times of the Earth about the Sun, and of the Moon about the Earth :**237 Therefore the Addition required is to the Gravity 1236 of the Moon towards the Earth, in a Ratio compounded of these Ratio's; that is, the abovementioned inverse Ratio of the periodical Times of the Earth and Moon, the other Ratio's deftroying one another. These Times are given, and their Squares are inverfly as I to 178,73.

Now if the Moon be at L, [Plate XXV. 1292 Fig. 1.] in which Situation the Sun attracts the Moon and Earth in the fame Line, but not equally; the Moon it draws with a greater Force, becaufe it is lefs diftant from it: The Difference of R 3 those those Forces is the Force by which the Moon is drawn back from the Earth, and by which the Gravity of the Moon towards the Earth is diminisched.

The Forces whereby the Moon at L, and the Earth at T, tend towards the Sun, are to one "1208 another as the Squares of the Lines S T and SL :* and the Difference of the Forces, that is, the difturbing Force, is to the Force by which the Earth defcends toward the Sun, as the Difference of those Squares to the Square of the Line L S; that is, nearly as double L T to L S or TS; for 1293 these Lines do but very little differ from one another; and the Difference of the Squares, whose Roots differ but little, is keeping the Proportion double that which is between the Roots.

If therefore TS, as before, reprefents the Force whereby the Earth defeends towards the Sun, L1 will reprefent the difturbing Force and diminifhing Gravity, when A T reprefents the diffurb-* 1286 ing Force in the Quadratures.* Let the Moon be 1294 at 1; it is again (together with the Earth) attracted by the Sun in the fame Line ; but the Earth, because less distant, moves more fwiftly to-*1208 wards the Sun; * fo that there is a Force which feparates the Earth from the Moon, namely, the Difference of the Forces which attract the Moon and the Earth; which Force always acts contrary to the Gravity of the Moon towards the Earth, and diminishes it, in the fame Manner as has been demonstrated from the greater Gravity of the Moon towards the Sun, fuppoling it at L. At / alfo the feparating Force fearce differs from the feparating Force at L; for this Force, as we have fhewn, is proportional to the Difference of the Squares of the Lines TS and LS; and that (as appears by fuch another Demonstration) proportional to the

the Difference of the Squares of the Lines l S and T S; which Difference, by reafon that L lis very fmall in refpect of TS, fcarce differ among themfelves; fo that the Force which diminifhes the Moon's Gravity at l, is alfo reprefented by L l.

Yet the Diffurbing Force is fomething greater at 1295 the Conjunction L, than at the Opposition 1; for fupposing the Differences between the Roots to be equal, the Squares, keeping the Proportion, will differ fo much the more, the lefs they are; and fo keeping the Proportion, the Forces differ more at L and T, than at T and 1 which also are lefs.*

From this we conclude, that the Force which 1296 diminifies the Gravity of the Moon in the Syzygies, is double that which encreafes it in the Quadratures; namely, as Ll to AT. Wherefore in the Syzygies, the Gravity of the Moon from the Action of the Sun is diminified by a Part, which is to the whole Gravity, as 1 to 89, 36; for in the Quadratures, the Addition of Gravity is to the Gravity, as 1 to 178,73.*

In the Syzygies, the difturbing Force follows the 1297 fame Proportion with half added to it; that is, with the difturbing Force in the Quadratures*; *1296 it is therefore directly as the Diftance of the Moon from the Earth,* and inversity as the Cube of the *1288 Diftance of the Earth from the Sun.*

At the Syzygies the Gravity of the Moon towards 1298 the Earth, receding from its Center, is more diminifhed, than according to the inverse Ratio of the Square of the Distance from that Center; for it would be diminished in that Ratio, if the Force to be taken away followed the same Ratio; but on the contrary, as it encreases when the Distance becomes greater, * the Diminution is always *1297 greater than in that Ratio.

R.4

Plate

1200 Plate XXV. Fig. 1.] Laftly, let the Moon be at F, in any intermediate Place between the Syzygy and Quadrature, it is drawn towards the Sun along FS; by which fince it is lefs diftant than the Earth T, it is attracted with more Force than the Earth : Let the Force, with which the Moon tends to the Sun, be to the Force, with which the Earth is carried towards it, as F M to T S, which also before has been made use of to express the fame Gravity of the Earth. Draw the Parallelogram FHMI, whole Diagonal is F M. and whofe Side F H is parallel and equal to the Line TS. The Motion of the Moon towards the Sun is refolved into two Motions, one along FH, the other along FI, and thefe Lines denote the Forces whereby the Moon endeavours *192 to move along them.* The Motion along FH is common to the Moon and the Earth, which with an equal Force, and in a Line parallel to it, does alfo tend to the Sun : So that by this Motion of the Moon, the Situation of it, in respect of the Earth, will not be changed, and the diffurbing Force will be only the Motion along FI.

By reafon of the immenfe Diftance of the Sun, the Part MS of the Line MF is fmall, in refpect of the Whole, and the Angle FST, where it is the greateft, as AST, is hardly more than the 6th Part of a Degree: Whence it follows, that the Lines M I and SN are very near one another, and that the Points I and N are fcarce fenfibly diftant, and may without any fenfible Error be confounded together; which Error, notwithftanding how little foever it need be regarded, in Confideration of one whole Revolution, is compenfated by a contrary Error when the Moon is at E. Therefore the diffurbing Force is exprefied by F N.

It is to be obferved, when only the Part E F 1300 of the Line ES is confidered, that it is to be looked upon as parallel to the Line L l, becaufe of the fmall Angle which thefe Lines make. From the Point N, draw NQ perpendicular to the 1301 Line F T, continued if Need be, in which the Moon gravitates towards the Earth ; and let the Rectangular Parallelogram E P N Q be drawn: Let us conceive the Force along F N refolved into two others, acting in the Directions F Q and F P, and reprefented by thefe Lines :* *192 By the Force along FQ, the Force of Gravity is diminished, in the Cafe represented by this Figure; but it is encreafed when the Point Q falls between F and T; but by the Force along F P, the Moon in its Orbit is drawn towards the next Syzygy L, and the Motion of the Moon is accelerated or retarded, according as this Force confpires with, or acts contrary to the Moon's Motion.

Near a Syzygy, the Gravity of the Moon is diminifhed, and the Line FQ, which follows the Proportion of this Diminution, grows lefs, receding from the Syzygy till it vanifhes, at the Diftance of about 54 Degr. 44 Min. from it : At a greater Diftance of the Moon from the Syzygy, Q falls in between F and T; and the Gravity of the Moon towards the Earth is encreafed by the Sun's Action. The Force along FP vanifhes in the Syzygy L; receding from it, it encreafes quite to the Octant, which is the middle Point between the Syzygy and the Quadrature; and then it diminifhes again till it vanifhes quite at B.

Between B and *l*, or *l* and A, the difturbing 1302 Motions are determined in the fame Manner, as in A L B, the opposite inferior Part of the Orbit; at E and F, the Diminution of Gravity is equal,

equal, and in that Pofition it is drawn in the Orbit with an equal Force towards the Syzygy *l*, with that with which at F it is impelled towards the Syzygy L.

1303 Hence it follows, that in the Motion of the Moon from the Syzygy to the Quadrature, between L and B, as also between l and A; the Gravity of the Moon towards the Earth is continually encreased,
1304 and the Moon is continually retarded in its Motion. But in the Motion from the Quadrature to the Syzygy, between B and l, as also between A and L, every Moment the Moon's Gravity is diminisched, and its Motion in its Orbit is accelerated.

You may determine the Forces upon which thefe Effects depend, by comparing them with the known Force, whereby Gravity is encreafed in the *1291 Quadratures,* and which is reprefented by the Moon's Diftance from the Center of the Earth.

1305 The Lines M I, HF, ST, are equal by Construction ; therefore when the Points I and N are confounded, MN is equal to ST, and MS is equal to NT. The Lines MF and S T reprefent the Forces, whereby the Moon at F and the Earth at T are carried towards the Sun S; therefore they are as the Square of the Line TS *1208 to the Square of the Line FS;* wherefore as FG is the Difference of those Lines, FM and T'S differ from one another double the Line *2293 G F,* and adding G F to the Line F M, the Difference between G M and T S, that is, M S will be triple the Line FG; and therefore this is alfo the Quantity of the Line NT : Now as FE *1300 is double FG;* therefore NT will be to FE as Three to Two.

Let F T be continued, if Need be, and from E draw EV perpendicular to it; the Triangles E V F and NQ T, which are rectangular, will be fimilar, by reafon of the alternate Angles V F E

V F E and QTN:* Therefore NT is to F E, *1300 that is, Three is to Two, as NQ, equal to F P, is to EV; which therefore is proportional to two third Parts of the Force, which is expressed by F P; but EV is the Sine of the Angle E TV at the Center, which is double the Angle E F V at the Circumference, equal to the Angle F T L, which is the Distance of the Moon from the Syzygy. Therefore, as the Radius T A, or T E, 1306 is to a Sine and an half of double the Distance of the Moon from the Syzygy, namely, F P, fo the Addition of Gravity in the Quadratures (which is expressed by the Radius T A) is to the Force which accelerates or retards the Moon in its Orbit.

The Computation of this Diminution of Gravity, and of its Encrease at a lefs Distance from the Quadratures, is deduced from the same Principles.

This Diminution is reprefented by the Line FQ, which is equal to Q T, minus the Radius; but from the Confideration of the Triangles above-mentioned, V F taken once and an half is equal to Q T; therefore V T and an half, with half the Radius added to it, expresses the required Diminution of Gravity; and the Radius is to the 1307 Sum or Difference of once and a half the Co-fine of double the Diffance of the Moon from the Syzygy and half the Radius : As the Addition of Gravity in the Quadratures, to the Diminution or Encrease of Gravity in that Situation of the Moon, concerning which the Computation is made.

We make use of the Difference of the Cofine from half the Radius, when the Angle, whose the Co-fine is, is greater than a right Angle; because in that Case we make use of the Co-fine of the Complement of the Angle to two right Angles: When in this fame Case the Co-fine and a half, which we make use of, is greater

greater than half the Radius, the Quantity found is to be added ; that is, encreafes the Gravity, which obtains every where between the Quadrature, and 35 Degr. 16 Min. from it.

1308 Thefe Forces, whatever is the Figure of the Moon's Orbit, are exactly determined; for they are compared with the Addition of Gravity in the Quadratures, fuppofing the Moon in the Quadrature to be at the fame Diftance from the Earth, at which it really is in the Place which is confidered; but this Addition is different in every
1291 Cafe.

Though it be foreign to the Purpole of this Work, to give a Computation of the Moon's Motion, I thought it neceffary to explain in a few Words, what is the Method whereby to difcover the Forces that govern the Moon; because the more exactly we know the Forces, the more easily we shall conceive their general Effect.

Now to examine the Moon's Motion, we must fingly confider its feveral Irregularities; which to do without Confusion, we must remove feveral Irregularities, and conceive the Moon as moving in a Circle about the Earth, in which Curve it is plain, that it can be retain-*247, ed by Gravity.* This Motion is diffurbed by 1208 the Action of the Sun, and the Orbit is more con-1309 vex in the Quadratures than in the Syzygies. The

Convexity of a Curve, which a Body defcribes by a central Force, is fo much the greater, as the central Force does more firongly every Moment turn the Body out of the Way; it is alfo the greater, the more flowly the Body moves; becaufe the central Force, acting the longer, has a greater Effect in inflecting the Way of the Body. From contrary Caufes the Convexity of the Curve is diminifhed. Both concur in encreafing the

the Convexity of the Orbit in the Quadratures, * *1303 * 1304 and diminishing it in the Syzgies.*

From this it follows, that the circular Figure 1310 of the Moon's Orbit is changed into an Oval. whole greater Axis goes through the Quadratures : fo that the more convex Parts are in the Ouadratures. Wherefore the Moon is less distant from the Earth at the Syzygies, and more at the Quadratures, and it is no Wonder that the Moon comes towards the Earth, when its Gravity is diminished ; because the Access is not the immediate Effect of this Diminution, but of the Inflexion of the Orbit towards the Quadratures.

The Motion of the Moon, taking away the Action of the Sun, is not in a Circle, but in an Ellipfe, one of whole Foci coincides with the Center of the Earth; * for the Orbit of the *967, Moon is Excentric, and it is retained in it by the Force of Gravity.

Therefore what has been demonstrated cannot be exactly applied to the Moon's Motion ; for as the Forces, which generate the Deviations explained, do really act upon the Moon, the Ellipfe, which the Moon would defcribe if the Sun was taken away, is changed, and cæteris paribus, 1311 the Propositions of Nº 1309, 1310, may be applied to the Moon's Motion.

In the Quadratures and Syzygies, the diffurbing 1312 Force acts in the fame Line as the Force of Gravity towards the Earth ;* therefore the Force which *1286 continually acts upon the Moon, and retains it 1292, 1294 in its Orbit, is directed towards the Center of the Earth, and the Moon describes Area's by Lines drawn to the Center, of the Earth proportional to the Times. * *215

Plate XXV. Fig. 1.] In other Points of the Or- 1313 bit. as F, befides the Force which acts in the Line F T, there is also another, whose Direction

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¹³⁰¹ ction is perpendicular to FT, * which is here reprefented by F P: The Direction of the Force compounded of both, is directed fometimes fidewife to the Line F T, and does not tend to the

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190 Center of the Earth ; wherefore the Area's, by Lines drawn to the Center of the Earth, are not ex-

226 attly proportional to the Times. In the Octants, F P is the greateft of all; and the Force, which is reprefented by that Line, is to the Gravity of the Moon towards the Earth, in that Point, in the Mean Diftances of the Sun and Moon, as I to *1306 I19,49.* Wherefore the Direction of the Force,

¹³⁰⁷ compounded of the Actions of the Sun and Earth upon the Moon, makes an Angle of above half a Degree with the Line FT.

The Motion of the Moon is fubject to feveral other Irregularities; fo that it defcribes a Curve wholly irregular; which Aftronomers, in order to fubject it to the most exact Computations 1314 that can be made, do reduce to an Ellipse which they conceive to be agitated by various Motions, and also to be changeable, left the Moon should go out of it.

In refpect to central Forces we have observed. that a Body does not defcribe an Ellipfe, if the central Force, by which it is retained in its Orbit, decreafes in any other Ratio than the Inverse Ratio of the Square of the Diftance; but that the Curve may be often reduced to a moveable El-*2.43 lipfe : * Concerning which it is to be obferv-1315 ed, that the Ellipse, in that Motion, turns about 1316 one of its Foci, and the Motion of the Elliple is directed the fame Way as the Motion of the Body in it, when the central Force decreases faster than in the Inverse Ratio of the Square of the Distance. But if the central Force decreases slower as you recede from the Center, the Ellipfe is carried the contrary Way; as these Things may be demonstrated Mathematically. Hence

Hence it follows, the Orbit of the Moon cannot be referred to an Elliptic Orbit, unlefs you suppose it agitated by four Motions every Revolution; that is, unlefs the Line of the Apfides, which goes through the Center of the Earth, goes forwards twice, and backwards twice.

The Apfides of the Moon go forward when the 1317 Moon is in the Syzygies,* or rather whilft the *1315 Moon moves between the Points, which are 54 1298 Degr. 44 Min. diftant from them. * In the Qua. *1307 dratures, and between the Points diftant from them 35 Degr. 16 Min. The Apfides go backwards, 1318 that is, move in Antecedentia.* *1316

The Forces, upon which the Progress and Regress of the Apsides depend, are the Forces which diffurb the Motion of the Moon, which have 1319 been before explained; therefore, fince the diffurbing Force in the Syzygies, is double the diffurbing Force in the Quadratures; the Frogress, *1296 confidering one entire Revolution of the Moon, exceeds the Regress, cæteris paribus.

In a Circle whole Center is in the Center of the Forces, the Diminution of the Force, in receding from the Center, produces no Effect; becaufe in fuch a Line the Body does not recede from the Center; therefore the Effect of this Diminution is fo much the greater, as the Curve defcribed by the Body differs more from fuch a Circle.

In an Elliptic Orbit, one of whole Foci co-1320 incides with the Center of the Forces, the Curvature in the Apfides differs most of all from fuch a Circle, and the Effect of the Diminution of the Force, in receding from the Center of the Forces, is the greatest of all. If this Orbit is but a little 1321 Excentric at the Ends of the leffer Axis, the Ellipfe differs very little from the Circle above-mentioned, and the Effect of the Diminution is the least of all. The

1322 The Progress and Regress of the Apsides depend upon the Proportion, according to which the Force of Gravity decreases, receding from the *1315 Center of the Earth ;* it is therefore the Effect of 1316 the Diminution of the Central Force.

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1323 Plate XXV. Fig. 2.] This Motion of the Apfides, which we have explained, undergoes feveral Changes; the Apfides go forward fafteft of all in a Revolution of the Moon, fuppofing the Line
*1317 of the Apfides in the Quadratures; * and in that 1322 very Cafe, they go back the floweft of all in the fame
*1318 Revolution, * becaufe, by reafon of the fame Ex-1321 centricity of the Moon, the Quadratures are but 1322 very little diftant from the Ends of the leffer Axis of the Orbit.

1324 Supposing the Line of the Apsides to be in the Qua-*1317 dratures, the Apsides are carried in Confequentia, 1321 the least of all in the Syzygies;* but they return the *1318 fwiftest in the Quadratures; * and in this Case, *1322 in one entire Revolution of the Moon, the Regress 1320 exceeds the Progres.

1325 Whilft the Earth is carried along in its Orbit, the Line of the Apfides does fucceffively go through all Situations in respect of the Sun; wherefore, confidering a great many Revolutions of the Moon

1319 taken together, the Apfides go forwards; and it is plain from Obfervations, that in the Space of about eight Years, the Line of the Apfides performs one entire Revolution.

We have also faid that the Excentricity of the Orbit is not constant.

1326 The Excentricity of a Body is encreased, if the central Force, the Diminution being continued, decreases faster than before, whilft the Body is carried from the lower to the upper Apsis; for then it is every Moment less attracted, than if the Force did not decrease; and therefore it recedes the more. The Excentricity of the Orbit is also thereby encreased, in the same Case, in the

the Motion from the upper to the lower Aplis, becaufe in this Cafe, coming towards the Center, the Force encreases fo much the faster, as the Body defcends more towards the Center; fo that in each Cafe, the Difference between the greateft and least Distance from the Center of the Forces, may become greater, and the Excentricity be thereby encreafed. By the fame 1327 Reafoning it appears, that the Excentricity is diminished, when the Central Force is encreased, in the Motion of the Body from the lower to the upper Apfis; and likewife when that Force is diminished in the Motion from the upper Apsis to the lower; that is, when it decreases more flowly than before, in the receding from the Center.

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Applying this to the Moon's Motion, it ap-1328 pears, that the Excentricity of the Orbit, every Revolution, undergoes various Changes. That it is 1329 the greatest of all, when the Line of the Apsides is in the Syzygies; becaufe the Forces in the Apfides. being compared, do decrease faster than in an inverse Ratio of the Square of the Diftance,* *1298 whence this Addition follows,* which prevails *1326 in this Polition :* But the Orbit is the least Ex-*1320 centric of all, when the Line of the Apfides is in the Quadratures, the Diminution of the Excentricity prevailing.* *1290

We have faid that the Moon moves in a Plane 1327 inclined to the Plane of the Ecliptic ; that the 1321 Line of the Nodes is carried round in Antecedentia:* and that the Inclination of the Orbit is *970 not conftant ;* thefe Effects are also deduced from *969 the Action of the Sun upon the Moon.

By reafon of the fmall Inclination of the Moon's Orbit, the Forces which we have hitherto confidered acting in the Plane of the Ecliptic, not regarding the Inclination of the Orbit, may (without any fenfible Error) be referred to the Plane of

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of the Orbit; and the Moon in it is fubject to the 1330 Motions before explained: But there is a Force which removes the Moon from the Plane of the Or-

1314 bit; fo that we must conceive that Plane to be agitated, otherwife the Moon would go out of the Orbit.

- 1331 Plate XXV. Fig. 1.] Let the Moon be at F; if we attend to what has been faid above con-
- *1299 cerning the Actions of the Sun,* it is plain, that the Plane of the Parallelogram FHM I goes through the Line T S, which joins the Centers of the Earth and Sun; and therefore it is in the Plane of the Ecliptic; fo that the Point N, to which is directed the Force FN, diffurbing on account of the Action of the Sun, is in that Plane.
- 1332 Plate XXV. Fig. 4.] Let the fame Force be reprefented by F1; at F let F R be raifed perpendicular to the Plane of the Orbit; and imagine the Parallelogram F R 1*i*, whofe Side F*i* is in the Plane of the Orbit, and whofe Diagonal is F1; the diffurbing Force along F1 is refolved into two, in the Directions F R and F*i* which *192 thefe Lines reprefent,* and of which this laft acts in the Plane of the Orbit : So that we must refer to this what relates to the diffurbing Force, of which we have treated in N° 1299; for the Lines F*i* and F1 fcarce differ, and the Plane of the Moon's Orbit.

1333 'The Line F R muft be determined, which reprefents the Force that acts perpendicular to the Plane of the Orbit, and removes the Moon from that Plane: Now the Relation of the Line FR or Ii, to the Radius E'T, is the Ratio of the diffurbing Force, which is fpoken of here, *12?6 to the Encrease of Gravity in the Quadratures.* In

I

In the Cafe of this Figure, in which the Line 1334 of Nodes Nn is in the Quadratures, we find out FR ; becaufe I T (which is N T of Fig. 1.) is given,* and becaufe I T is to Ii, or to F R, as *1325 the Radius to the Sine of the Inclination of the Orbit.

But in every Cafe the Force must be determin-1335 ed which drives the Moon out of the Plane; let us therefore fuppose the Line of Nodes carried to the Situation M m, whereby, every thing elfe remaining as before, I i is changed. To M m continued, if Need be, let i X and I X be drawn perpendicular, which make an Angle equal to the Inclination of the Plane of the Orbit.

The Ratio between ET and Ii; that is, the 1336 Ratio between the Addition of Gravity in the Quadratures, and the Force, which we feek, which removes the Moon out of its Orbit, is compounded of the Ratio's of the Line ET to TI, of the Line TI to IX, and laftly of the Line IX to I i. The first is the Ratio between the Radius and three Times the Sine of the Diftance of the Moon from the Quadrature ;* the fecond is the Ratio of the * 1305 Radius to the Sine of the Angle I TX, that is, of the Diftance of the Node from the Syzygy. Laftly, the third is the Ratio of the Radius to the Sine of the Inclination of the Orbit: And the Ratio, compounded of thefe, is the Ratio of the Cube of the Radius to three times the Product of the Sines of the Distances of the Moon from the Quadrature, and of the Node from the Syzygy, as allo of the Inclination of the Plane. To this Force is alfo to be referred Nº 1308.

This Force vanishes in the Quadratures, because 1337 the Point I coincides with the Point T, which is the Center of the Earth, and the Line Ii vanifhes; the Lines Fl and Fi concurring in the Plane of the Orbit, which also follows from the Computation above-mentioned ;* the Sine of * 1136 VOL. II S 2

the

the Diftance of the Moon from the Quadrature vanifhing, and confequently, the whole Product which is multiplied by that Sine.

- 1331 That fame Product vanishes alfo, and with it the Force which it reprefents, when the Sine of the Distance of the Node from the Syzygy vanishes, that is, supposing the Line of the Nodes in the Syzygies : It is also deduced from this, that the Line of Nodes N n (Plate XXV. Fig. 5.) continued goes through the Sun, wherefore the Sun is in the Plane of the Orbit itself, and therefore cannot draw the Moon but in that Plane.
- 1339 The Force alfo, which we examine, is increased as the Moon advances towards the Syzygy, and as
 1336 the Node recedes from it.
- 1340 Plate XXV. Fig 6.] Let P p be the Plane of the Ecliptic, P A the Orbit of the Moon; when the Moon is come to A, that is, is receded a little from the Node, it is removed out of the Plane of the Orbit, and in the fecond Moment it is not carried along AB (the Continuation of the Orbit P A) but along Ab; because it comes towards the Plane of the Ecliptic along Bb; therefore it is moved as if it came from a more di-
- 1341 ftant Node p. Whence it appears, that the Nodes go backward, whilf the Moon moves in its Orbit, as long as it recedes from the Node: The Nodes alfo go back, whilf the Moon is going to the oppolite Node; becaufe as the Moon is continually driven out of its Orbit towards the Plane of the Ecliptic, it is continually directed to a Point lefs diftant, and comes fooner to the Node, than if, not being agitated by fuch a Motion, it had continued in Motion with the fame Celerity.

1342 Confidering one entire Revolution of the Moon cæteris paribus, the Nodes move in Antecedentia
* 1339 fwiftest of all, when the Moon is in the Syzygies,* then flower and flower, till they are at Rest, when
* 1337 the Moon is in the Quadratures. Whilft

Whilft the Earth is carried round the Sun (even when we do not attend to the above-mentioned Motion of the Nodes) the Line of Nodes 1343 does fucceffively acquire all poffible Situations in refpect of the Sun; and every Tear goes twice through the Syzygies, and twice thro' the Quadratures

If now we confider feveral Revolutions of the 1344 Moon, the Nodes in one whole Revolution go back very fast, the Nodes being in the Quadratures ;* * 1339 then flower, till they come to rest, when the Line of Nodes is in the Syzygies.* * 1339

By the fame Force with which the Nodes are 1345 moved, the Inclination of the Orbit is also changed; it is increased as the Moon recedes from the Node; and diminisched as it comes to the Node.

Plate XXV. Fig. 6.] For the Angle bpL is 1346 lefs than the Angle A P L, and for the fame Caufe it is continually diminished, and the Inclination becomes greater ; but when the Moon is come to the greatest Distance from the Plane of the Ecliptic, and is going towards the oppofite Node, the Direction of the Moon is continually inflected towards the Plane of the Ecliptic, and lefs inclined to it than if it continued in Motion in its Orbit : Let N pn be the Plane of the Ecliptic; the Curve Nn the Orbit of the Moon ; by the Force whereby the Moon is continually removed out of it, the Way of the Moon is changed, and it goes in the Curve N p, which is more inclined to Npn at N than at p; fo that we must conceive the Inclination of the Plane of the Orbit to be twice changed ;* whilft the Moon *13'4 moves from one Node to the other; therefore this happens four times in each Revolution of the Moon, 1347 it is twice diminished, and twice again encreased.

Plate XXV. Fig. 4.] Supposing the Nodes N n, 1348 to be in the Quadratures, the Forces which in one Revolution encrease the Inclination, and diminish it, are equal to one another; for by reason of the S 3 equal

equal Diffance of each Node from the Syzygies, the Forces that change the Inclination at N D and n E are equal to the Forces in the correspon-*1336 dent Points in the Arcs D n and EN :* By the former the Inclination is encreased, by the latter *1345 it is diminished ;* the Diminution of the Angle of Inclination, on account of the first, is reftored by the Action of the fecond, and here it is not *1343 changed. In the Motion above-mentioned* of the Line of Nodes in respect to the Sun, which depends upon the parallel Situation of this Line, the Node N is carried to the Syzygy E. When (for Ex.) the Line of Nodes is come to the Situation M m, the Moon in its Receis from the Nodes goes through the Quadratures Nn, in which the Force which changes the Inclination *1337 vanifhes,* and near which it is the leaft of all :* *1326 but in coming towards the Nodes, the Moon is every where diftant from the Quadratures, and *1336 a greater Force acts upon it ;* therefore con-1349 fidering one entire Revolution, the Encrease of the *1345 Angle of Inclination exceeds its Diminution ;* that is, that Angle is encreafed, or which is the fame, the Inclination is diminished; which obtains

every where in the Motion of the Nodes from the Quadratures to the Syzygies.

1350 When the Nodes are come to the Syzygies, the In1351 clination of the Plane of the Orbit is the leaft of all; for in the Motion of the Nodes from the Syzygies to the Quadratures, the Plane of the Orbit is continually more and more inclined; for in that Cafe as the Moon goes to the Node, it paffes through the Quadratures; in its Receifs from them the Moon is diffant from the Quadratures, and in one whole Revolution of the Moon, the Force which
*1337 encreafes the Inclination * exceeds that which
1341 diminifies it 3 therefore the Inclination is en1352 creafed; and it is the greateft of all when the Nodes are in the Quadratures, where the Diminution of the

the Angle made by the Plane of the Orbit with the Plane of the Ecliptic is terminated.* *1349

All the Errors in the Moon's Motion, that we have 1353 explained, are fomething greater in the Conjunction than in the Opposition.*

All the diffurbing Forces are determined by dif-1354 covering their Relation with the Addition of Gravity in the Ouadratures ;* for which Reafon *1306 1307 they all undergo the fame Changes as that Ad-1336 dition does; that is, they are inverfly, as the Cube of the Distance of the Sun from the Earth, * *1289 which when it remains the fame, they are as the *1288 Distance of the Moon from the Earth.* Confidering all the disturbing Forces together, the Diminu-1355 tion of Gravity prevails ;* which follows immedi-*1296 ately from the Progress of the Apfides ; for it ap-*970 pears from thence, that confidering feveral Revolu-1325 tions together, the Effect of the Diminution of *1317 Gravity exceeds the Effect of the Encrease of it.* 1318

Therefore the Motion of the Moon being confider-1356 edin general, the Gravity of the Moon towards the Earth is diminifhed, coming near the Sun 3* there-*1355 fore when it is lefs attracted by the Earth, it recedes more from it than it would recede, if there was no fuch Diminution of Gravity: Therefore in that Cafe the Moon's Diftance is encreafed, as 1357 alfo the Periodical Time 3* and that Time is the *224 greateft, as alfo the Diftance of the Moon(cæteris paribus) the greateft, when the Earth is in the Peribelion, **1354 becaufe then it is leaft diftant from the Sun.

CHAP. XVII.

Concerning the Figures of the Planets.

I F we confider the Figures of the Planets, we fhall find that they have fuch Figures, which follow from these very Laws by which

the

the Syftem is governed; which is very agreeable to that admirable Order which we observe every where, that no Forces act upon the Planets to 1358 deftroy them; that is, that the Figure of a Planet, whether it be a Primary or Secondary Planet, is juch as it would acquire, if it wholly confifted of fluid Matter; which agrees with the Phwnomena.

1359 Whence it follows, that all the Primary and Secondary Planets are Spherical; for they confift of a Matter whofe Particles gravitate towards one
1206 another; from which mutual Attraction a Sphe-1207 rical Figure is generated in the fame Manner, as a Drop becomes round from another Sort of At*24 traction of the Parts.*

1360 This Spherical Figure of the Planets is not changed from their Motion round the Sun, or from the Motion of the Secondary Planets about their Primary ones; becaufe all the Particles are carried by the fame Motion: But this Figure undergoes fome Change by the Motion round the Axis, and fo much the greater as this Motion is fwifter.

1361 Plate XXV. Fig. 7.] Let PP be the Axis of a Planet ; E e the Diameter of the Æquator perpendicular to the Axis; let there be a Canal PCE filled with a Liquid; this Fluid will defcend by its Gravity in both Legs towards C, and will not be at reft, till the Preffure in both Legs be equal. If the Planet be at reft, the *1359 Height of the Fluid in both Legs will be equal :* but if the Planet be moved about its Axis Pp, all the Liquid in the Leg CE will endeavour to *217 recede from the Center by its centrifugal Force,* *223 which Force acts contrary to Gravity,* and therefore diminishes the Gravity; fo that there is no Æquilibrium till C E exceeds C P. Now if the Canal be taken away, the lateral Preffure of the Fluid, of which the Planet confifts, does not change the Gravity towards C, nor the Difference

ference between the Heights of the Columns CE, C P;* therefore the Planet is every where higher *280 in the Æquator than in the Poles, and acquires, by its Motion round its Axis, the Figure of a 1362 Spheroid depressed in its Poles; for the Elevation is continually diminished as you go towards the Pole, because the centrifugal Force is diminished by reason of the Diminution of the Distance from the Axis.*

If what has been demonstrated be compared with the Phænomena, it will appear why all the Bodies in our System are spherical;* but that *924 this Figure is not exact, but a little changed by their Motion round their Axis,* though this can-*1362 not be observed in most of them, may be deduced from Observations made upon *Jupiter* and the Earth. Astronomers have observed that the Axis of 1363 Jupiter is shorter than its Æquatorial Diameter; although this Planet be the greatest of all the Planets, it is moved the switch about its Axis;* *962 and therefore this Difference may be observed.

The Elevation of the Earth at the Aquator is 1364. determined by us, although perhaps to the Inhabitants of the other Planets, if there are any, it may not be more fensible than the Elevations on Mars and Venus are to us, which are fo fmall that we cannot perceive them.

Suppose the Earth to be fluid, it will acquire 1365 the aforefaid spheroidical Figure:* If the Parts *1362 cohere towards the Center, the Position of the other Parts will not be changed thereby, nor will it be changed, if in some Places the Parts cohere together quite to the Surface; fo that the Surface of the Sea must necessarily acquire a spheroidical Figure depressed at the Poles. But fince the Shores are every where but a little elevated above the Surface of the Sea, it is certain that the Continent acquires the fame Figure.

Now

Now to measure this Elevation, that is, how much the Diameter of the Æquator of the Earth is longer than the Axis, we must confider its Motion round its Axis in the Space of 23 Hours, *960 56 Min. 4 Sec.* and supposing the Earth homo-

geneous, the Computation will be made in the following Manner.

1366 The Periphery of the Earth is 128,202,185 Rbynland Feet; therefore in one Second of Time, a Point of the Æquator goes through 1488 Feet; the verfed Sine of which Arc is 0,054, a Space which could be gone through by a Body in fuch a Time by the centrifugal Force.

By a Gravity, a Body, in one Second, as we have shewn before, falls through 15, 607 R bynland Feet : but thefe Experiments were made at the Diftance of 48 Degr. from the Æquator E e (Plate XXV. Fig. 7.) at the Point A, the centrifugal Force at E is to the centrifugal Force at A, as C E to C A, for thefe Lines are very little dif-*232 ferent at A B:* Let this centrifugal Force be A b; having drawn the Perpendicular b a to C A continued, let the Force through A b be refolved into two other Forces directed along A a *192 and a b ;* the Gravity is diminished only by the former, and A b is to the Force diminishing it, as C A to A B, by reafon of the fimilar rectangular Triangles Aba and A BC, which have their oppofite vertical Angles equal at A; therefore the centrifugal Force at the Æquator, with which a Body in one Second goes through 0, 054; is to the Force which diminifhes the Gravity at A, in a duplicate Proportion of the Radius AC to A B, which is the Co-fine of the Latitude A E of 48 Degr. So that from this diminifhing Force the Body in one Second goes through 0,0243; wherefore, if the Earth was at reft, in falling it would not go through 15, 607 Feet, but 15, 632; with

with which Gravity a Body falls under the Poles. because these Points are not moved. At the Æquator, by the centrifugal Force, a Body goes through 0, 054. and falls as much in the fame Time from the Height of 15, 578 Feet ; whence it appears that the Gravity under the Poles is, to the Gravity under the Æquator, as 289 to 288.

If Fig. 7. reprefents the Figure of the Earth. the Weight of a Column of Liquid C E will be to the Weight of a Column of Liquid CA, the Earth being at reft, as 289 to 288; for otherwife, the Earth moving, there will not be an Æquilibrinm; because $\frac{1}{2n}$ of the Column CE is furtained by the centrifugal Force; for the centrifugal Force decreafes as you come towards the Center, in the Ratio of the Diftance, * in which Ratio alfo the *232 Gravity decreafes ;* fo that in all the Points of * 1233 the Column, the fame Part of the Weight is fuflained, as towards the Surface.

Whence we deduce, that the Height CP at the 1367 Pole is, to the Height EC at the Æquator, as 229 to 230; for fuppoling this Ratio between the Axis and the Æquatorial Diameter, if a Computation be made of the Gravities in the Places P and E, the Earth being at reft, they are found to be to one another, asII2I, 7I. to II20, 7I; which Ratio obtains every where in correspondent Points, that is, which are diftant from the Center as CP to PE; becaufe in both Legs the Gravity decreafes in Proportion to the Difrance from the Center.* You have the Weight * 1233 by multiplying the Quantity of Matter by the Gravity; for the Weight encreafes in a Ratio of both : By multiplying 1121, 71. by 229, and 1120, 71. by 230. the Products are to one another as 288 to 289; which is the Ratio of the Weights before difcovered. The mean Diameter

of

976 of the Earth is 3,400,669 Perches, therefore the Axis PP is 3,393,261, and the Æquatorial Diameter E e is 3,408,078 Perches, which exceeds
1368 the Axis by 14,817 Perches (viz.) = 30, and the Æquator is more elevated by 7408,5.

1369 In this Computation, as we have faid, we have looked upon the Earth as homogeneous; but if it be more denfe towards the Center, the Matter which is added to it may be looked upon as a feparate Body, from whofe Center the Points P and E are unequally diftant, and towards which therefore the Bodies P and E have a difference is fo much the greater as thefe Differences are greater; and it will be alfo fo much the greater in refpect of the whole Gravity, as the Quantity of Matter which is added, or which is the fame, as the Denfity is greater towards the Center.

It is plain that the Forces of Gravity, at the Poles and the Æquator, differ from one another more than $\frac{1}{289}$ Part, by comparing together Experiments made at feveral Diftances from the Æquator, by the Help of Pendulums, by which the Forces of Gravity may be compared together, *164 as we have fhewn;* and which Difference is truly 165 nearly double that which is found by Computa-1370 tion; whence it follows, that the Elevation of the Æquator is nearly double that which we have de-*1368 termined to be 7408, 5 Perches.*

Now if we confider the fpheroidical Figure 1371 of the Earth, we fhall fee that beavy Bodies do not tend directly to the Earth's Center, unlefs at the Poles and the Æquator, but every where perpendicularly to the Surface of the Spheroid; for a Liquid will not be at reft, unlefs its upper Surface forms a right Angle with the Direction of heavy *272 Bodies; * and the Figure of a Spheroid is formed by the Surface of a quiefcent Fluid. We alfo

alfo deduce this Direction of heavy Bodies from 1372 the centrifugal Force (Plate XXV. Fig. 7.) The Body A, by its Gravity, tends towards C, and is carried by its centrifugal Force along A b: This Force at the Point A is to the Gravity along A C, as I to 430, 8. having formed a Parallelogram with the Sides A c and Ab, fuppoling thefe to one another, as 430, 8. to 1, the Diagonal will fhew the Direction of heavy Bodies,* *190 forming a fmall Angle with the Line AC. The Force along A b encreases as you go towards the Æquator, whereby this Angle is encreafed. but is diminished by the Encrease of the Angle CAb; fo that in the Æquator, where the centrifugal Force is greateft, the Direction of heavy Bodies coincides with EC; at the Pole it coincides with PC; becaufe there is no centrifugal Force there.

In this fpheroidical Figure, the Latitude of the 1373 Place is determined by an Angle, as ACE, which is made with the Æquator, by a Line drawn from the Place of the Center. Dividing this whole Arc PAE, by this Method, into 90 Parts, that is, into Degrees, it will eafily appear, that going 1374 towards the Poles, the Degrees are encreased on the Surface; but this Difference is fo very fmall, that in meafuring Degrees that are not very diftant, it cannot be difcovered ; becaufe the Error, arifing from the Make and Ufe of the Inftruments, exceeds this Difference. Whence Degrees measured at the South and North of France. as alfo in England, differ little from one another, and the middle one is the leaft of all ; wherefore nothing can be concluded concerning the Earth's 1375 Figure from thefe Meafures.

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CHAP. XVIII.

The Physical Explanation of the Motion of the Axis of the Earth.

HAT the Nodes of the Moon go back, *1341 that is, are moved in Antecedentia,* and that the Inclination of its Orbit is liable to *1347 Change,* we have already demonstrated; let us conceive feveral Moons to be at the fame Diftance, revolving in equal Times about the Earth, in a Plane inclined to the Plane of the Ecliptic; it is plain they will all be agitated by the fame Motions : Let us conceive the Number of the Moons to be encreased, so as to touch one another, and form a Ring, whofe Parts cohere; whilft one Part of the Ring is attracted, to encrease the Inclination, the other Part is 1376 agitated by a contrary Motion, to diminish its *1345 Inclination ;* the greater Force in this Cafe pre-1377 vails, that is, in the Motion of the Line of Nodes, from the Quadratures towards the Syzygies, the Inclination of the Ring is diminished in each of its *1349 Revolutions ;* and it is the leaft of all, when the *1350 Line of Nodes is in the Syzygies.* On the contrary, its Inclination is encreased, when the Line of Nodes is carried from the Syzygies towards the Quadra-*1351 tures; * and it is the greatest of all, when the Line of 1378 the Nodes is in these last.* The Line of the. *1352 Nodes is continually carried in Antecedentia, un-1344 lefs when it is at reft in the Syzygies.*

1379 If the Quantity of Matter in the Ring be diminished, its Motions will not be changed; because they depend upon Gravity, which acts equally up-*1027 on every Particle of Matter.*

If the Diameter of the Ring be diminisched, these 1380 Motions are diminisched in a Ratio of this Diminution,* but none of them wholly vanish; and *1354 it is agitated by the same Motions.

Let us conceive the Earth to be fpherical; and 1381 in the Plane of the Æquator, which makes an Angle of 23 Degr. 29 Min. with the Plane of the Ecliptic, let there be a Ring, revolving in the fame Time as the Earth; let it be diminifhed fo as to touch the Earth, and cohere with it; by this the aforefaid Motion of the Ring will not be deftroyed; for fince the Earth is kept in a determinate Situation by no Force, it yields to the Impreffions of the Ring, whofe Agitations are yet diminifhed, the Matter to be moved being encreafed, and the moving Power remaining the fame.

And this is truly the Cafe, for the Figure of the Earth is fpherical, encompafied with a Ring at the Æquator, whereby the Earth is more elevated towards the Æquator,* the Line of Nodes *1370 of which Ring is the Section of the Planes of the Æquator and Ecliptic. Whence we deduce the following Conclusions.

In the Æquinoxes, the Inclination of the Æ-1382 quator is leaft of all;* and therefore the Incli-*1377 nation of the Axis is the greateft; for it makes a right Angle with the Plane of the Æquator.**1075 The Inclination of the Æquator is encreafed, that is, the Inclination of the Axis is diminiscued, till 1383 the Suncomes to the Solftices, where this Inclination of the Axis is least of all, and that of the Æquator the greateft.* Therefore twice in a Tear the In-*1377 clination of the Axis of the Earth is diminiscued, and twice encreased. And the Selftion of the Plane of the 1384 Æquator with the Plane of the Ecliptic, which is 1385 at rest in the Æquinoxes, the rest of the Time is moved in Antecedentia.* *1378

The

Mathematical Elements

1386 The Plane of the Æquator is enclined to the Plane of the Moon's Orbit; for it makes a fmall

Book IV.

969 Angle with the Plane of the Ecliptic : And therefore the Moon acts in the fame Manner upon the Ring as the Sun ; and although the Moon be lefs, yet, becaufe it is much lefs diftant than the Sun, it exerts a greater Action upon the Ring.

1387 Wherefore also the Inclination of the Axis of the *1384 Earth to the Plane of the Moon's Orbit* (and confequently to the Plane of the Ecliptic) is twice changed in every Revolution, and twice reftored by

the Action of the Moon: And the Section of the Plane of the Equator, with the Plane of the Moon's

1385 Orbit, is carried in Antecedentia: From which Motion neceffarily follows, that the Section of the Plane of the Æquator, with the Plane of the Ecliptic, changes its Place.

1388 The Changes of the Inclination of the Axis are too fmall to be observed; but the change of Place of the

1389 Line of the Equinoxes, and the Motion of the Axis which follows from it, being always carried the fame Way, at last become fensible; and from these *1185 follow the Phænomena before explained.*

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CHAP. XIX.

Concerning the Tides.

T HAT we may explain the Tides from the Principles already laid down, we muft confider, that the Earth, as alfo all Bodies near it, *1206 gravitate towards the Moon ;* therefore the Particles of Water in the Earth's Surface which tend towards the Center of the Earth (for here we neglect the Confideration of N° 1371.) are carried with it towards the Moon. Since alfo the folid Mafs of the Earth is carried towards the Moon, according to the Laws, which would obtain

obtain, if all the Matter of which it confifts was collected in its Center; * what has been demon-*1126 ftrated in Chap. 16. of the Action of the Sun upon the Moon falling towards the Earth, whilft with 1390 the Earth it goes towards the Sun, may be applied to the Action of the Moon upon the Particles of Water in the Earth's Surface, which do not cohere with the Mafs of the Earth, but tend towards its Center, and continually with its Mafs fall towards the Moon; by which Force, as we have fhewn,* the Earth is keptinits Orbit, about *1280 the common Center of Gravity of the Earth and the Moon.

Plate XXV. Fig. 1.] Let S be the Moon; 1391 ALBL/ the Surface of the Earth, whole Mais tends towards the Moon, as if it was all collected at T ; by the Action of the Moon, the Particles of Water, A and B, acquire a greater Gravity towards T ;* on the contrary, the Parti-*1287 cles at L.1, lofe of their Gravity ;* whence we *1296 deduce, that if the whole Earth was covered with Water, there would not be an Aquilibrium, unlefs this Water was more elevated in the Points L and I, than in a whole Circle 90 Degrees diftant from these Points; and therefore paffing through the Points A and B. Therefore by 1302 the Astion of the Moon, the Water acquires a (pheroidical Figure, formed by the Revolution of an Oval about its greater Axis, which, being continued, goes through the Moon.

Let us suppose the Moon in the Æquator; all the Sections of the Earth which are parallel to the Æquator, as they are also parallel to the Axis of the Spheroid* are Oval, whose greater Axes *1392 pass through the Meridian of the Moon; whence it follows, that the Earth being at Rest, in any 1393 Circle of Latitude, the Water is more elevated VOL. II. 'T in

in the Meridian in which the Moon is, and in the opposite Meridian, than in the intermediate Places.

DEFINITION.

1394 The Lunar Day is the Time Spent between the Moon's going from the Meridian and coming to it again. This Day is divided into 24 Lunar Hours: It is 50 Minutes longer than the Natural Day.

From the Motion of the Earth round its Axis, every Lunar Day, every Place paffes through the Meridian of the Moon and the opposite Meridian, that is, twice paffes through that Place, 1395 where the Water is raifed by the Action of the Moon, and twice through that Place, where the *1393 Water is dispersed by the fame Action*; and fo in a Lunar Daythe Sea is twice elevated, and twice

depressed, in any assigned Place.

- 1396 By the Motion of the Earth round its Axis, the elevated Water continually recedes from the Meridian of the Moon; yet by the Action of the Moon, the Axis of the Spheroid paffes
- *1392 through the Moon;* therefore the Water is continually agitated, that the Elevation, which (on Account of the Motion of the Earth) is removed, may be brought under the Moon. Therefore the Water continually flows from A and B (Plate XXV. Fig. 1.) towards L and I, whilft, by the Motion of the Earth, the Elevation is carried from L towards B, and from I towards A; that is, between L and B, as alfo between I and A, there are two contrary Motions, by which the Water is accumulated; fo that the greatest Elevations are between these Points (viz.) not directly under the Moon, but on one Side of that Point, and likewife afide of 1397 the opposite Point. That is, in any Place, the Water most elevated, two or three Hours after the Moon

Moon has paffed the Meridian of the Place, or the opposite Meridian.

The Elevation towards the Moon a little exceeds 1398 the opposite one.* The Ascent of the Water is dimi-*1390 nished, as you go towards the Poles, because there 1295 is no Agitation of the Water there.

What has been demonstrated, in relation to 1399 the Moon, may be applied to the Sun; therefore 1400 from the Action of the Sun, every natural Day, the 1401 Sea is twice elevated, and twice depressed *. This *1395 Agitation is much less, on account of the immense Distance of the Sun, than that which depends upon the Moon; yet it is subject to the same Laws. 1402

The Motions, which depend upon the Actions of 1403 the Moon and Sun, are not diftinguished, but confounded; and from the Action of the Sun, the Lunar Tide is only changed; which Change varies every Day, by reason of the Inequality between 1404 the Natural and Lunar Day*.

In the Syzygies, the Elevation, from the Actions 1405 of both Luminaries, concur, and the Sea is more elevated; the Sea afcends lefs in the Quadratures; for where the Water is elevated by the Action of the Moon, it is difperfed by the Action of the Sun, and fo on the contrary. Therefore, whilf 1406 the Moon paffes from the Syzygy to the Quadrature, the daily Elevations are continually diminified: On the contrary, they are encreafed when the Moon moves the Quadrature to the Syzygy. At a New 1407 Moon alfo, cateris paribus, the Elevations are greater, and those that follow one another the fame * 1398 Day, are more different than at a Full-Moon.* 1402

The greatest and least Elevations are not observed, 1408 till the second or third Day after the New or Full Moon; because the Motion acquired is not presently destroyed from the Attrition, and other Causes, by which acquired Motion the Ascent of the Water is encreased. Although the Ascion by Vol. II. T 2 which

which the Sea is raifed be diminished; fomewhat like to what we have demonstrated elfewhere con-*1174 cerning Heat*.

1409 If now we confider the Luminaries receding from the Plane of the Æquator, we shall perceive that the Agitation is diminished, and becomes less, according as the Declination of the Luminaries becomes greater. Which plainly appears, if we conceive them to be in the Poles ; for then the Axis of the fpheroidical Figure coincides with the Action of the Earth ; and all the Sections that are parallel to the Æquator, are perpendicular to the Axis of the Spheroid; and therefore circular. So that the Water, in every Circle of Latitude, will have every where the fame Elevation; and fo in the Motion of the Earth, the Height of the Sea is not changed in particular Places. If the Luminaries recede from the Poles, it is eafy to find, that the Agitation will be more and more encreafed, till it be the greateft of all, the Spheroid revolving about a Line perpendicular to the Axis, the Axis of the Spheroid being fuppofed in the Plane of the Æquator.

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1410 Hence it is plain, why in the Syzygies, near the Æquinoxes, the Tides are observed to be the greateft, both Luminaries being in or near the Æquator. The Actions of the Moon and Sun are greater, the IAII * 1354 less those Bodies are distant from the Earth*; but

*1390 when the Diftance of the Sun is lefs, and it is in the South Signs, often both the greateft Æquinoctial Tides are observed in that Situation of the Sun; that is, before the Vernal, and after the Autumnal Æquinox, which yet does not happen every Year; because fome Variation may arife from the Situation of the Moon's Orbit, and the Diftance of the Syzygy from the Æquinox.

Plate XXV. Fig. 8.] In Places diftant from the 1412 Æquator, as the Luminaries recede from the Æquator, the Elevations, that happen the fame Day, are unequal. Let PP be the Axis of the Earth, EE the Æquator, Ll a Circle of Latitude; A B the Axis of the fpheroidical Figure which the Water forms: When a Place in the Circle L1 is given at L or l, it is given in the fame Meridian with the Axis of the Spheroid, and the Water is most elevated, in both Cafes; yet at L it is more elevated than at 1; for CI exceeds C1, which Lines meafure the Heights of the Waters, that is, the Diftances from the Center : Thefe Lines would be equal if A L and B/ (which are the Diftances from the Axis of the Spheroid) were equal. But Cl is lefs, becaufe Bl exceeds AL. which arifes from the Inclination of the Axis of the Spheroid to the Æquator.

As long as the Moon is on the fame Side of the 1413 Æquator in any Place, that is, towards the Line CA continued, the Elevation of the Water is ob- 1414 ferved to be greatest every Day, after the Moon has passed the Meridian of the Place; for there is the greateft Elevation when the Place is come to L; but if the Æquator separates, or is between the Moon and the Place, of which we fpeak, that is, if the Moon be towards the Line CB continued, the Water again at L will come to the greatest Height, and, every Day, the greatest Elevation of the Sea will be, after the Moon has paffed through the opposite Meridian.

All Things, which have been hitherto explained, would exactly obtain, if the whole Surface of the Earth was covered with Sea; but fince the Sea is not every where, fome Changes arife from thence, not indeed in the open Sea; becaufe the Ocean is extended enough to be fubject to the Motions we have spoken of. But the

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1415 the Situation of the Shores, the Streights, and many other Things depending upon the particular Situation of the Places, difturb these general Rules. Yet it is plain from the most general Observations, that the Tide follows the Laws which we have explained. What remains is, to determine the Forces with which the Sun and Moon act upon the Sea, that it may appear, that they are able to produce the Effects which we have mentioned; and that the Actions of those Bodies, upon Pendulums and other Bodies, are infensible.

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1416 The Encreafe or Addition to the Gravity of the Moon in the Quadratures, from the Action of the Sun, is to the Gravity of the Moon to*1291 wards the Earth, as 1 to 178, 73*. In which Computation we have fuppofed the mean Diftance of the Moon from the Center of the
1285 Earth to be 60 Semidiameters of the Earth: Therefore the Gravity of the Moon is to the Gravity of the Earth's Surface, as 1 to 60×60
1208 =3600. Therefore the above-mentioned Encreafe is to the Gravity on the Earth's Surface, as 1 to 643428, in which Computation there is an Error to be corrected.

This Computation would be exact, if the Encreafe, of which we fpeak, was to the Force, with which the Earth defcends towards the Sun, as the Diftance of the Moon (which is 60 Semidiameters of the Earth) to the Diftance of *1287 the Earth from the Sun*; but it is as the true mean Diftance of the Moon, which is 60¹/₂ Semidiameters of the Earth to the Diftance of the Earth from the Sun. Wherefore the Encreafe, that we have just determined, ought to be $\frac{1}{120}$ Part encreafed, and will be to the Force of Gravity on the Earth's Surface, as $1 \frac{1}{120}$ to 643428, or as 1 to 638110, 4.

This Encreafe of the Gravity of the Moon, in the Quadratures from the Action of the Sun,

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is to the Encrease of the Gravity of the Water on the Earth's Surface, in Places which are 90 Degr. diftant from the Sun (from the fame Action of the Sun) as 601 to 1*. Therefore this * 1288 Encrease of the Gravity is to the Gravity of the Water, as I to 38605679. The Diminution of the Gravity, under the Sun, and in the oppofite Place, is double this Encrease *; therefore it is to *1296 the Gravity as I to 1930 839, and the whole 1417 Change in the Gravity, arising from the Action of the Sun, is to the Gravity it [elf, as 1 to 12868560.

In order to compare the Action of the Moon 1418 with the Action of the Sun, we must make Experiments in Places, in which, by reafon of the Narrownefs, the Sea is fenfibly raifed. Near Briftol at the Autumn and Spring, at which Times the Agitation of the Sea is greateft*, the *1410 Water afcends in the Syzygies about 45 Feet, more or lefs, in the Quadratures about 25 Feet, more or lefs. Which Numbers are to one another as 9 to 4.

The Determination of the Forces, which we would find, if the greateft and leaft Elevations were exactly at the Time of the Syzygies, would be very eafy, which we have fhewn before not to happen fo*. *1408

The Diftance of the Moon from the Syzygy, or the Quadrature, is not always the fame in the greateft or leaft Elevation; for this Diftance varies, because the Moon is fometimes more, and fometimes lefs diftant from the Meridian, when it goes through the Syzygy or Quadrature. The mean Diftance of the Moon from the Syzygy, or Quadrature, to which the aforefaid Obfervations ought to be referred, is about 18 Degr. 30 Min. fo that the whole Action of the Sun neither confpires with the Action of the Moon in the Syzygies, nor acts contrary to it in the Quadratures. Alfo in fuch a Cafe, if at the

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the Syzygy, both the Luminaries be in the Æquator, in the faid Diftance from the Quadrature, the Declination of the Moon is 22 Degr. 13 Min. more or lefs; whereby the Force of the Moon to *1409 move the Seais diminished*. Besides, cæteris paribus, the Diftance of the Moon from the Earth *1310 at the Syzygies is lefs than at the Quadratures ;* 1311 whence also the Action of the Moon is diminish-*1411 ed at the Quadratures :* By attending to all which Things we may difcover, that the mean Force of 1419 the Sun, to move the Sea, is to the mean Force of the Moon to move the fame, as 1 to 4,4815. But the Force of the Sun is to the Force of Gravity, as *1417 I to 12868560; * wherefore the Force of the Moon is to the fame Force of Gravity as 1 to 2871485. 1420 From whence it follows, that these Forces of the Moon and Sun are too fmall to be fenfible in Pendulums and other Experiments; but it is eafily proved, that they are capable of agitating the Sea.

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By diminishing the Gravity 2. Part, the Sea *1368 is raifed to the Height of 88902 Rhynland Feet*. 1417 For each Perch contains 12 Feet; whence, by the 1421 Rule of Proportion*, we find that the Action of 1422 the Sun changes the Height of the Sea two Feet, *1419 and that the Action of the Moon changes it 8, 95;* and that from the joined Action of both, the mean Agitation is of about eleven Feet, which agrees pret-1423 ty well with Obfervations; for in the open Ocean, as the Sea is more or lefs open, the Water is raifed to the Height of Six, Nine, Twelve, or Fifteen Feet; in which Elevations also there is a Difference arifing from the Depth of the Waters. But those Elevations, which far exceed these, happen, where the Sea violently enters into 1424 Streights or Gulphs; where the Force is not broken, till the Water rifes higher.

CHAP.

CHAP. XX.

Of the Moon's Denfity and Figure.

THE Forces of the Sun and Moon, for 1425 giving Motion to the Sea, are to one another, in a Ratio compounded of the Ratio of the Quantities of Matter in thefe Bodies* (for all the *1207 Particles of Matter act) and the inverse Ratio of the Cubes of the Diftances of the Sun and Moon from the Earth * *1390

The Quantities of Matter are in a Ratio com-1402 1354 pounded of the Ratio of the Bulks, that is, of the Cubes of the Diameters, and the Ratio of the Denfities;* wherefore the Forces abovementioned are directly as the Denfities and the Cubes of the Diameters, and inverfly as the Cubes of the Diftances.

The apparent Diameters of Bodies, that is, the Angles under which they are feen, encreafe as the Diameters themfelves, and diminish as the Diftances ; that is, they are directly as the Diameters, and inverfly as the Diftances ; therefore the Ratio compounded of the Ratio's of the Cubes of the apparent Diameters of the Sun and Moon, and of the Ratio of the Denfities, will be the Ratio of the Forces, whereby those Bodies act upon the Sea. Therefore the Densities of those 1426 Bodies are directly as the Forces, whereby they move the Sea, and inverfly as the Cubes of their apparent Diameters ; and dividing the Forces by the Cubes of those Diameters, you have the Ratio of the Denfities.

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The Force of the Sun is to the Force of the *1419 Moon, as I to 4, 4815;* the mean apparent Diameter of the Sun is 32 Min. 12 Sec. and the mean apparent Diameter of the Moon is 31 Min. 16 Sec. That is, they are to one another, as 1427 3864 to 3753. Therefore the Density of the Sun is to the Moon's Density, as 10000 to 4.8911 : which Denfity of the Moon may be compared with the *1260 Denfities of Jupiter, Saturn, and the Earth*, and the Moon is denfer than the Earth.

The Quantities of Matter in two Bodies are to one another, in a Ratio compounded of the Denfities and Bulks;* that is, if the Body be a *79 Sphere, in a Ratio compounded of the Denfities and the Cubes of the Diameters,

The Densites of the Moon and Earth are to one 1428 *1427 another, as 48911 to 39214*; the Diameters 1260 as 11 to 40, 2. therefore the Quantities of Matter, in those Bodies, are as 1 to 39, 13. Though the Denfities be difcovered, if you suppose the Bodies to be homogeneous; yet the Quantities of Matter will be rightly defined, though the Bodies are not homogeneous; for we determine the Denfity which that Body would have, if the Matter, of which the Body really confifts, was equally diffused all over it.

The Gravities, on the Surfaces of the Earth and 1429 Moon, are determined, by multiplying the Denfi-*1230 ties by the Diameters*, that is, they are to one

another, as 2, 93 to 1, or as 407,8 to 139, 2. Which Number alfo does express the Relation of Gravity on the Surface of the Moon, with the Gravity on the Surfaces of the Sun, Jupiter and *1258 Saturn.*

The common Center of Gravity of the Moon and 1430 Earth, about which both Bodies are moved, is determined; for its Diftance from the Center of the

the Earth, is to the Diftance between the Centers of both Bodies, as the Quantity of Matter in the Moon to the Quantity of Matter in both Bodies;* therefore 40, 13. is to 1: as the Diftance *234 of the Moon from the Earth, is to the required ²³⁵ Diftance of the Center of Gravity, from the Center of the Earth, which is found to be of \$126950 Perches, as is deduced from the known Diameter of the Earth*, and the Diftance of the *976 Moon.

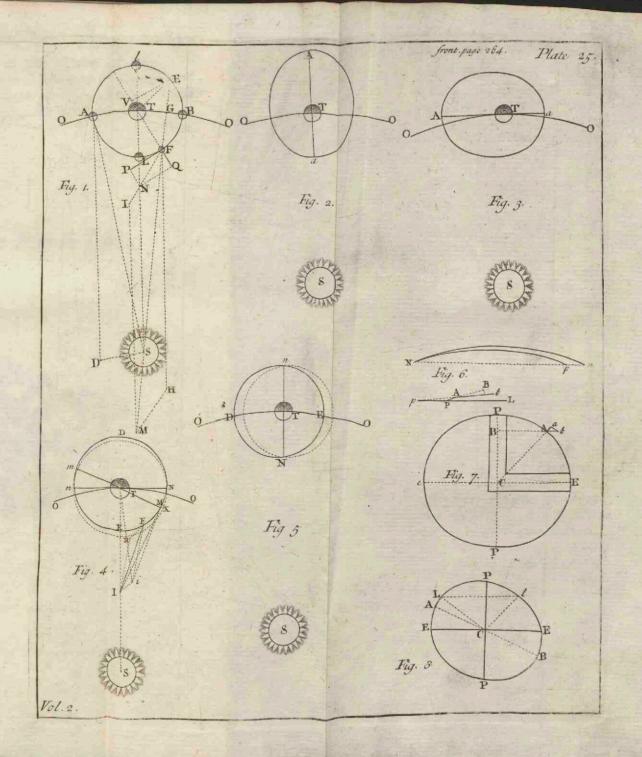
To determine the Figure of the Moon, we 1431 must examine what Figure it would have, if it was fluid*. If we confider the Moon alone at *1358 reft, it would be fpherical :* If we confider the *1359 Action of the Earth upon the Moon, the Moon would acquire the Figure of a Spheroid, whole Axis would go through the Earth*. The Force *1392 of the Earth, for changing the Figure of the Moon, is to the Force of the Moon upon the *1428 Earth, as 39, 13. to 1.* and as the Diameter of 1428 the Moon to the Earth's Diameter,* which are *1390 to one another as II to 40, 2, and it is a Ratio 1354 compounded of thefe 10 7. to 1. This Force of the Moon is to the Gravity upon the Earth's Surface, as I to 2871485 ;* which Gravity, on *1420 the Earth's Surface, is to the Gravity on the Surface of the Moon, as 407, 8, to 139, 2.* or *1429 as 2871485, to 980028 : Wherefore the Action of the Earth, for changing the Moon's Figure, is to the Gravity upon the Moon's Surface, as 10, 7. to 1432 980028, or as I to 91524, the Gravity being changed on the Earth's Surface, by 2871483 Part the Water is raifed 8, 95 Feet;* and therefore if *1420 Gravity was to be changed off Part, the Eleva-1422 tioned would be of 280, 7 Feet, as it is found by the Rule of Three. If, keeping this Diminution of Gravity, we confider a lefs Body, this Height muft

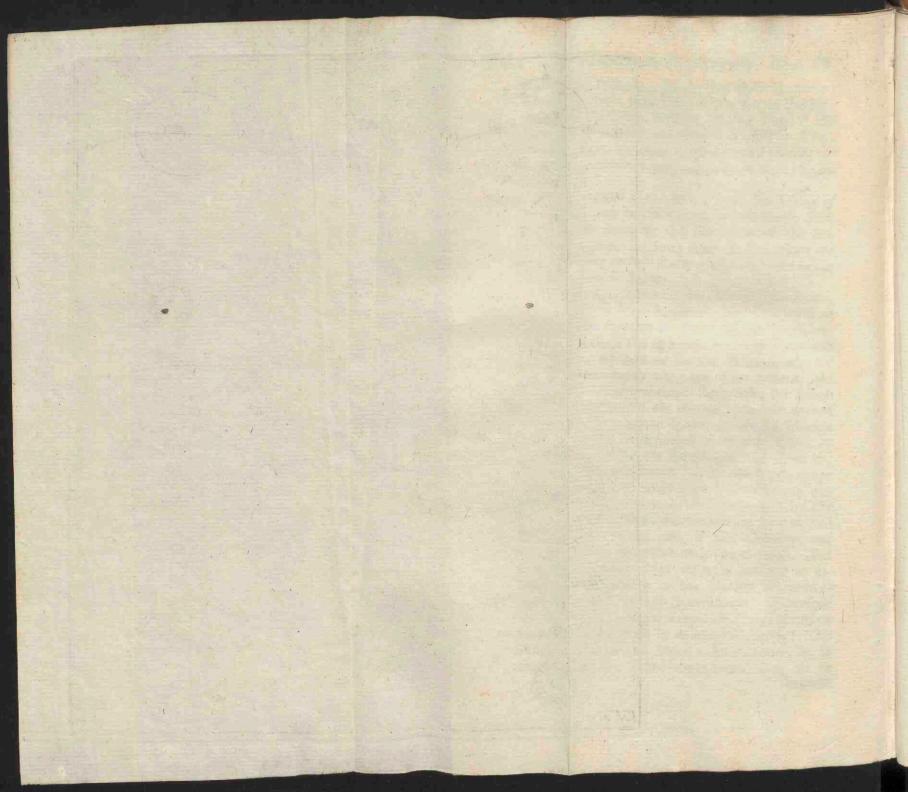
must be diminished, in Proportion to the Diameter : Therefore, from the Action of the Earth, the Elevation of the Moon is of 76, 8 Feet ; and if the

1433 Moon be homogeneous, there will not be an Æquilibrium, unless the Axis of the Spheroid exceeds the Diameter, which is perpendicular to it, 153, 60 Feet.

1434 The Elevation of the Moon, from the Action of the Earth, may be differed by one fingle Proportion, by knowing the Elevation of the Sea from the Moon's Action; forthefe Elevations are in a duplicate inverse Ratio of the Gravities on the Surfaces of those Bodies.

1435 If, Supposing this to be the Figure of the Moon. we conceive the Parts to cohere, there will not be an Æquilibrium between the Parts of the Moon, unless the Axis of the Spheroid be directed towards the Earth ; whence we fee the Reafon, why the Moon always turns the same Face towards the Earth; by which continual Agitation, the Moon 1436 has at least acquired the Motion about its Axis of *970 which we have before fpoken;* which Motion 1092 must necessarily be performed in the same Time, as the Moon performs one Revolution; for from the Action above-mentioned, it must neceffarily adapt itfelf to fuch a Celerity; for if the Celerity was greater, it would be continually retarded by the Force, whereby the fame Face is always directed towards the Earth; and, if this Celerity was lefs, it would be continually accelerated. Yet this Force is not great enough, fenfibly to difturb the Æquibility of the Motion acquired a-1437 bout the Axis, every Revolution : Therefore the Motion about the Axis is equable, though the Moon *966 be moved in its Orbit by an unequal Motion*. The Polition alfo of the Moon's Axis cannot be fo changed by the Force above-mentioned, as to become





become perpendicular to the Plane of the Orbit, when its Inclination is changed;* there-*1345 fore the Axis of the Moon is fometimes inclined 1438 to the Plane of the Orbit, as we have before fhewn*.

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The End of the Fourth Book.





A N

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